

Workshop on Multi-front Exotic phenomena in Particle and Astrophysics (MEPA 2023)



Cluster of Excellence
PRISMA+
Precision Physics,
Fundamental Interactions
and Structure of Matter

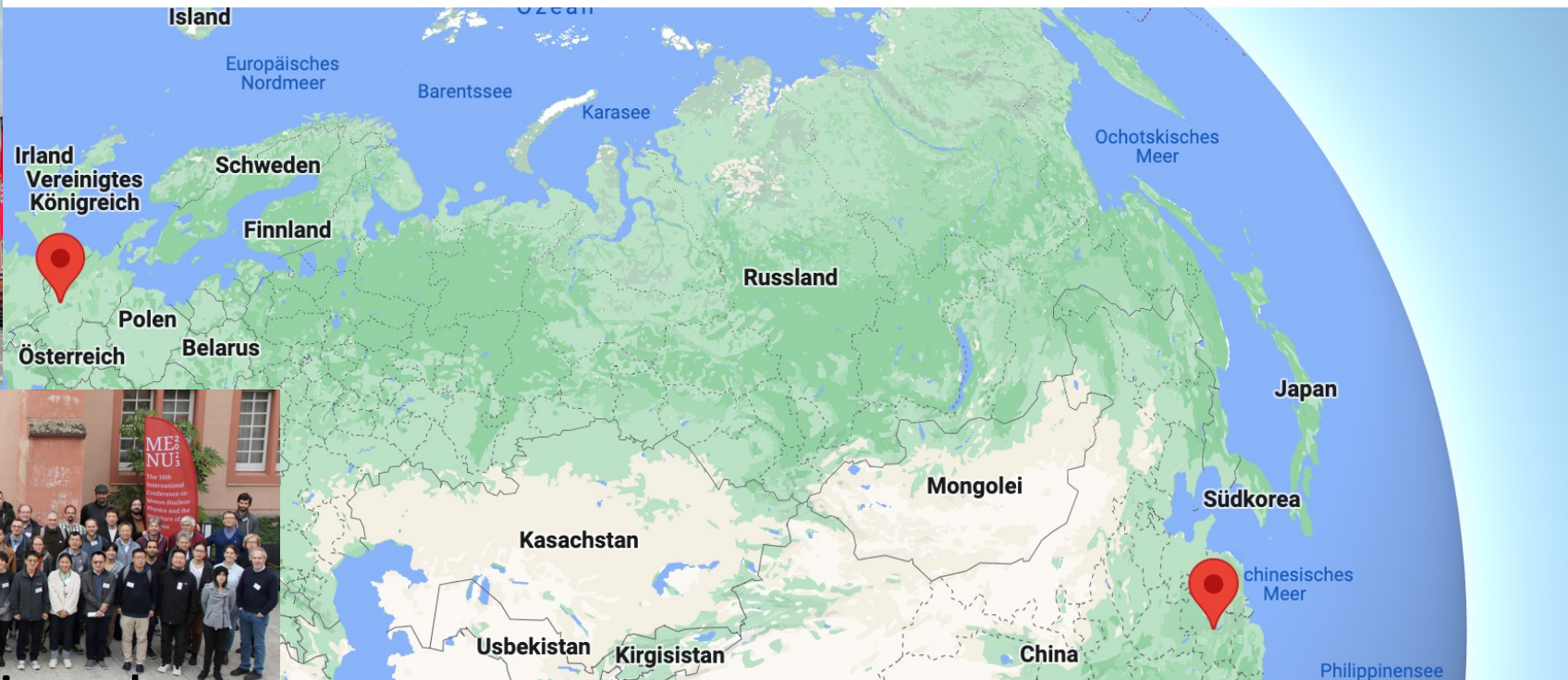
Achim Denig
October 20, 2023

Status of the DarkMESA Experiment

*Workshop on multi-front exotic phenomena
in particle and astroparticle physics (MEPA2023)
Hefei, China*



Mainz



MENU conference this week



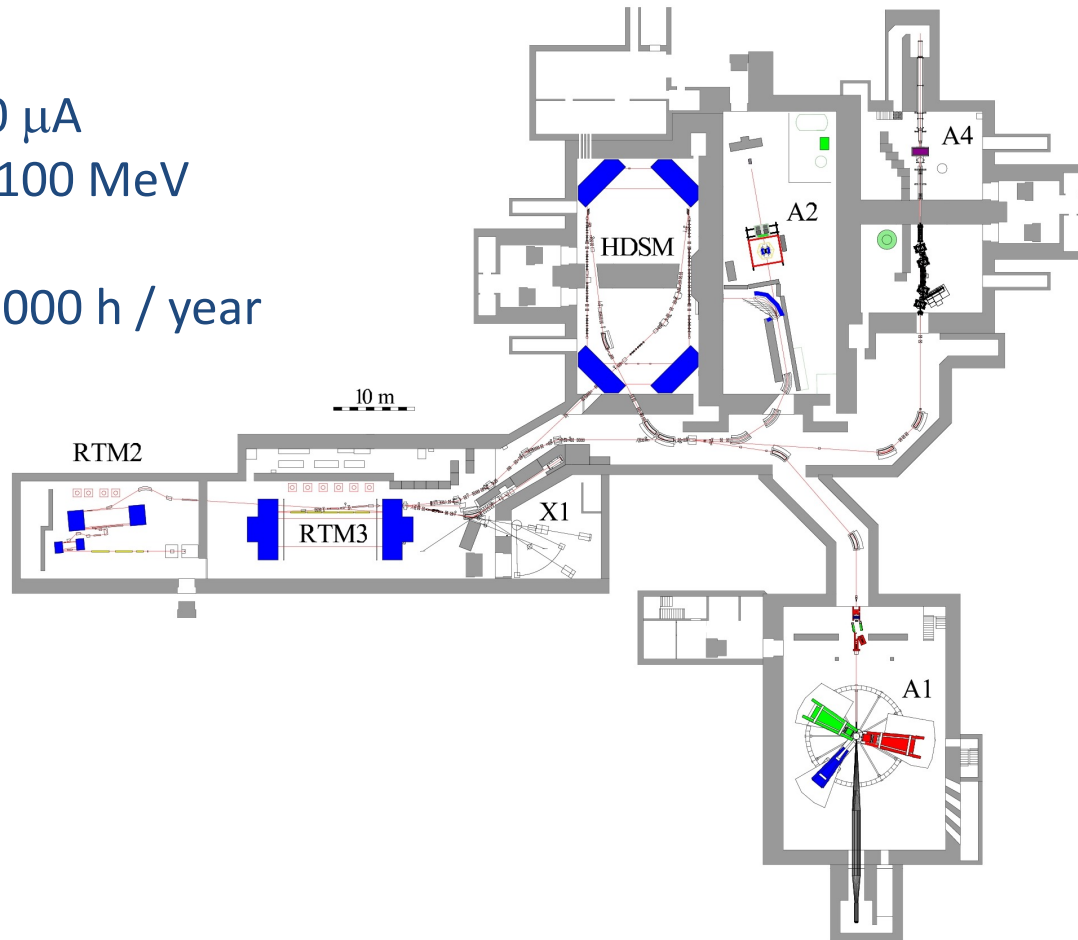
Hefei

The Mainz Microtron MAMI

**Electron Accelerator $E_{\text{max}} = 1.6 \text{ GeV}$ (CW)
operated at JGU Mainz**

Hallmarks

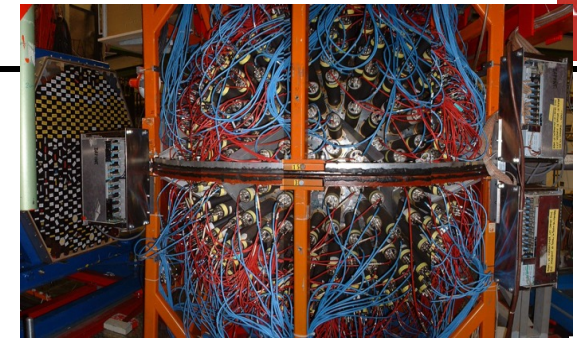
- Intensity max. $100 \mu\text{A}$
- Resolution $\sigma_E < 0.100 \text{ MeV}$
- Polarization 85%
- Reliability: up to 7000 h / year



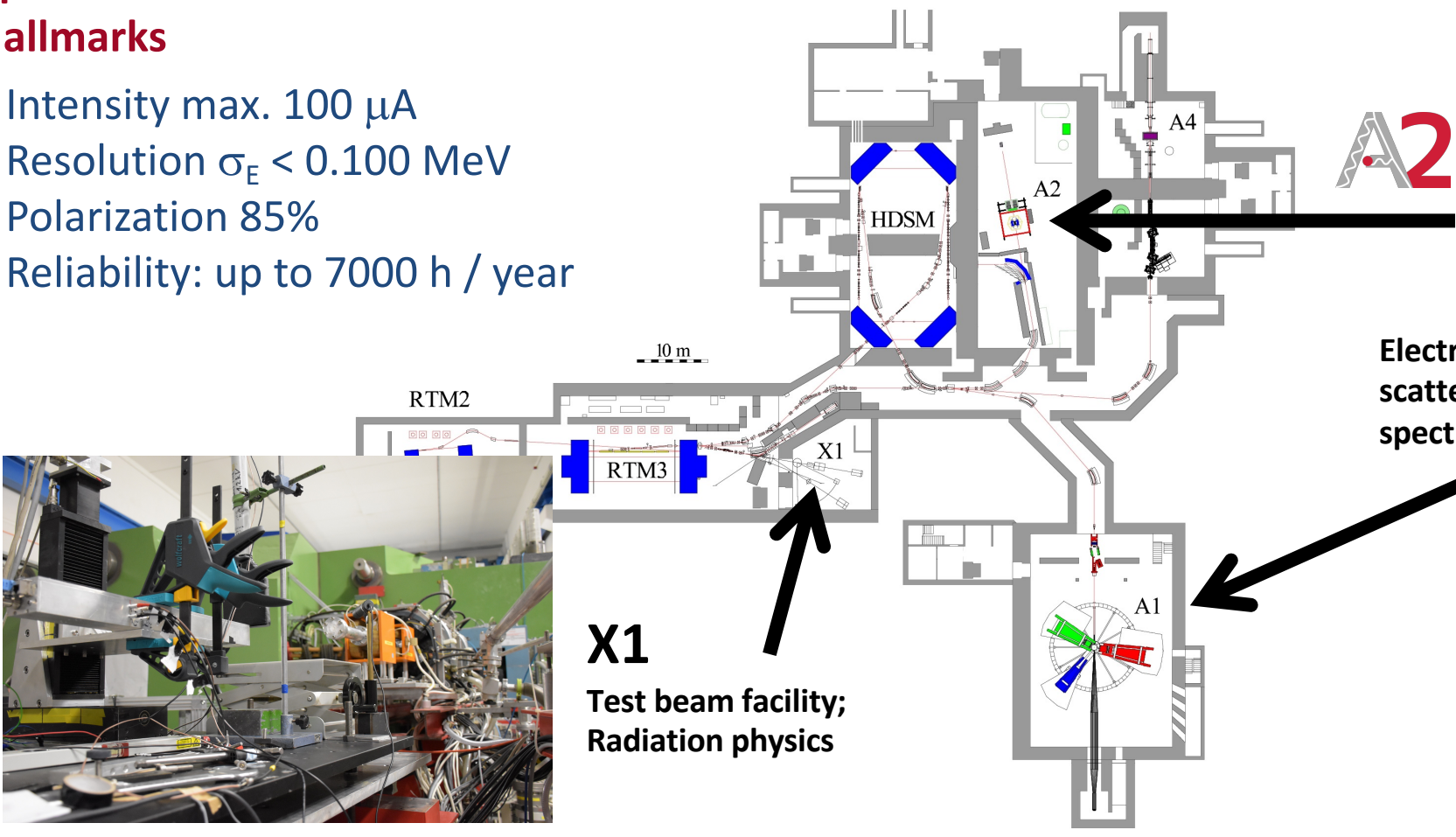
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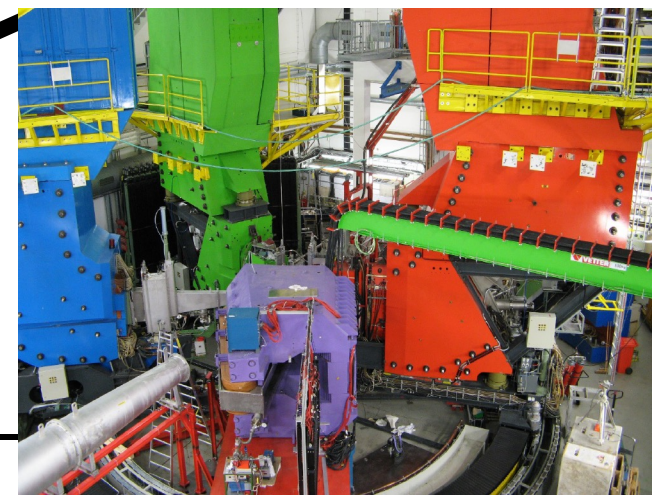
Photon scattering (A2 hall)
 (Crystal Ball / TAPS calorimeters;
 Polarized frozen-spin target
 → currently at Univ. Bonn)



Electron scattering (high resolution spectrometer setup)



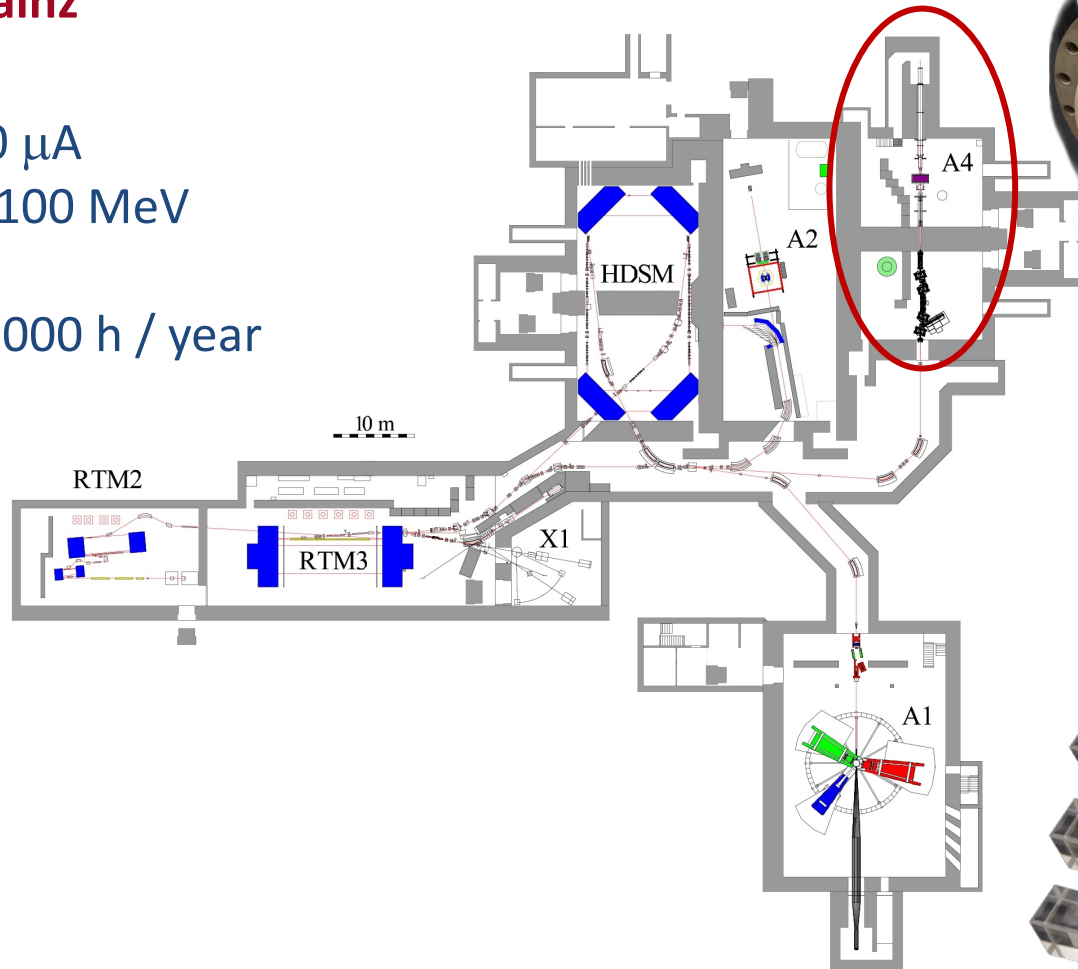
X1
 Test beam facility;
 Radiation physics



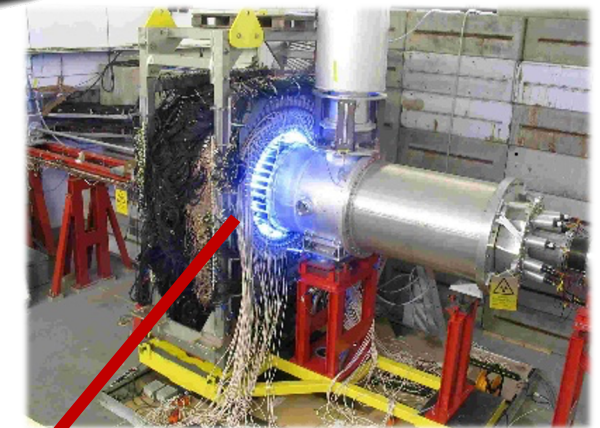
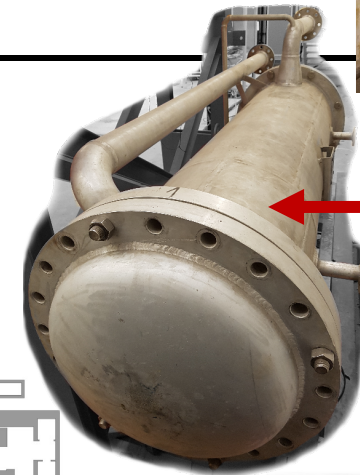
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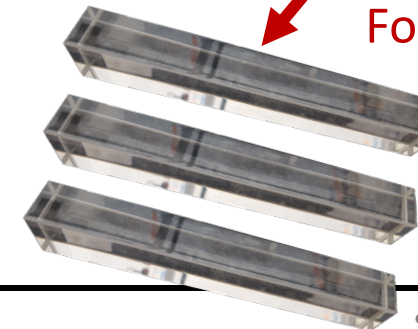
- Intensity max. 100 μ A
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Former A4 beam dump:
Al, H₂O, Cu \rightarrow 20 X₀



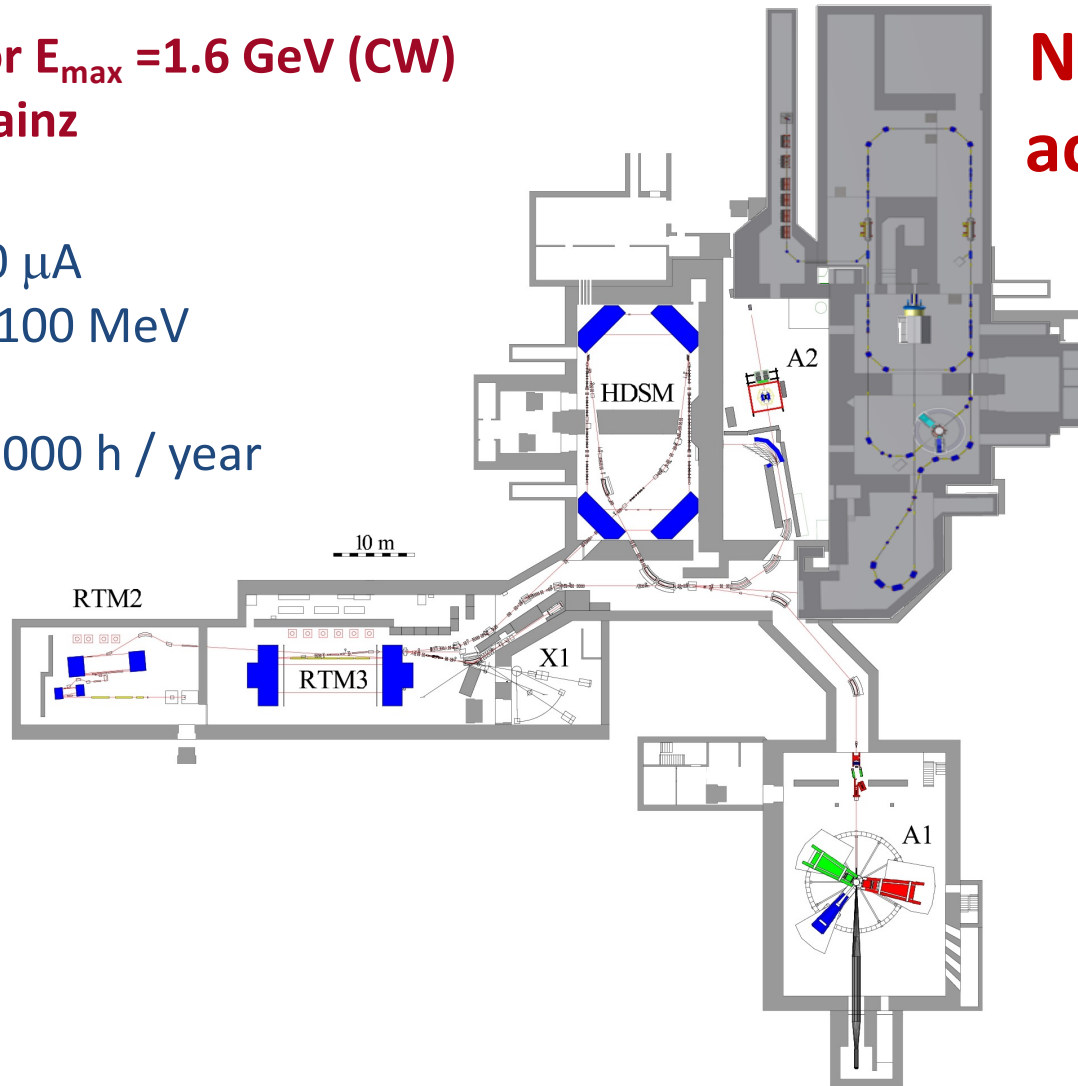
Former A4 experiment:
1000 PbF₂ crystals
and PMTs



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**New MESA
accelerator**



New experimental hall



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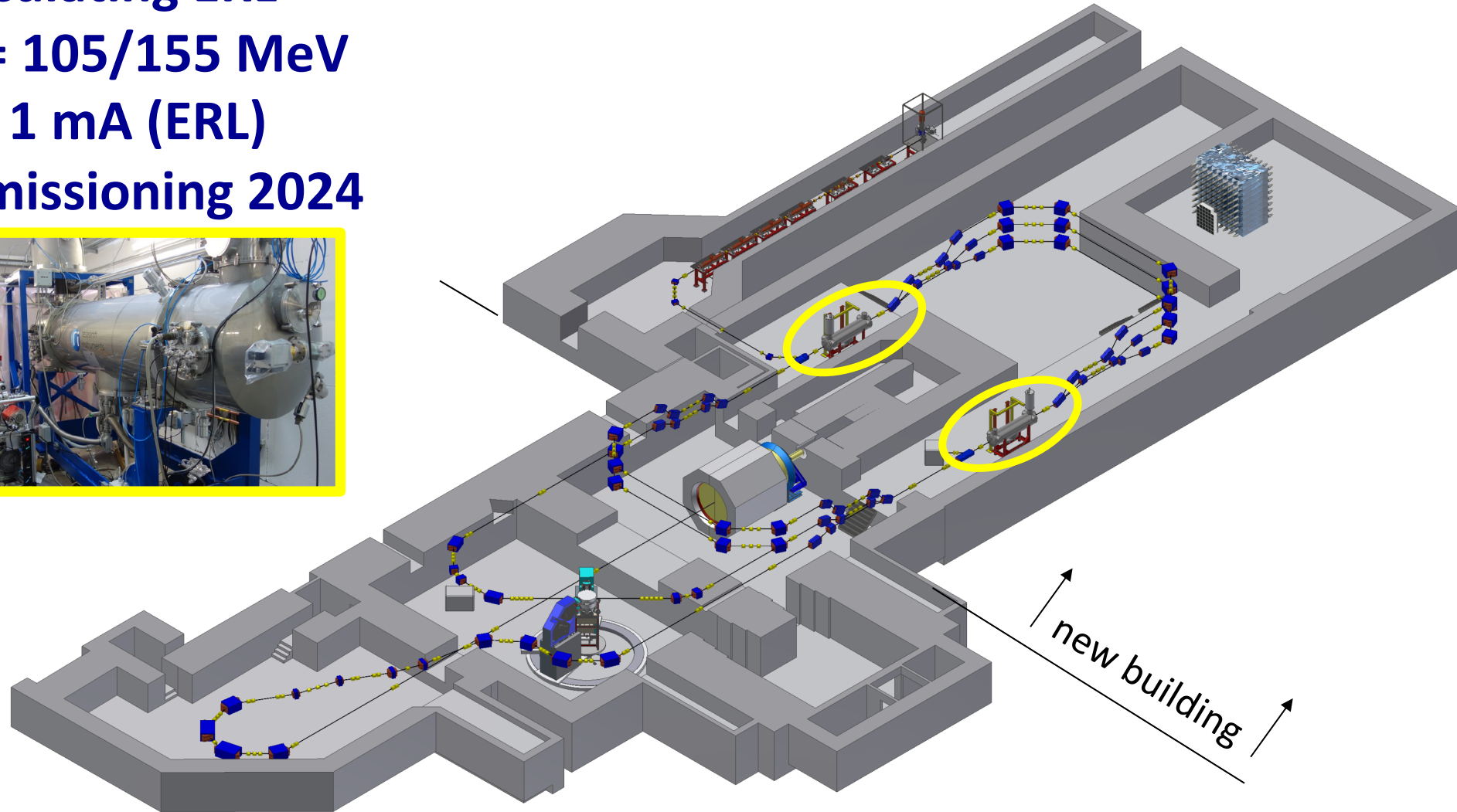
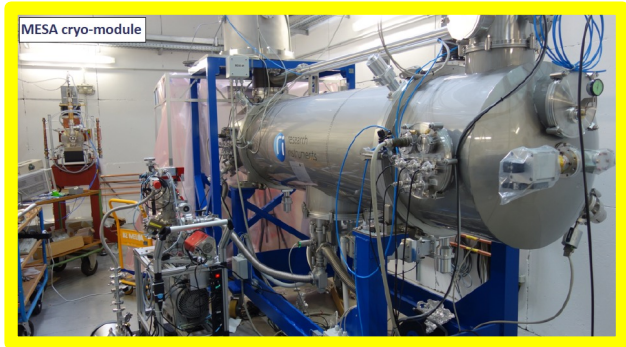
Mainz Energy-Recovering Superconducting Accelerator (MESA)

Recirculating ERL

$E_{\max} = 105/155 \text{ MeV}$

$I_{\max} > 1 \text{ mA (ERL)}$

Commissioning 2024



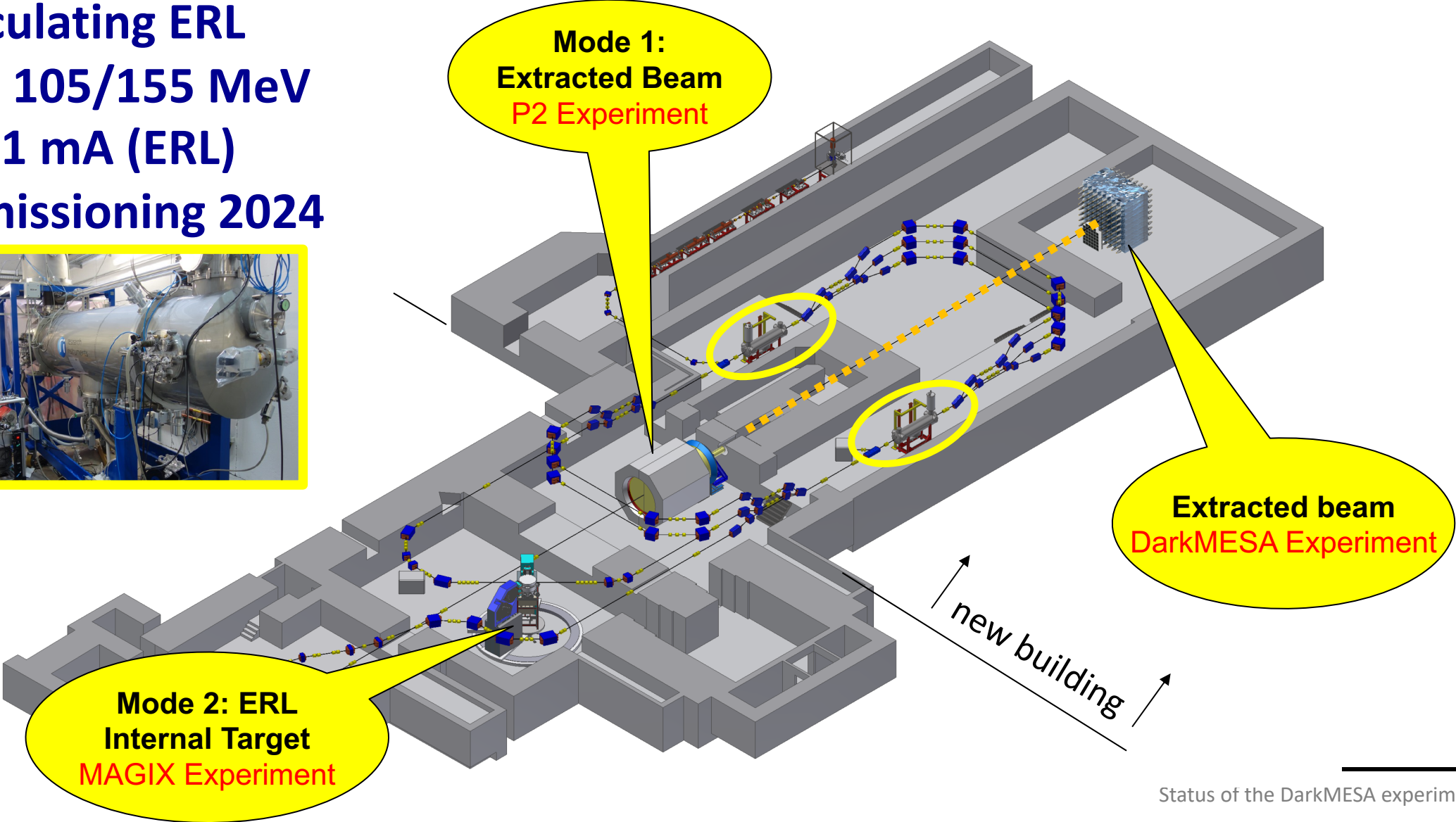
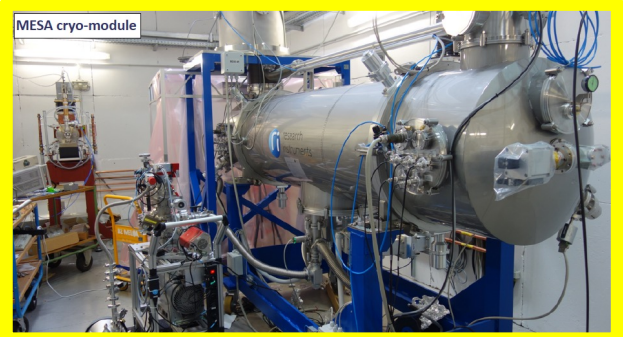
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High-Resolution MAGIX Spectrometers

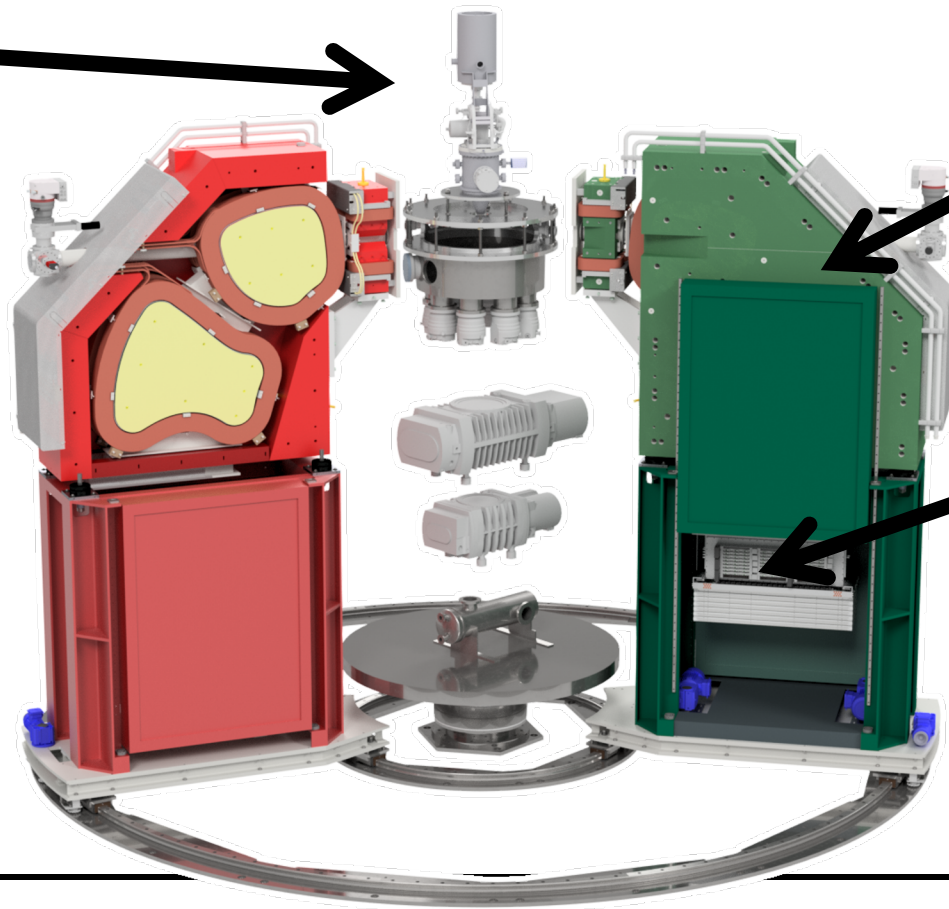
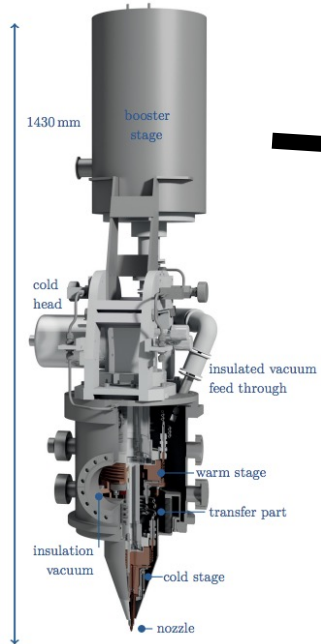
Operation of a high-intensity (polarized) ERL beam
in conjunction with light internal target
→ a novel technique in nuclear and particle physics



High-Resolution MAGIX Spectrometers

Operation of a high-intensity (polarized) ERL beam
in conjunction with light internal target
→ a novel technique in nuclear and particle physics

NIM A1013 (2021)



Two identical spectrometers

- Two dipoles each
- One quadrupole each

TPC-based focal plane detector

- 10^{-4} momentum resolution
- Requires spatial resolution of $O(100 \mu\text{m})$
- Open field cage
- GEM readout

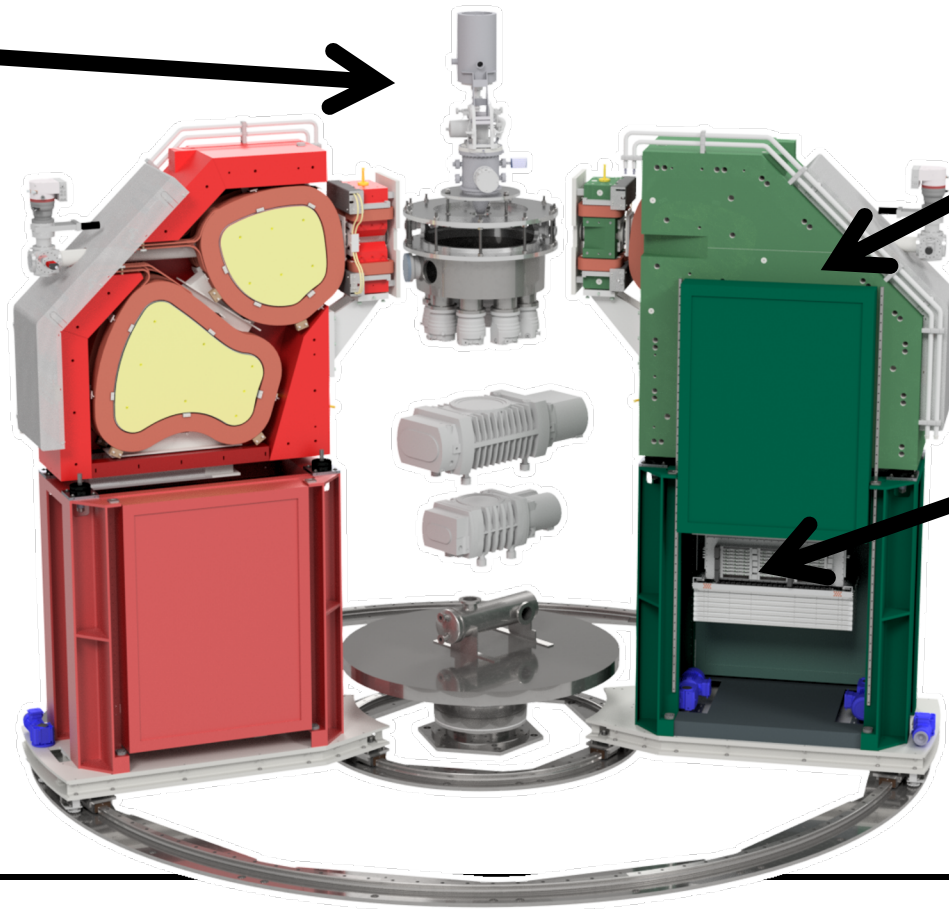
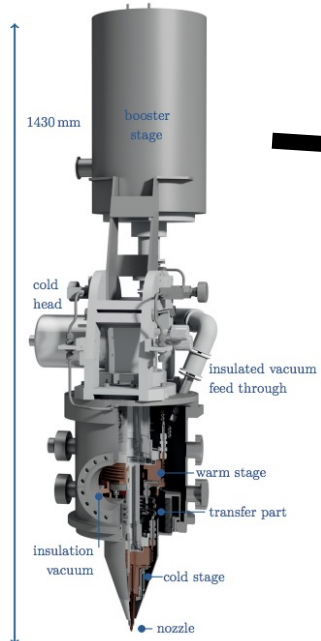
Supersonic cryogenic gas jet target

- Windowless environment
- Commissioned at A1/MAMI
- Design density $10^{19}/\text{cm}^2$

High-Resolution MAGIX Spectrometers

Operation of a high-intensity (polarized) ERL beam
in conjunction with light internal target
→ a novel technique in nuclear and particle physics

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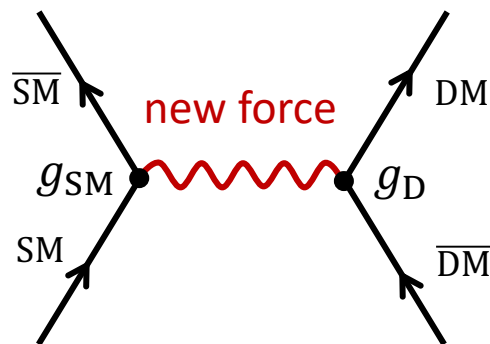
MESA Physics Programme

	ERL Mode MAGIX expt.	Extracted Beam Mode P2 expt.	Extracted Beam Mode DarkMESA.
Nucleon From Factors	✓		
EW Mixing Angle		✓	
Nuclear Astrophysics	✓ $^{12}\text{C} (\alpha, \gamma) ^{16}\text{O}$	neutron skin of nuclei ✓	
Few Body Physics	✓		
Light Dark Matter Search	✓		✓

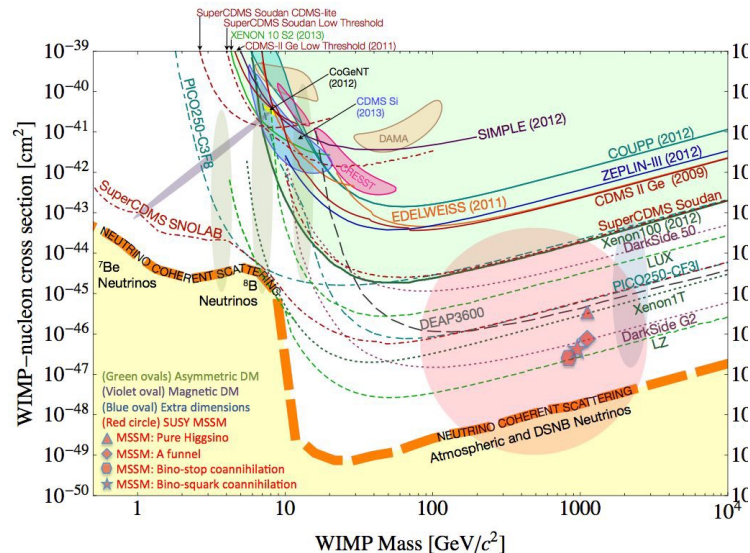
Light Dark Matter (LDM)

Light Dark Matter

$$\text{MeV} < m_{\text{DM}} < \sim \text{GeV}$$

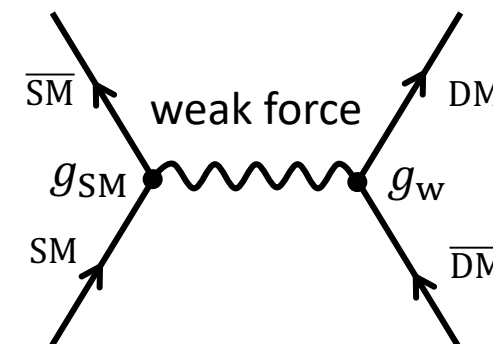


- Thermal relic targets exist for the MeV-GeV mass scale
- LDM requires a beyond SM force
- Rich phenomenology of portals: vector, higgs, neutrino, axion



WIMPs

$$\text{GeV} < m_{\text{DM}} < \sim \text{TeV}$$



- Matching relic abundance for the electroweak mass scale
- WIMPs require only SM interaction
- No positive evidence after LHC and galactic DM searches

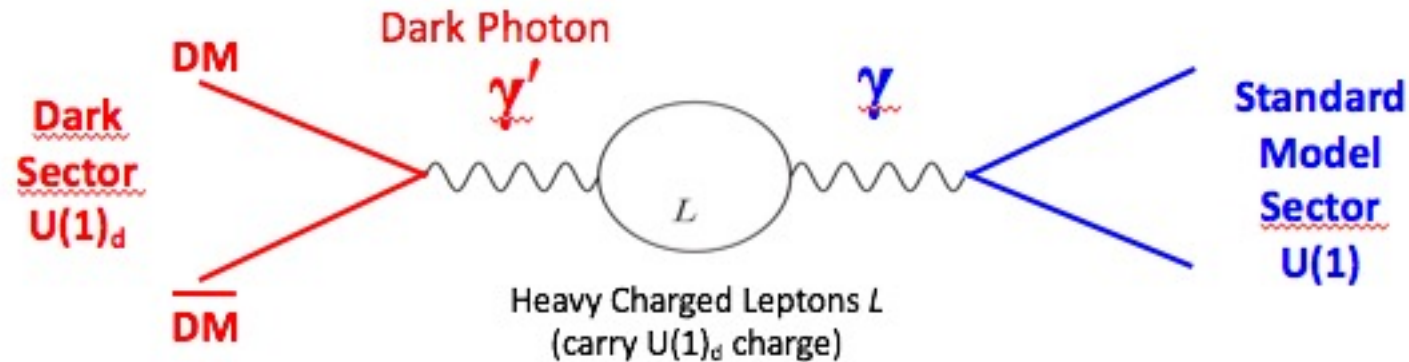
Vector Mediator: Dark Photon

Model 1: $m_{\gamma'} \ll m_{\text{DM}}$

Dark Photon decaying into SM particles – coupling ϵ
 → MAGIX

$$\frac{1}{2} \epsilon_Y F_{\mu\nu}^Y F'^{\mu\nu}$$

Holdom [1986]

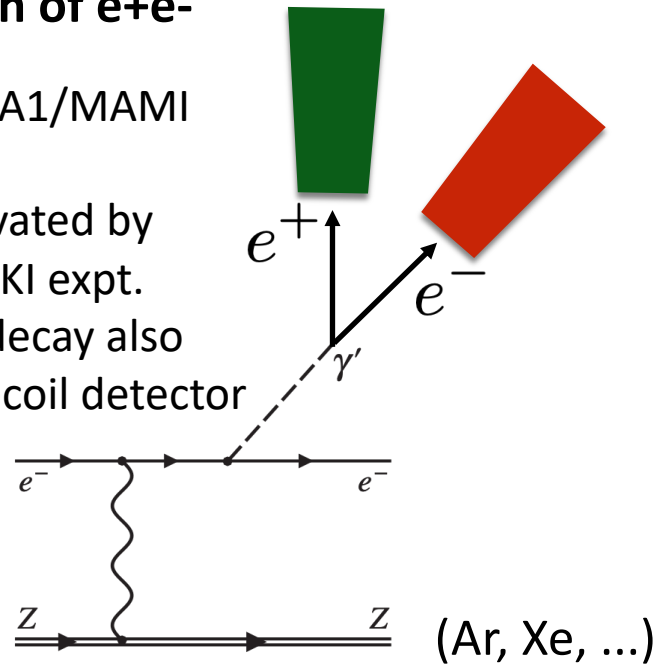


Dark Photon Search at MAGIX

Visible dark photon search

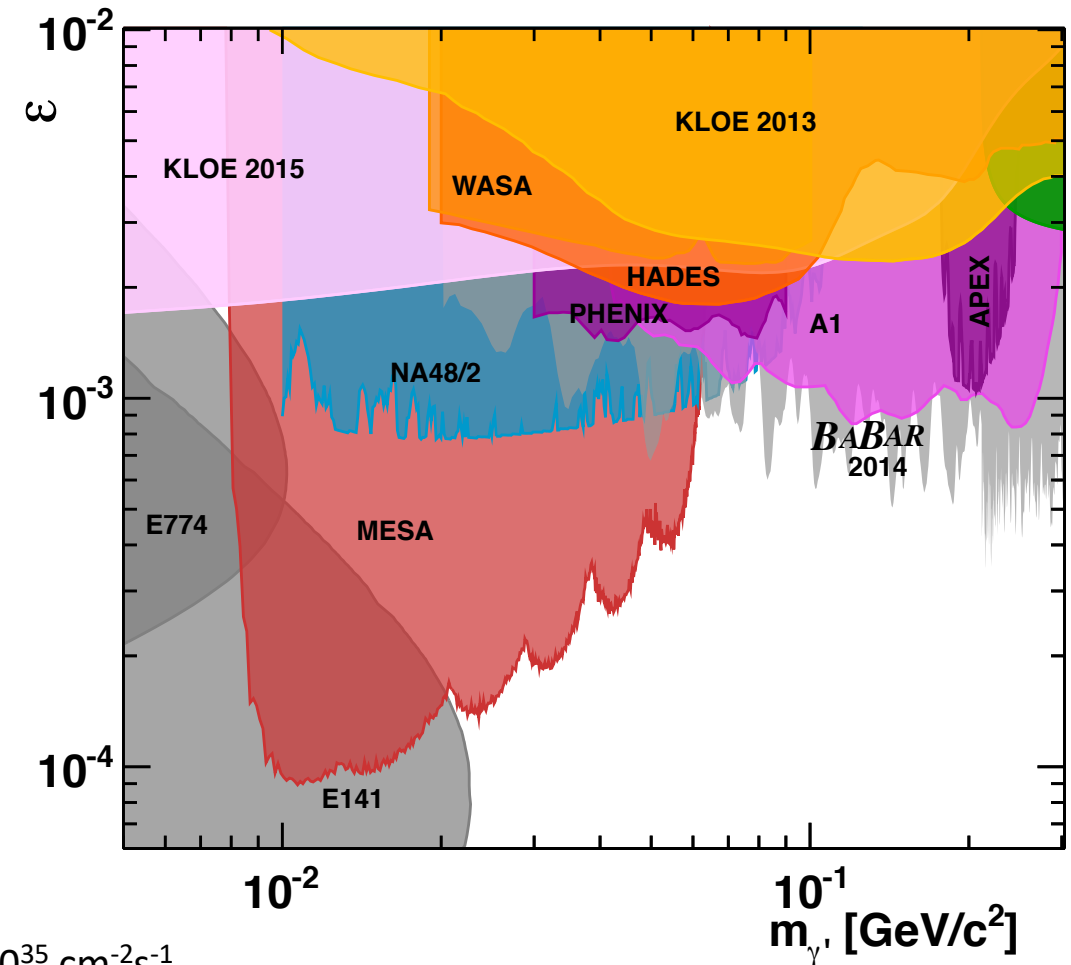
→ Coincident detection of e^+e^-

- Method pioneered by A1/MAMI and APEX/JLAB
- Parameter range motivated by X17 anomaly of ATOMKI expt.
- Invisible dark photon decay also possible thanks to Si recoil detector



Features:

- Xe gas target
- Luminosity $10^{35} \text{ cm}^{-2}\text{s}^{-1}$
- 6 month of data taking



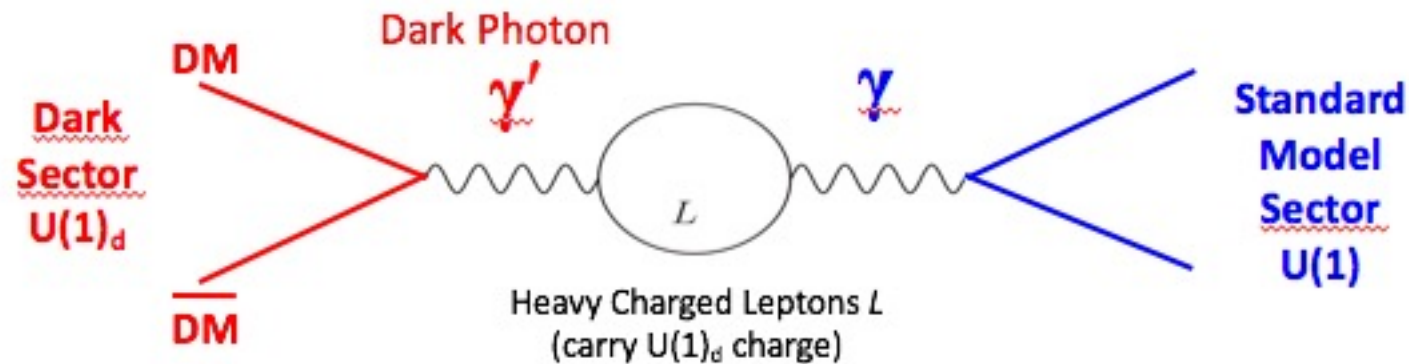
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 → MAGIX visible decay

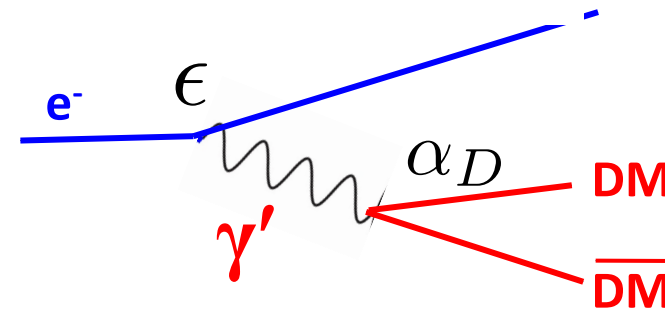
$$\frac{1}{2} \epsilon_Y F_{\mu\nu}^Y F'^{\mu\nu}$$

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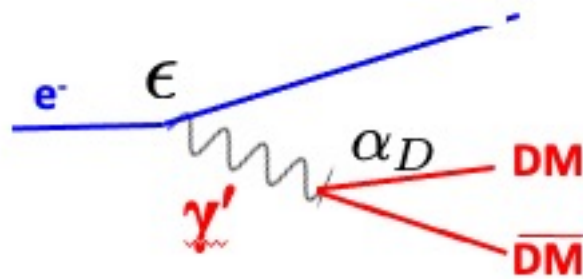


Model 2: $m_{\gamma'} > 2m_{\text{DM}}$

Dark Photon decaying into Dark Matter
 → MAGIX invisible decay and DarkMESA

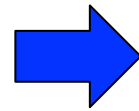


LDM Portals – A rich Phenomenology

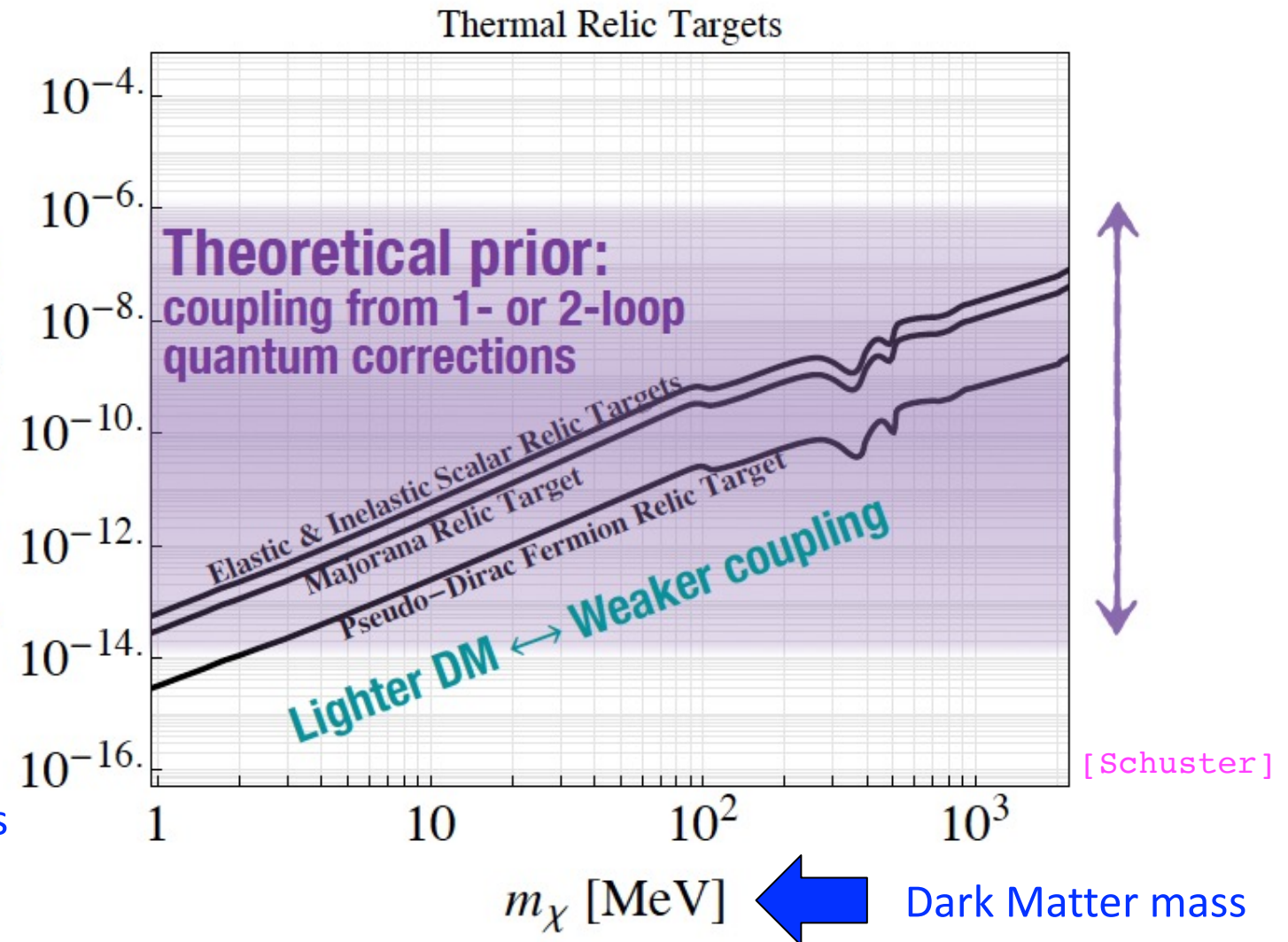


Combination of

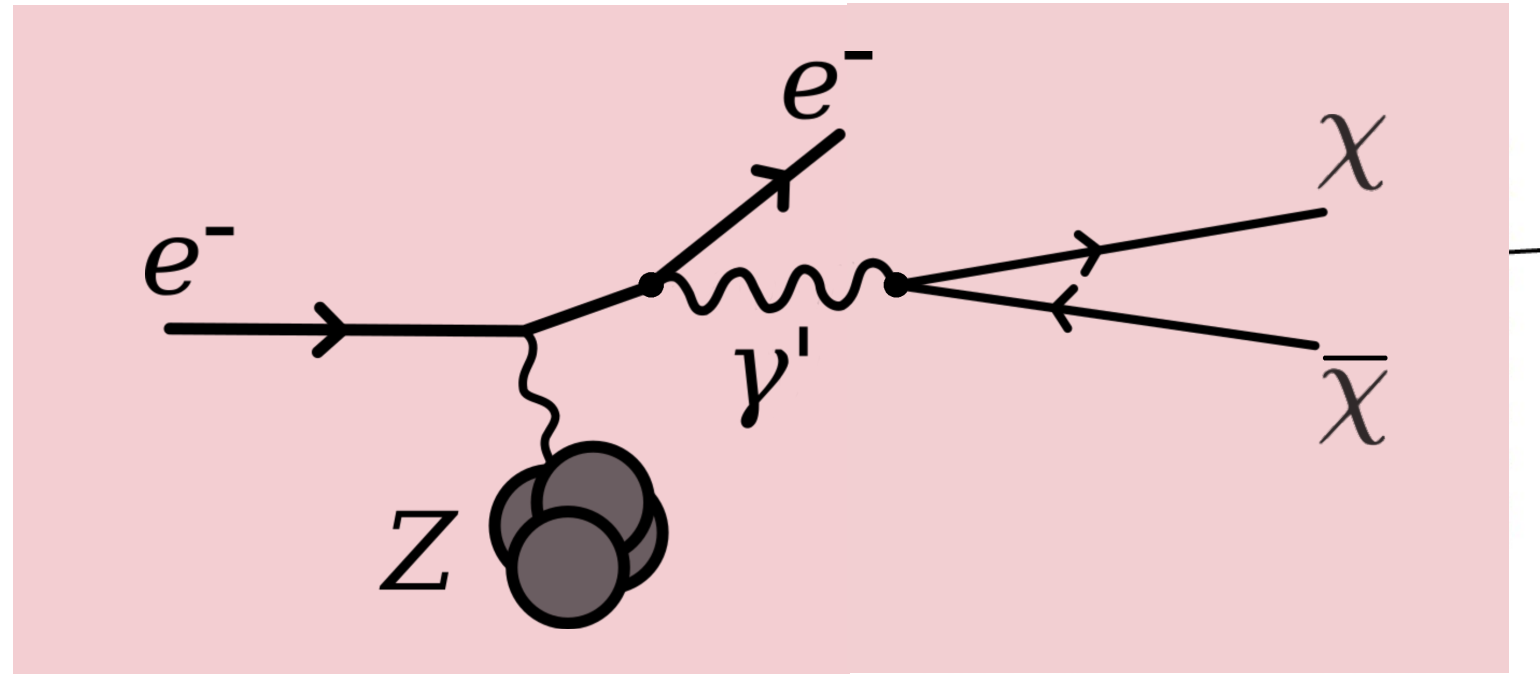
- Dark photon -Electron coupling
- Dark photon – DM coupling
- Ratio of DM/Dark photon masses



$$y = \epsilon^2 \alpha_D (m_\chi / m_{A'})^4$$

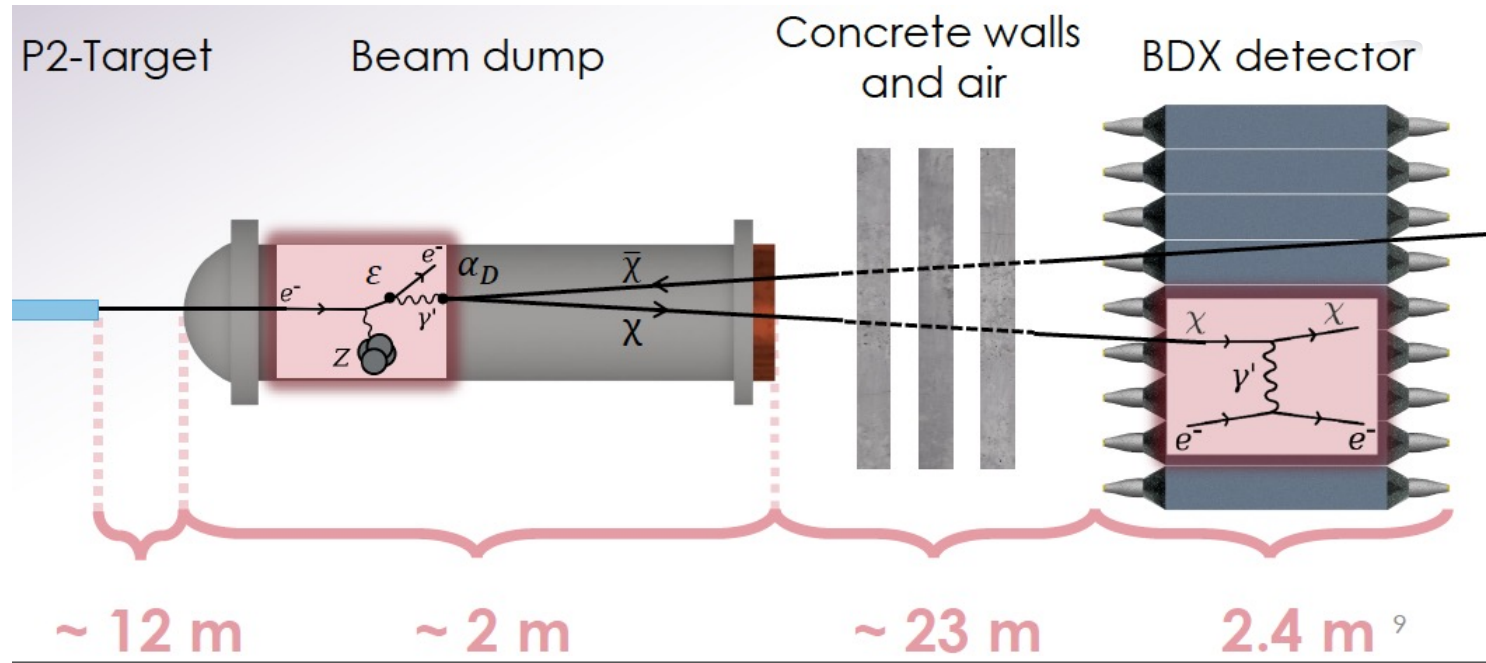


Model 2: $m_{\gamma'} > 2m_{\text{DM}}$



**Electron Scattering (MESA) on Beam Dump
→ Collimated pair of Dark Matter particles !**

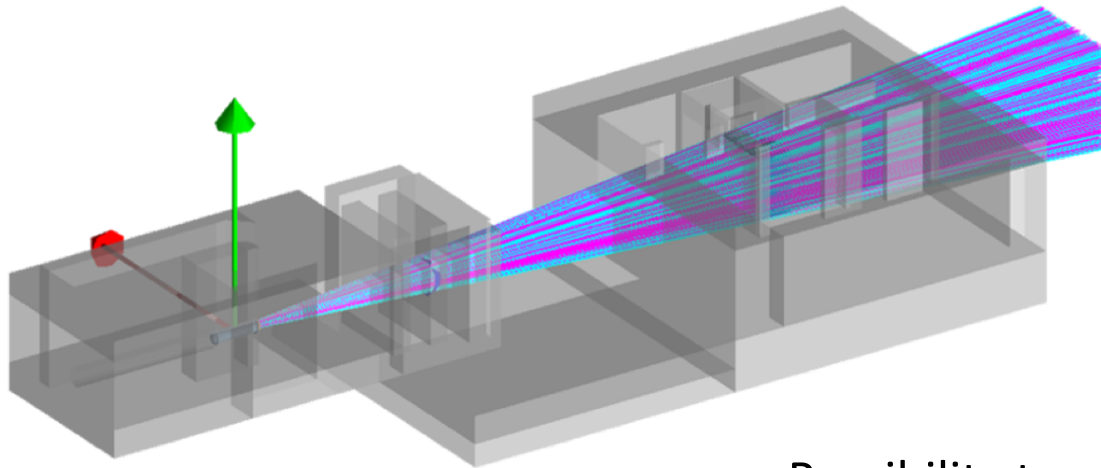
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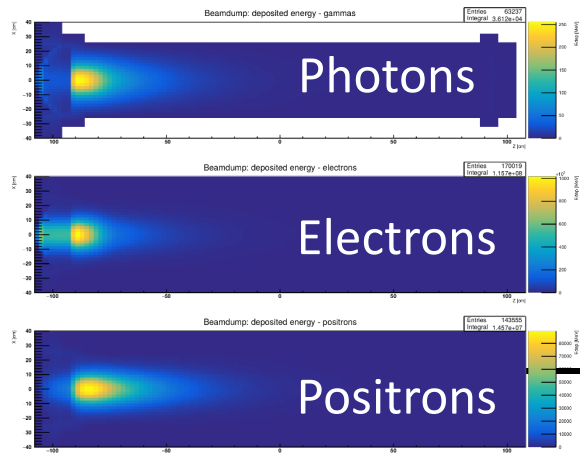
**Electron Scattering (MESA) on Beam Dump
→ Collimated pair of Dark Matter particles !**

Beam Dump Experiment DarkMESA

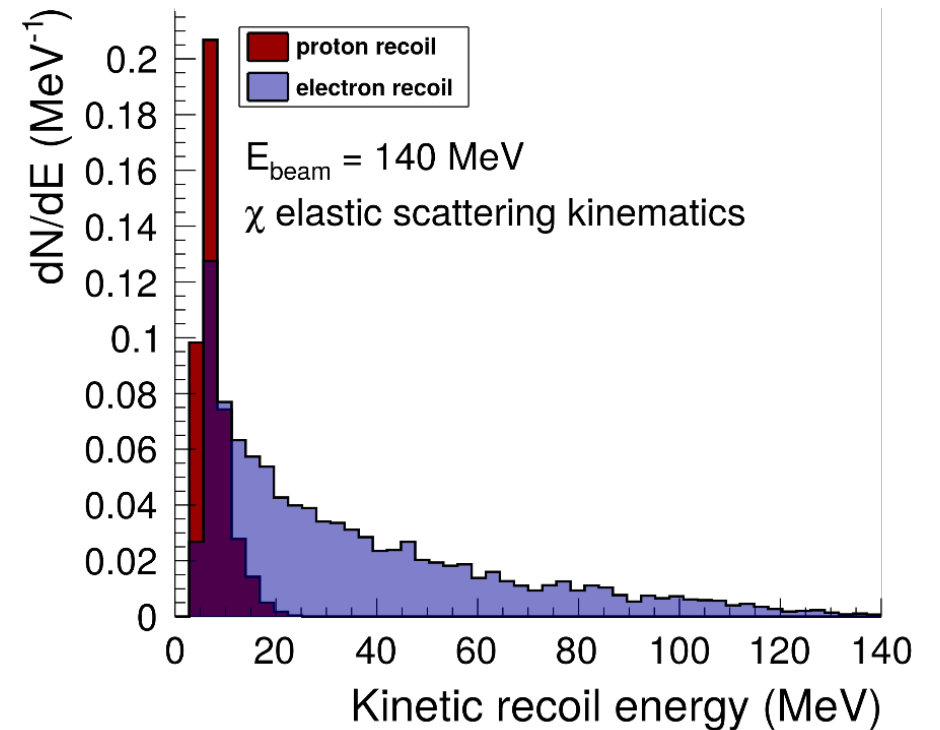
Full GEANT4 simulation:
 P2 target, beam dump, BDX detector volume, walls etc.



Possibility to place a second detector off-axis



→ LDM interaction with
 DarkMESA material (electron recoil)



Detector Concept for DarkMESA

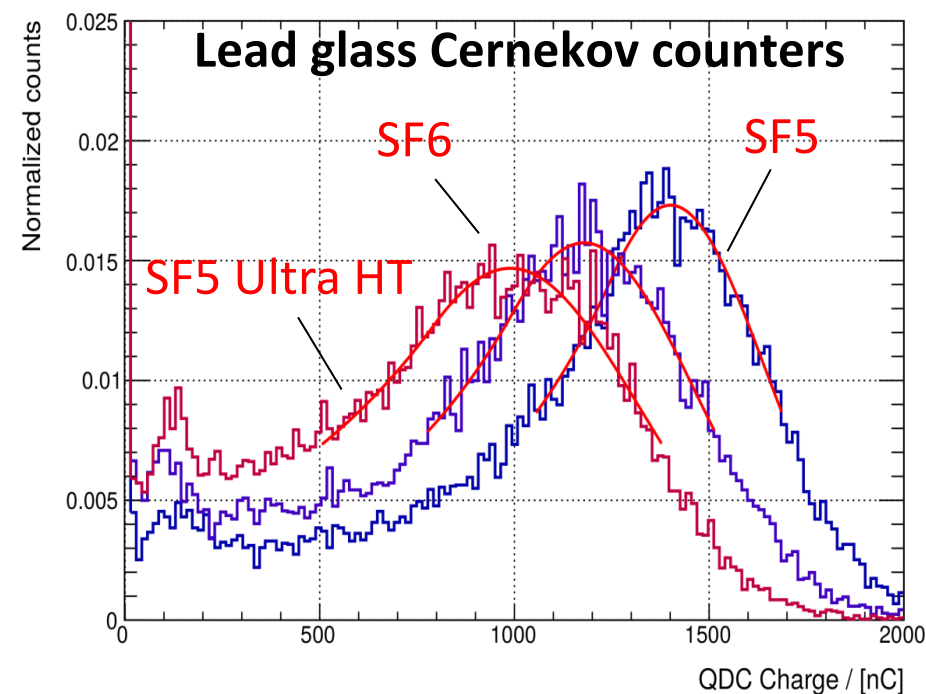
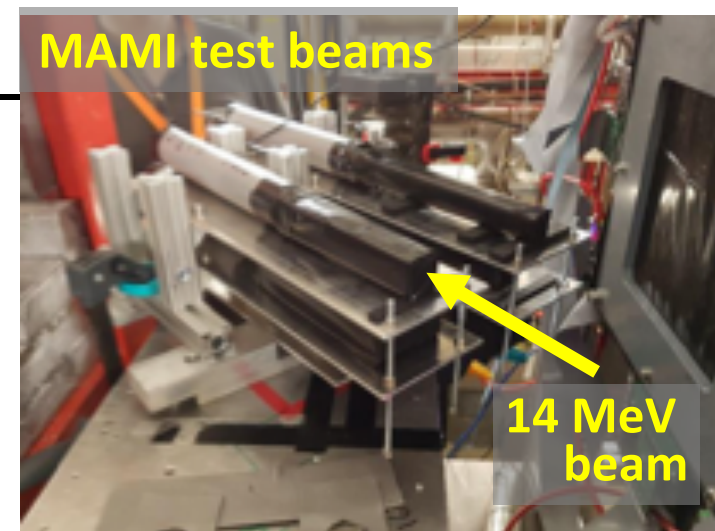
Ideal Requirements:

1. Electron Detection $>$ few MeV
2. Large Surface (Acceptance)
3. Large thickness (Int. Prob.)
4. Reliability (long running time)
5. Background rejection
 - Cosmics
 - Natural Backgrounds
 - Beam Backgrounds (n, ν)



Baseline Concept

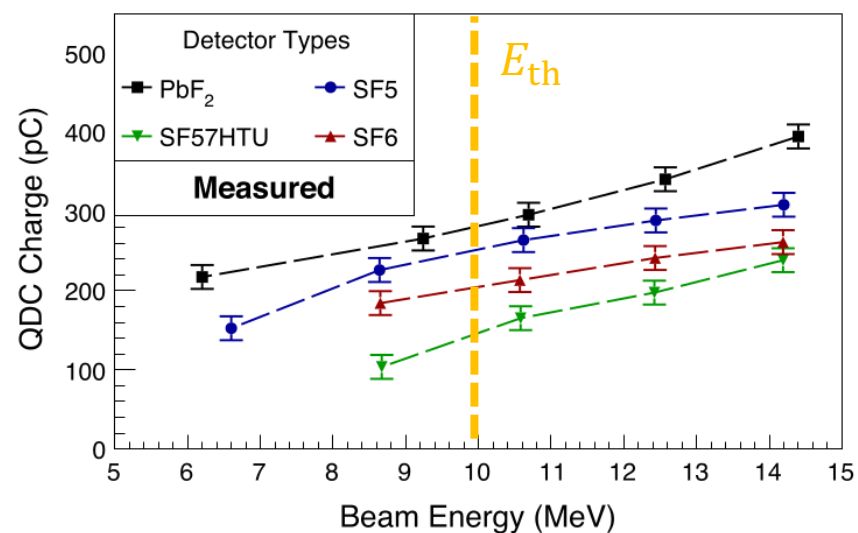
- Inorganic crystal calorimeter
- Cherenkov (fast, no neutrons)
 - Scintillator (higher light yield)



Detector Concept for DarkMESA

PbF₂ (old A4/MAMI crystals) and the Pb-glass SF5 offer proper electron sensitivity and neutron insensitivity

Electron sensitivity study

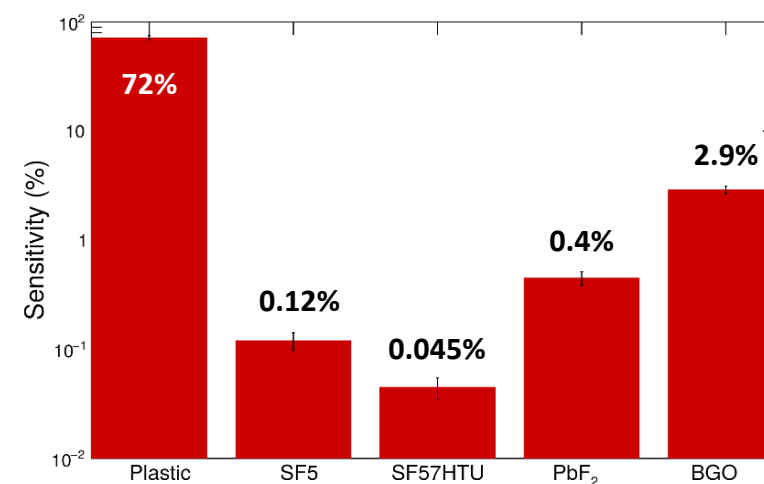


Electron beam tests

Mirco Christmann, et al.

- NIM A 958 (2020) 162398
- NIM A 960 (2020) 163665

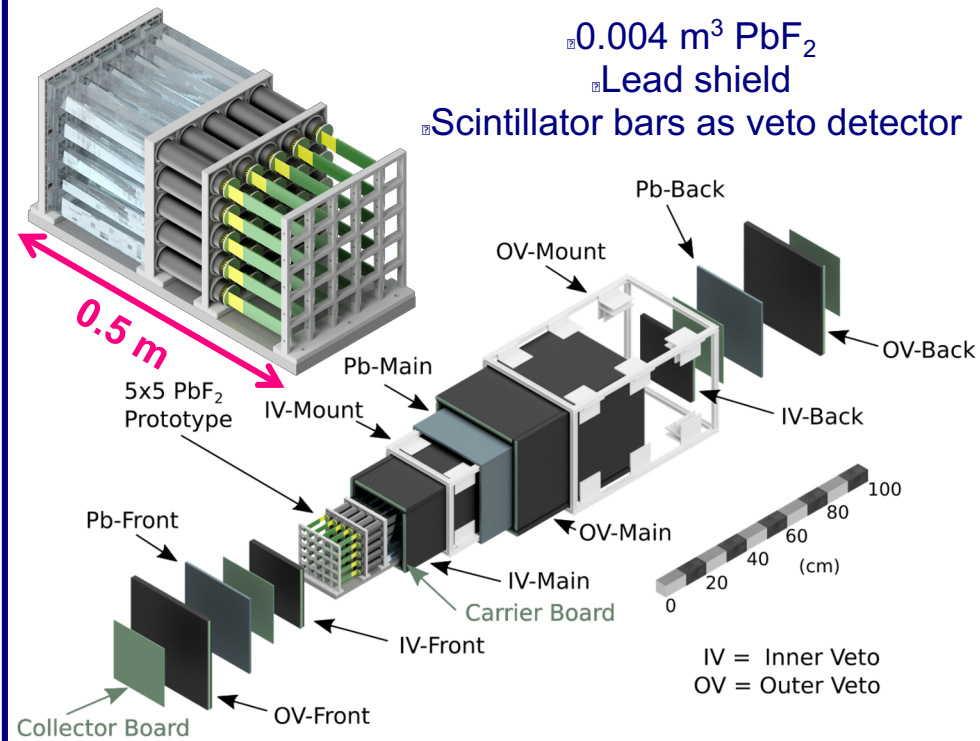
Neutron sensitivity study



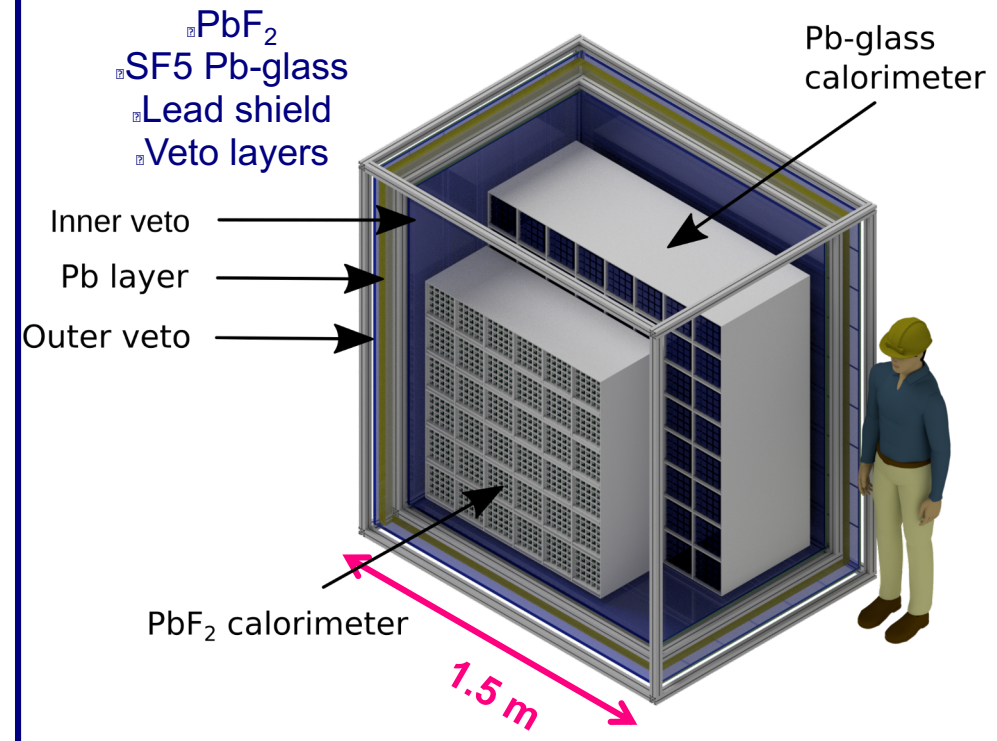
Neutron sensitivity study with an AmBe source

Detector Concept for DarkMESA

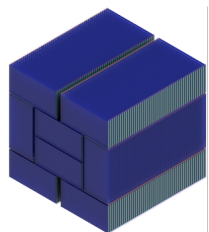
Phase A (prototype)



Phase B (increase volume / yield)



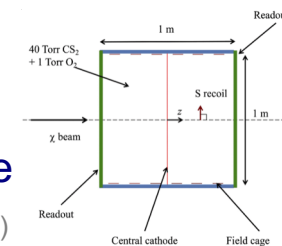
Phase C (range extension)



Radiation protection glas
e.g., SCHOTT-RD30

DRIFT -Time Projection Chamber
CS₂ at 50 mbar ~ 1 m³ active volume

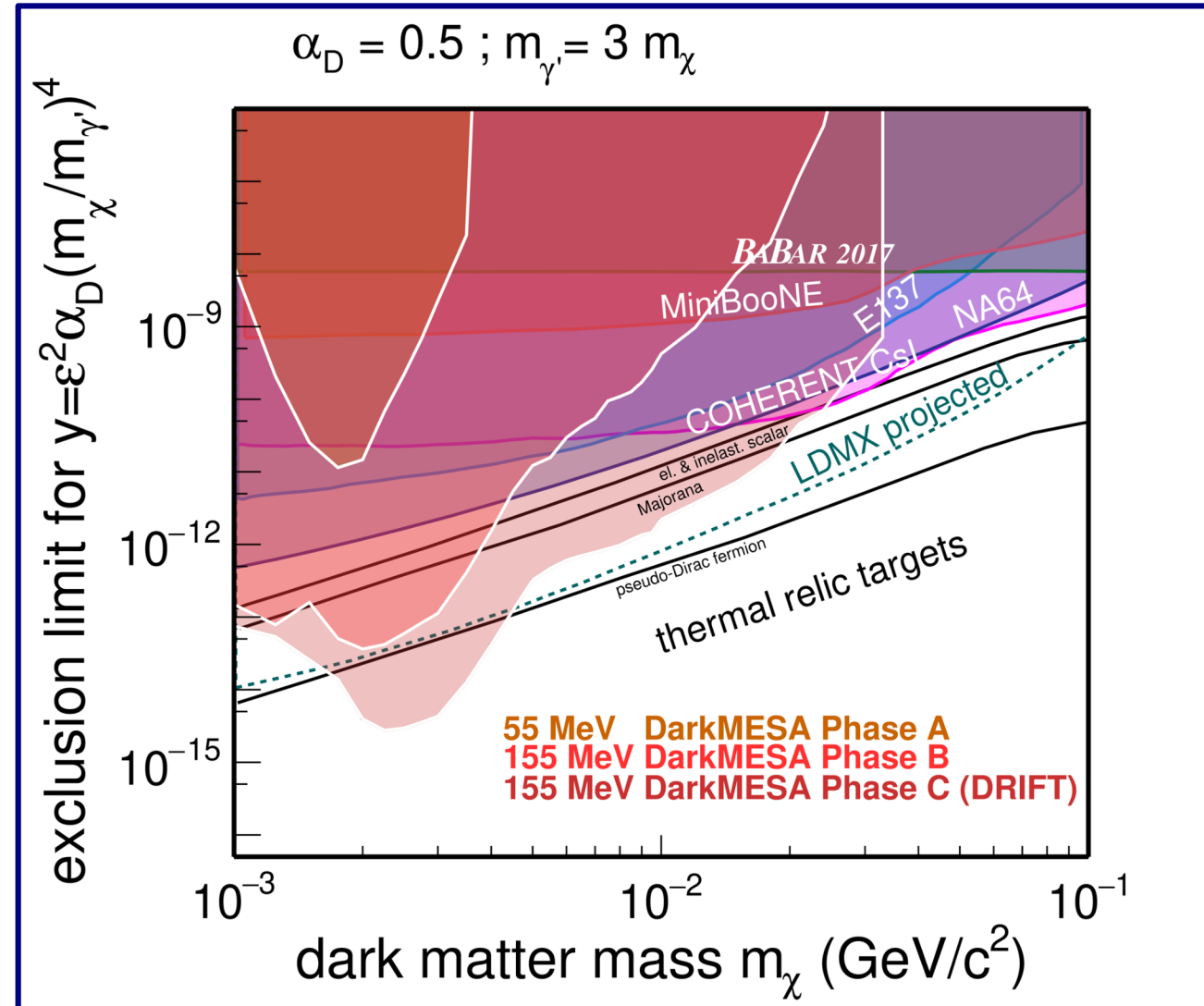
PRD 99 (2019) 061301(R)



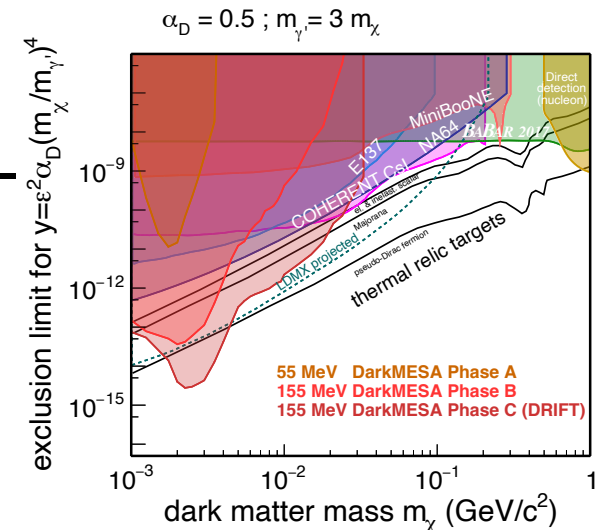
DarkMESA Physics Reach

Run plan for DarkMESA

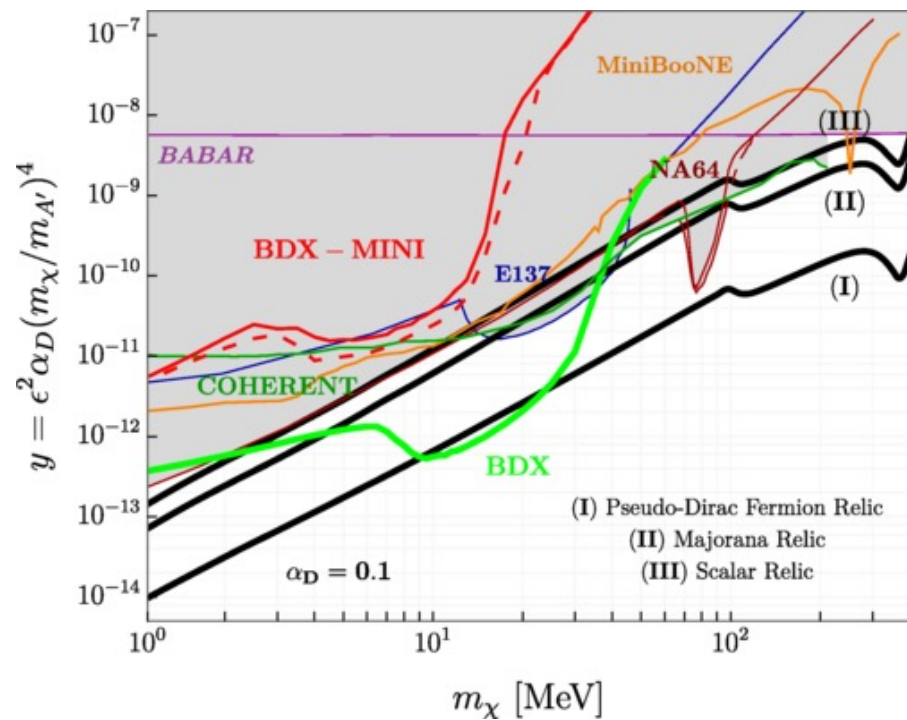
Phase	Detector	Period	Time	EOT
A	Prototype	1.-4. year	2,200 h	$7.42 \cdot 10^{21}$
B	PbF ₂ , SF5	4.-6. year	6,600 h	$2.22 \cdot 10^{22}$
C	+TPC	7.-12. year	13,200 h	$4.45 \cdot 10^{22}$



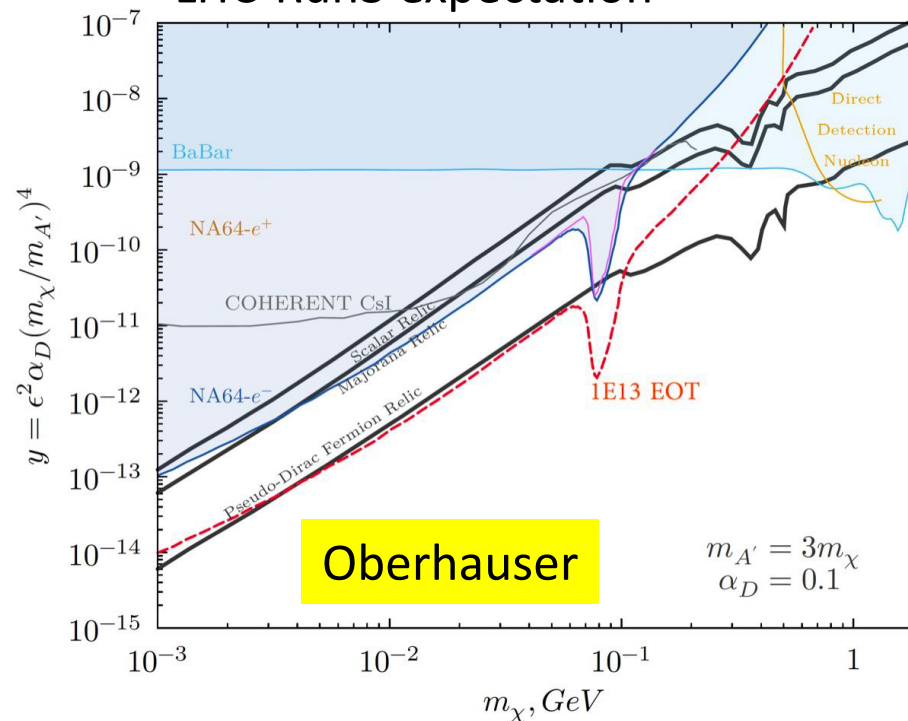
Competitors to DarkMESA



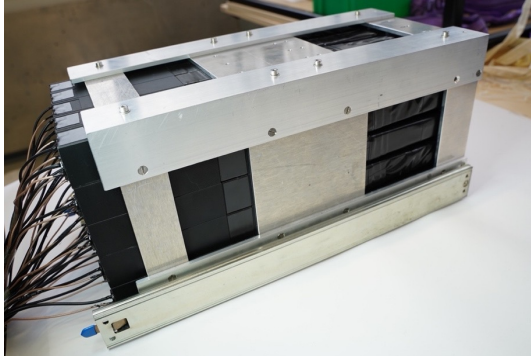
complementarity to BDX@JLAB



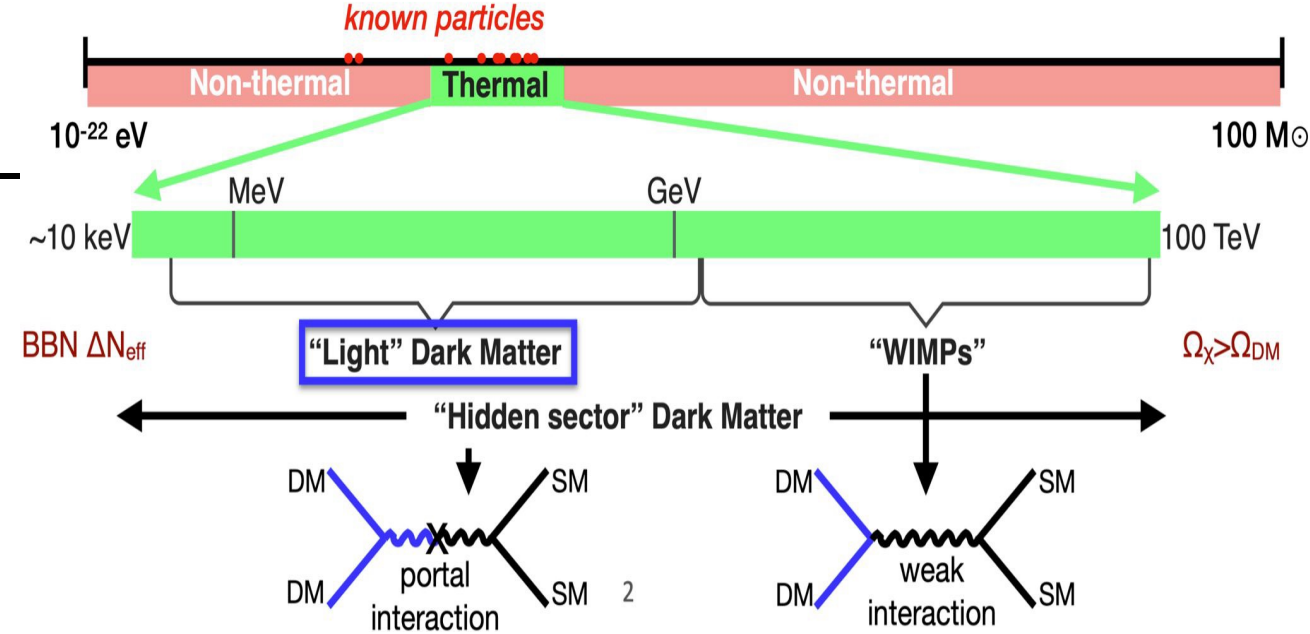
strong competition from NA64
LHC-Run3 expectation



Conclusions



Phase A
DarkMESA detector



- DarkMESA allows to test parameter space in well-motivated thermal DarkMatter model (complementary to WIMP model, for which room for discovery is narrowing)
- Based on rather conservative calorimeter design
Phase C might need more inovative detector aspects
- Strong competition world-wide, however complementarity (direct vs. indirect searches, different probes)

Potential to make a very fundamental discovery in physics !

DARK MESA