Exploring New Physics with Electronic Recoil Data from PandaX-4T Detector

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Explosion of New Physics



Dual Phase Xenon TPC



ZEPLIN · XENON · LUX · LZ · PandaX...



- Large A: large cross section & self-shielding; 1.
- 3D reconstruction and fiducialization 2.
- Scalable; 3.

(NR)

4. **Discrimination power**

- WIMPs, v, n \rightarrow Nuclear Recoil (NR)
- Axion, γ, β

Dark matter: nuclear recoil

γ background: electron recoil (ER)

 \rightarrow Electronic Recoil(ER)







(S2/S1)_{NR}<<(S2/S1)_{FR}

10/20/23

NR Searches V.S. ER Searches



Production Mechanisms of Solar Axion

Atomic recombination and deexcitation (Axio RD in figure), Bremsstrahlung, and Compton (ABC):

$$\Phi_{\rm a}^{\rm ABC} \propto g_{\rm ae}^2$$

□Primakoff effect:

$$\begin{split} \frac{d\Phi_{\rm a}^{\rm Prim}}{dE_{\rm a}} = & \left(\frac{g_{\rm a\gamma}}{{\rm GeV}^{-1}}\right)^2 \left(\frac{E_{\rm a}}{{\rm keV}}\right)^{2.481} e^{-E_{\rm a}/(1.205~{\rm keV})} \\ & \times 6 \times 10^{30}~{\rm cm}^{-2}{\rm s}^{-1}{\rm keV}^{-1}, \end{split}$$

■M1 nuclear transition of ⁵⁷Fe (14.4 keV):

$$\Phi_{\rm a}^{^{57}\rm Fe} = \left(\frac{k_{\rm a}}{k_{\gamma}}\right)^3 \times 4.56 \times 10^{23} (g_{\rm an}^{\rm eff})^2 \ \rm cm^{-2} s^{-1}$$



ER Signals from Axio-electric Effect



The approximate cross section of axio-electric effect is:

$$\sigma_{Ae}(E) = \sigma_{\rm pe}(E) \frac{g_{Ae}^2}{\beta} \frac{3E^2}{16\pi\alpha m_e^2} \left(1 - \frac{\beta^2}{3}\right)$$

where $g_{Ae} = C_{ae}m_a/f_a$, C_{ae} is model dependent.

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Other signals

Weakly Interacting Slim Particles (WISPs)

• Axion-like Particles (ALPs)

$$R \simeq \frac{1.5 \times 10^{19}}{A} g_{\rm ae}^2 \left(\frac{m_{\rm a}}{\rm keV/c^2}\right) \left(\frac{\sigma_{\rm pe}}{\rm b}\right) \rm kg^{-1} d^{-1}$$
(1)

• Dark Photons

$$R \simeq \frac{4.7 \times 10^{23}}{A} \kappa^2 \left(\frac{\text{keV}/c^2}{m_{\text{V}}}\right) \left(\frac{\sigma_{\text{pe}}}{\text{b}}\right) \text{kg}^{-1} \text{d}^{-1}$$
(2)

Neutrino Magnetic Moment Enhanced Neutrino-electron Scattering

$$\frac{d\sigma_{\mu}}{dE_{\rm r}} = \mu_{\nu}^2 \alpha \left(\frac{1}{E_{\rm r}} - \frac{1}{E_{\nu}}\right) \tag{3}$$

Particle and Astrophysical Xenon Experiments



PandaX-4T Detector System Layout



PandaX-4T Commissioning Run (Run 0)

□Sensitive volume: 3.7 tonne xenon

Commissioning started from Nov/2020 (95 days)

- 0.63 tonne-year exposure, 1058 candidates
- Sensitivity improved from PandaX-II final analysis by 2.9 times (30 GeV/c2);







New DM searching channel: $\chi \longrightarrow \nu$



- Elastic scattering v.s. Inelastic scattering (Absorption).
 - DM being absorbed, with an outgoing neutrino v

Fermionic DM absorption model through neutral current:

- Detectable NR signals with higher recoil energy .
- DM mass range: MeV

New DM searching channel: $\chi \longrightarrow \nu$



Phys. Rev. Lett. 129, 161803

□ First search for fermionic dark matter absorption signal in direct detection experiments

□Strongest limit of 1.5 x 10⁻⁵⁰ cm² achieved at 40 MeV/c² fermionic DM mass

 \Box Constraints on the coupling g_{χ} to the order of $10^{-10} (\text{TeV}\cdot\text{cm})^{0.5}$

New DM searching channel: $\chi e \rightarrow e \nu$



- General fermionic (sterile neutrino-like) dark matter absorption on e⁻;
- Strong sensitivity to vector and axialvector mediators; Complementary to astrophysical constraints, with much smaller theoretical uncertainties;

□ Competitive constraint in 20-55 keV/c²

Phys. Rev. Lett. **129**, 161804

Electro-magnetic Properties of DM

Minute photon-mediated interactions through millicharge or higher-order multipole is still possible;

Direct search for effective EM interactions: first experimental constraint on DM charge radius;

• 4 orders of magnitude smaller than neutrino

Other EM properties.

Table 1	Comparison of electromagnetic properties
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	Dark matter	Neutrino	Neutron
Charge radius (fm²)	<1.9×10 ⁻¹⁰	(−2.1,3.3)×10 ^{-6 a}	–0.1155 °
Millicharge (e)	<2.6×10 ⁻¹¹	<4×10 ^{-35 a}	(-2±8)×10 ^{-22 a}
Magnetic dipole ($\mu_{\rm B}$)	<4.8×10 ⁻¹⁰	<2.8×10 ^{-11 a}	-1×10 ^{-3 a}
Electric dipole (ecm)	<1.2×10 ⁻²³	<2×10 ^{-21 b}	<1.8×10 ^{-26 a}
Anapole (cm²)	<1.6×10 ⁻³³	roughly 10 ^{-34 c}	roughly 10 ^{-28 d}
			7.5

^aData are taken from the Particle Data Group³³.

X. Ning et al. Nature 618, 47-50 (2023)



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After Run 0

□Tritium removal

• xenon distillation, gas flushing, etc

2021/11 – 2022/05: physics run (Run1)

• 164 days: ~ 1 tonne-year

2022/09 - 2023/10: hall construction

- xenon recuperation
- detector upgraded

Expect to resume by the end of 2023

Commissioning (Run 0)	Calibration	Distillation	Physics Run (Run 1)	Calibration	Detector Upgrade
2020/11/28 _ 2021/04/16	2021/04/17 _ 2021/06/09		2021/11/15 _ 2022/05/15	2022/05/16 _ 2022/07/08	





Data Taking Condition of Run 1

□Gate -6kV, Cathode -16kV;

- □Xenon purity monitor: Maximum electron lifetime reaches 1800 us;
- Liquid level is monitored through the drift time of gate events and single electron gain (SEG);

□Additional 10 top PMTs turn-off.





Self-Calibration of PMT Gain and Signal Yield

Degrading of PMTs:

- LED calibration: once a week, not instant monitoring;
- correction factor derived from single hit distribution.

□Instability of signal yield: S1 & S2 of monoenergetic peak evolve by time:

 correction factor derived from S1 and S2 with 5.5MeV alpha events from ²²²Rn decay;



• likely related to the liquid level.



164keV centerE vs runNumber

Detector Calibration

Rn, D-D neutron and AmBe neutron for low energy region



ER Calibration



□ The detector response is modelled by **NEST (Noble Element Simulation** Technique). Detector parameters are fit to Rn calibration data using unbinned likelihood fitting with emcee; □ Calibrated NEST fits well with ^{83m}Kr data, both in spectra and resolutions: Energy resolution @ 41.5 keV: 6.8% Phys. Rev. Lett. 127, 261802

Phys. Rev. Lett. 129, 161803

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Solar Axion Signals in PandaX-4T



Expected solar axion signals in energy space produced by different mechanisms:

 $\Box \text{ ABC process: } g_{Ae} = 5 \times 10^{-12};$

 \Box Primakoff effect: $g_{A\gamma} = 2 \times 10^{-10}$

$$\Box$$
 ⁵⁷Fe nuclear transition: $g_{An}^{eff} = 1 \times 10^{-6}$

Main ER Backgrounds: ²²²Rn

- Rn level varies with running conditions
- Update the analysis

Rn level	µBq/kg
Run 0	7.07 ± 0.02 (stat.) ± 0.23 (sys.)
Run 1	8.67 ± 0.01(stat.) ± 0.27(sys.)

Circulation system to be upgraded









Main ER Backgrounds: ⁸⁵Kr

Compare to Commissioning run

- tightening beta-gamma coincidence selection
- less contributions from accidental events





	β–γ events	accidental events	Kr/Xe [ppt (10 ⁻¹²)]
Run0 0.6 tonne-year	4	0.14 ± 0.04	0.5 ± 0.3
Run1 1.0 tonne-year	12	0.25 ± 0.05	0.9 ± 0.3

Main ER Backgrounds: Tritium



□ Tritium spectrum identified in the data

- □ Likely originated from a tritium calibration at the end of PandaX-II;
- □ Preliminary estimation of tritium level in Run 1: fitting S1 spectrum,

keeping S2 blinded



Background Summary



Different from WIMP analysis, we extended the ROI to 30keV

- Rn emanation;
- Krypton contamination;
- Tritium;
- Radioactive isotopes of xenon: ¹²⁷Xe, ¹³⁶Xe, ¹²⁴Xe;
- Material: radioactivity of materials are assayed by HPGe;
- pp neutrino: theoretical estimation;
- others: surface, neutron, accidental, ⁸B.

The combined analysis of run0 and run1 is under final check and the data is ready to unblind.

Summary and Outlook

- PandX-4T has finished two physics runs;
- Combined analysis of Run0 and Run1 are updated;
- **ER** responses are calibrated with ²²⁰Rn;
- Expected background contributions are estimated respectively; tritium level has significantly reduced in Run1;
- > Exploring new physics with Run0 and Run1 data of PandaX-4T is ongoing!

PandaX Collaboration







Combined Analysis of Run0 + Run1

□New active time determination

• window-size of removal time depending on the charge of large signal in front

New event window based on S2

• fix window: 1ms before and after

New event builder

 S1-S2 pairing requires quality of S1 in prior

□New selection criteria

charge-dependent cut threshold



10/20/23

Main ER Backgrounds: Radon Emanation







¹²⁷Xe (Cosmogenically Activated)



Electronics



- V1725 Digitizer, 250 MS/s;
- Self-trigger mode: read out pulses above 20 ADC (~ 1/3 PE);



- Higher sampling rate;
- · Accept out-trigger mode;