

# USTC-PNP-Nuclear Physics Mini Workshop Series

## Polarization of Unstable Light (Hyper-) Nuclei in Heavy-Ion Collisions

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arXiv:2405.12015 (Accepted by PRL)

# Outline

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- 1. From polarization of hadrons to polarization of loosely-bound nuclei**
- 2. Production of light (hyper-)nuclei in heavy-ion collisions**
- 3. (Anti-)Hypertriton polarization and its spin structure**
- 4. Discussions: Effects of baryon spin correlations & Polarization of nucleons**
- 5. Summary and outlook**

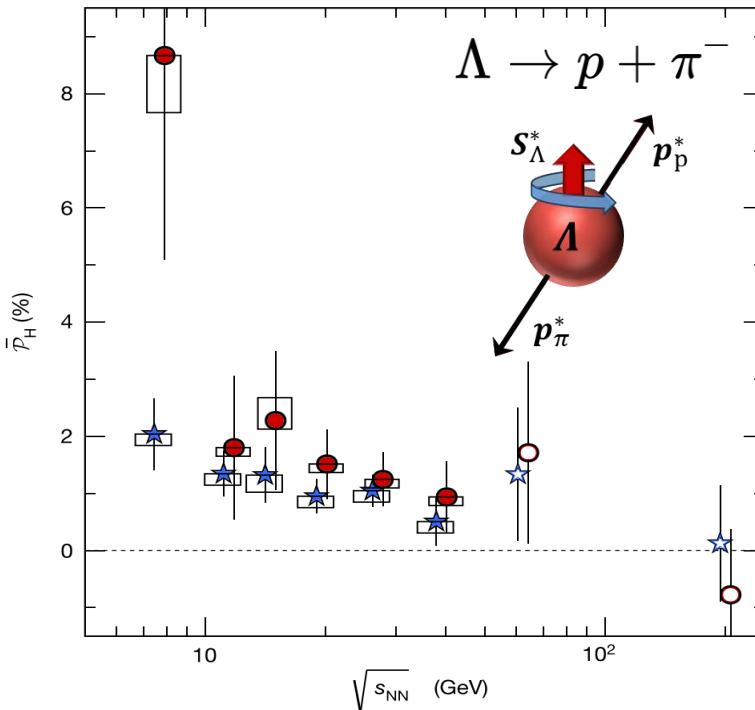
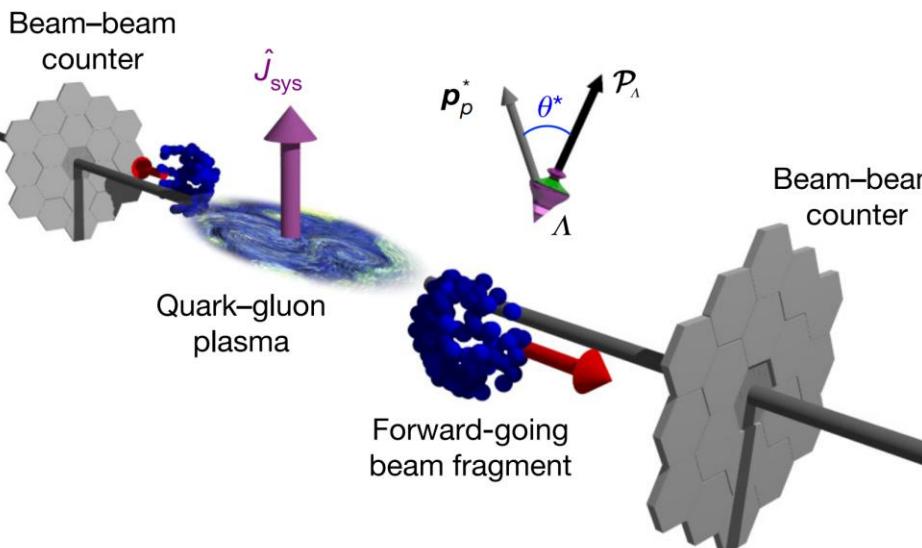
# 1. Polarization of hadrons in relativistic heavy-ion collisions (1)

STAR, Nature 548, 62 (2017)

Z. T. Liang and X. N. Wang PRL 94, 102301 (2005)

F. Becattini, F. Piccinini, and J. Rizzo, PRC 77, 024906 (2008)

F. Becattini, M. Buzzegoli, T. Niida, S. Pu, and A. Tang, Int.J.Mod.Phys.E 33 (2024) 06, 2430006



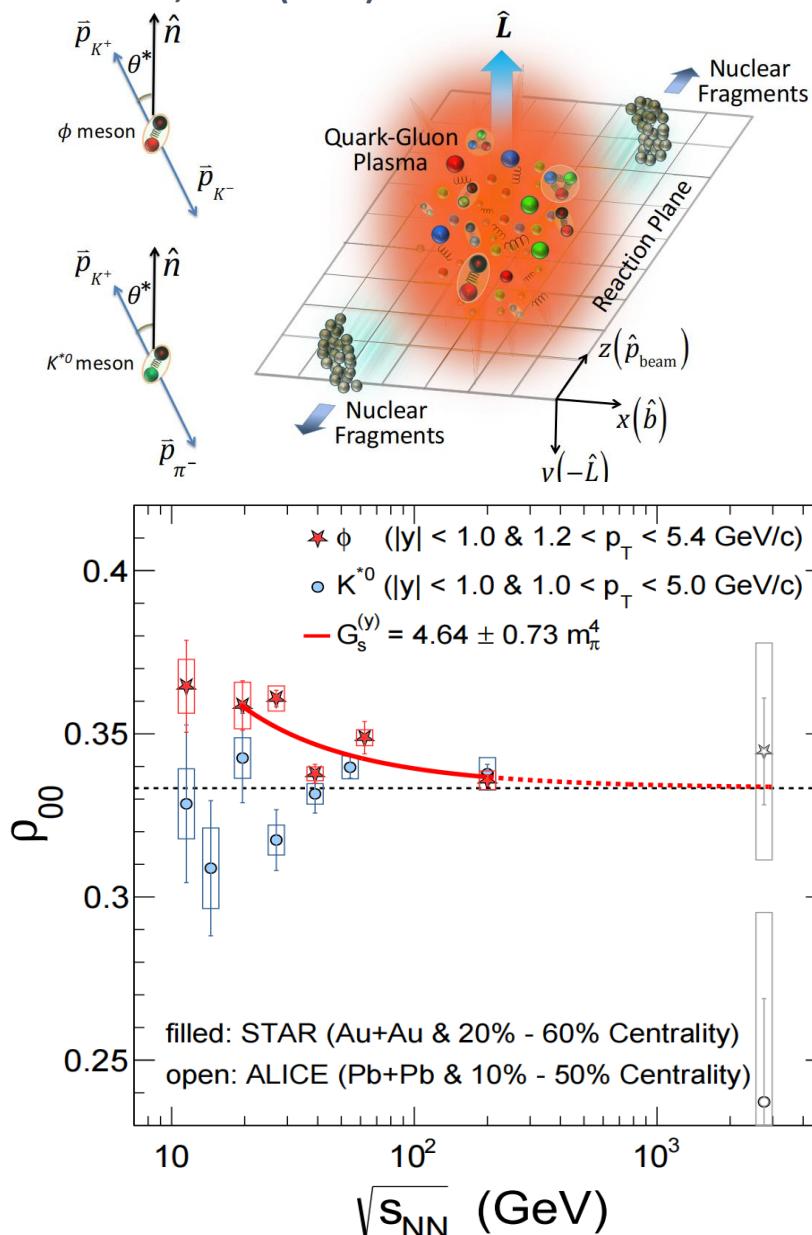
$$\frac{dN}{d \cos \theta^*} = \frac{1}{2} (1 + \alpha_H |\mathcal{P}_H| \cos \theta^*)$$
$$P_H = \frac{8}{\pi \alpha_H} \frac{\langle \sin(\Psi_1^{\text{obs}} - \phi_p^*) \rangle}{\text{Res}(\Psi_1)}$$

$$\omega \approx k_B T (\bar{\mathcal{P}}_{\Lambda} + \bar{\mathcal{P}}_{\bar{\Lambda}}) / \hbar$$
$$\approx (9 \pm 1) \times 10^{21} \text{ s}^{-1}$$

Spin polarization of Lambda hyperon → Vorticity of QGP

# 1. Polarization of hadrons in relativistic heavy-ion collisions (2)

STAR, Nature 614, 7947 (2023)



## Spin alignment of mesons

$$\frac{dN}{d(\cos\theta^*)} \propto (1 - \rho_{00}) + (3\rho_{00} - 1)\cos^2\theta^*$$



## Fluctuation/correlation of strong force field

X. L. Sheng et al., PRL 131, 042304 (2023)

$$G_s^{(y)} \equiv g_\phi^2 \left[ 3\langle B_{\phi,y}^2 \rangle + \frac{\langle \mathbf{p}^2 \rangle_\phi}{m_s^2} \langle E_{\phi,y}^2 \rangle - \frac{3}{2} \langle B_{\phi,x}^2 + B_{\phi,z}^2 \rangle - \frac{\langle \mathbf{p}^2 \rangle_\phi}{2m_s^2} \langle E_{\phi,x}^2 + E_{\phi,z}^2 \rangle \right]$$

## Quark-antiquark spin correlation

J. P. Lv et al., Phys.Rev.D 109 (2024) 11, 114003

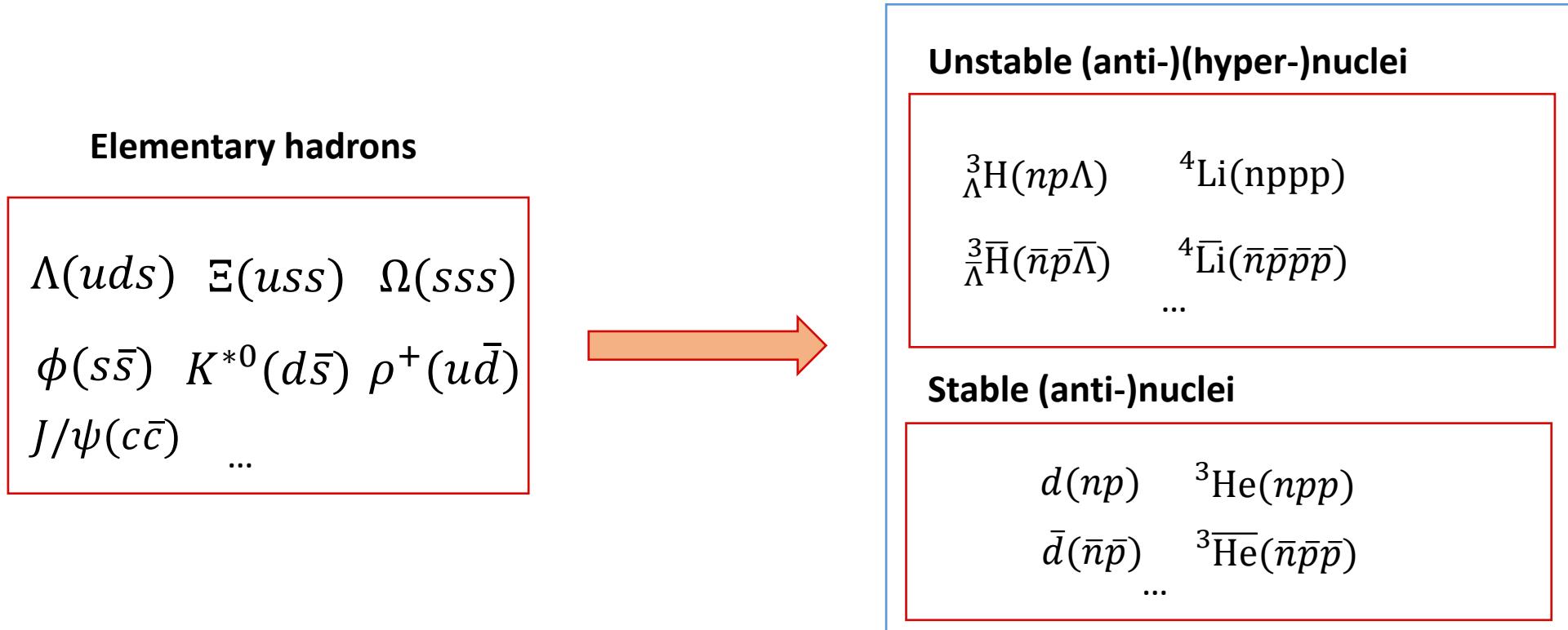
## Meson spectral property

F. Li and S. Liu, arXiv:2206.11890

Y. L. Yin, W. B. Dong, J. Y. Pang, S. Pu, and Q. Wang, Phys. Rev. C 110 (2024) 2, 024905

# 1. Polarization of light (anti-)(hyper-)nuclei

(3)



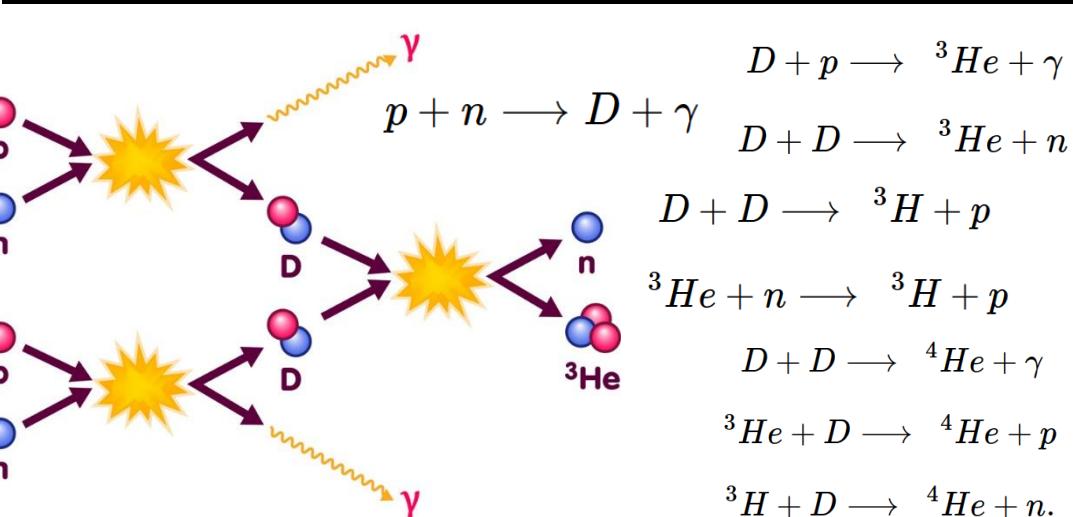
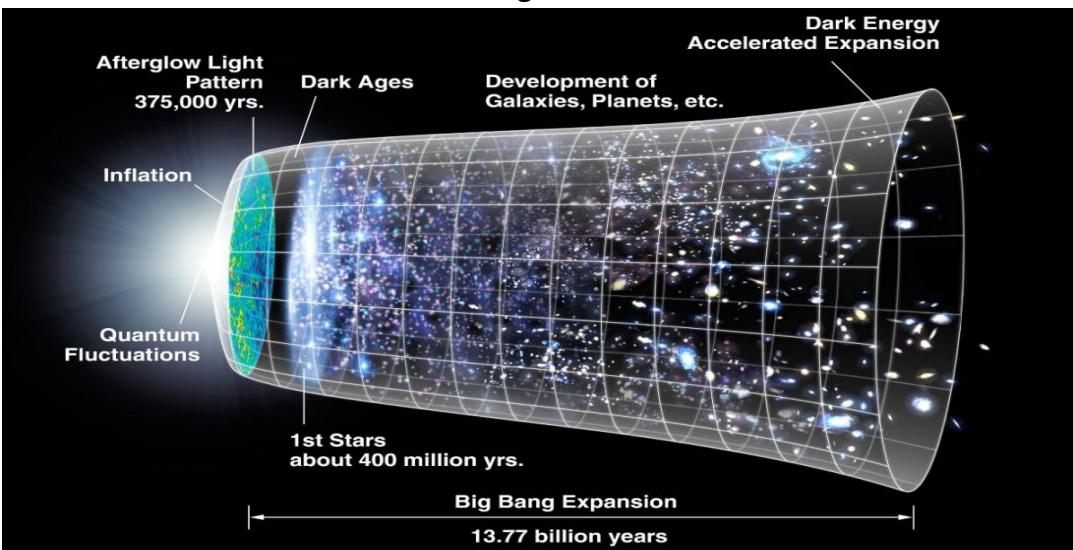
## 2. Little-Bang Nucleosynthesis

**Big-bang nucleosynthesis is responsible for the formation of light nuclei in our Universe.**

$$t \sim 100 \text{ s}, kT < 1 \text{ MeV}$$

K. A. Olive et al., Phys. Rept. 333, 389–407 (2000);

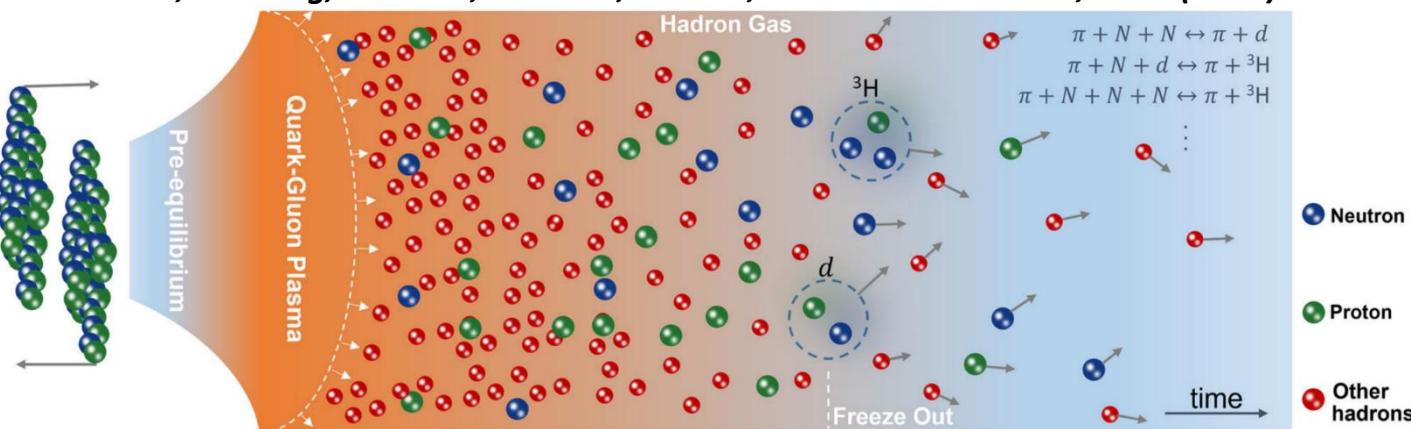
《The First Three Minutes》 S. Weinberg



**Synthesis of antimatter nuclei in little bangs of relativistic heavy-ion collisions**

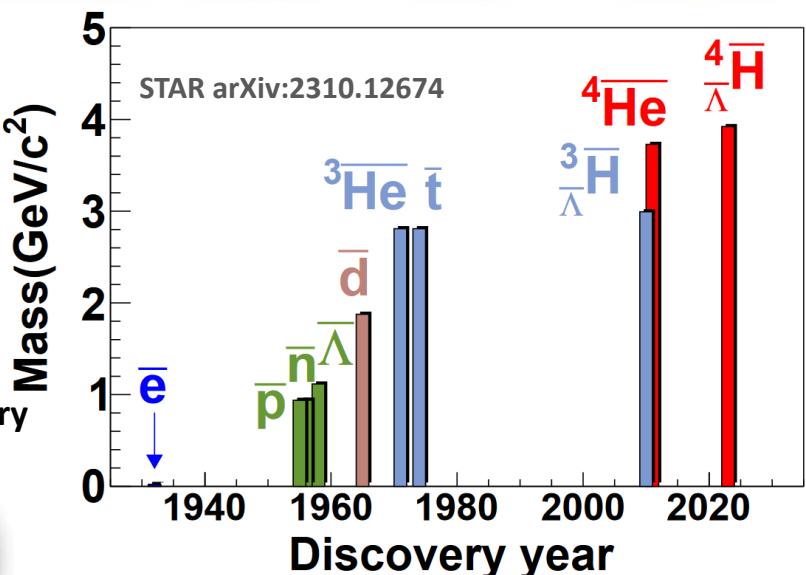
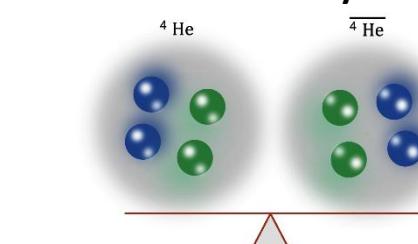
$$t \sim 10^{-22} \text{ s}, kT \sim 100 \text{ MeV}$$

K. J. Sun, R. Wang, C. M. Ko, Y. G. Ma, C. Shen, Nature Commun. 15, 1074 (2024)



**Antimatter factory**

**Matter-antimatter asymmetry**



J. Chen et al., Phys. Rep. 760, 1 (2018);

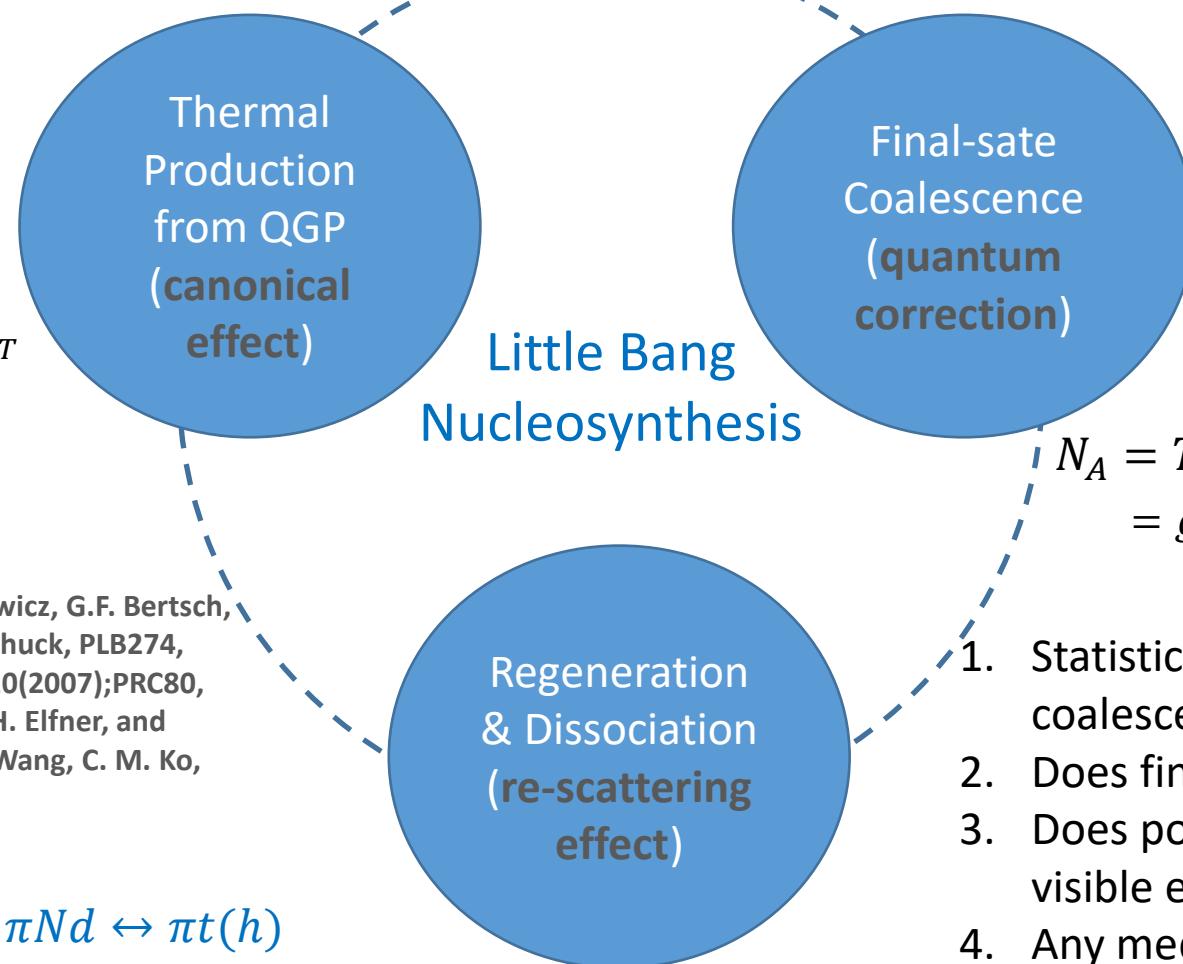
P. Braun-Munzinger and B. Donigus NPA987, 144 (2019)

## 2. Production Mechanisms : When? Where? How?

(5)

A. Andronic, P. Braun-Munzinger, J. Stachel, H. Stöcker, PLB 697, 203 (2011)  
A. Andronic, P. Braun-Munzinger, K. Redlich, J. Stachel, Nature 561, 321 (2018)  
V. Vovchenko et al., PLB785, 171 (2018);PLB800, 135131 (2020) (Saha Eq.);  
T. Neidig et al., PLB827,136891(2022)(Rate Eq.);...

$$N_A \approx g_A V (2\pi m_A T)^{3/2} e^{(A\mu_B - m_A)/T}$$



A.Z. Mekjian, PRC17,1051 (1978); P. Danielewicz, G.F. Bertsch, NPA533, 712 (1991); P. Danielewicz and P. Schuck, PLB274, 268 (1992); Y. Oh and C. M. Ko PRC76, 054910(2007);PRC80, 064902(2009);D. Oliinychenko, L. G. Pang, H. Elfner, and V. Koch, PRC99, 044907 (2019); K. J. Sun, R. Wang, C. M. Ko, Y. G. Ma, and C. Shen, 2207.12532(2022)



J. I. Kapusta, Phys. Rev. C 21, 1301 (1980)  
H. Sato and K. Yazaki, PLB98, 153 (1981);  
E. Remler, Ann. Phys. 136, 293 (1981);  
M. Gyulassy, K. Frankel, and E. Remler, NPA402,596 (1983);  
S. Mrowczynski, J. Phys. G 13, 1089 (1987);  
S. Leupold and U. Heinz, PRC50, 1110 (1994);  
R. Scheibl and U. W. Heinz, PRC59. 1585(1999);  
K. J. Sun, C. M. Ko and B. Donigus, PLB 792, 132 (2019);  
S. Sombun et al., PRC99, 014901 (2019)  
F. Bellini et al., PRC99,054905(2019);  
PRC 103, 014907(2021);  
W. Zhao et al., PLB 820, 136571(2021);  
S. Wu et al., arXiv:2205.14302(2022);...

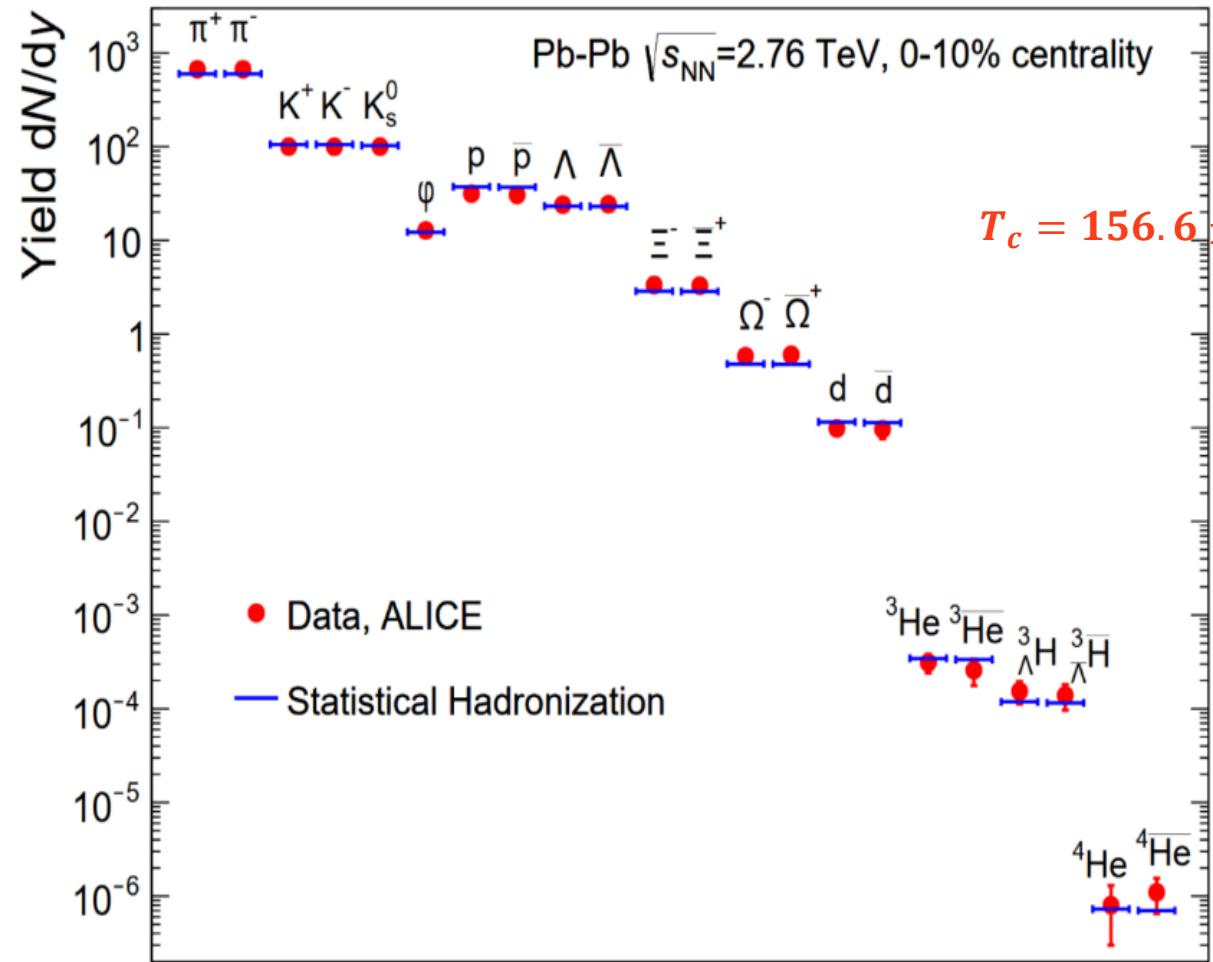
$$N_A = Tr(\hat{\rho}_s \hat{\rho}_A) \\ = g_c \int d\Gamma \rho_s(\{x_i, p_i\}) \times W_A (\{x_i, p_i\})$$

1. Statistical hadronization or final-state coalescence?
2. Does finite nuclear size play any role?
3. Does post-hadronization dynamics have visible effects?
4. Any medium effect?

## 2. Statistical hadronization

(6)

Andronic, Braun-Munzinger, Redlich, Stachel, Nature 561, 321 (2018)

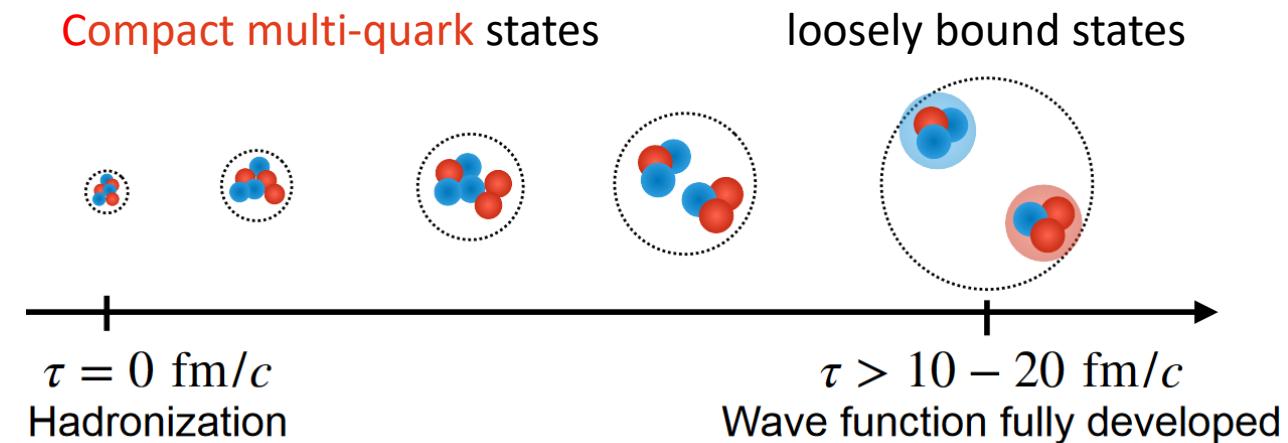


$$N_h \approx \frac{g_h V_C}{2\pi^2} m_h^2 T_C K_2\left(\frac{m_h}{T_C}\right)$$

$$\approx g_h V_C \left(\frac{m_h T_C}{2\pi}\right)^{3/2} e^{-m_h/T_C}$$

$T_C$ : Chemical freeze-out temperature, which is close to the chiral transition temperature (LQCD)

All (stable) particles including light (hyper)nuclei are produced at the QCD phase boundary and share a common chemical freeze-out

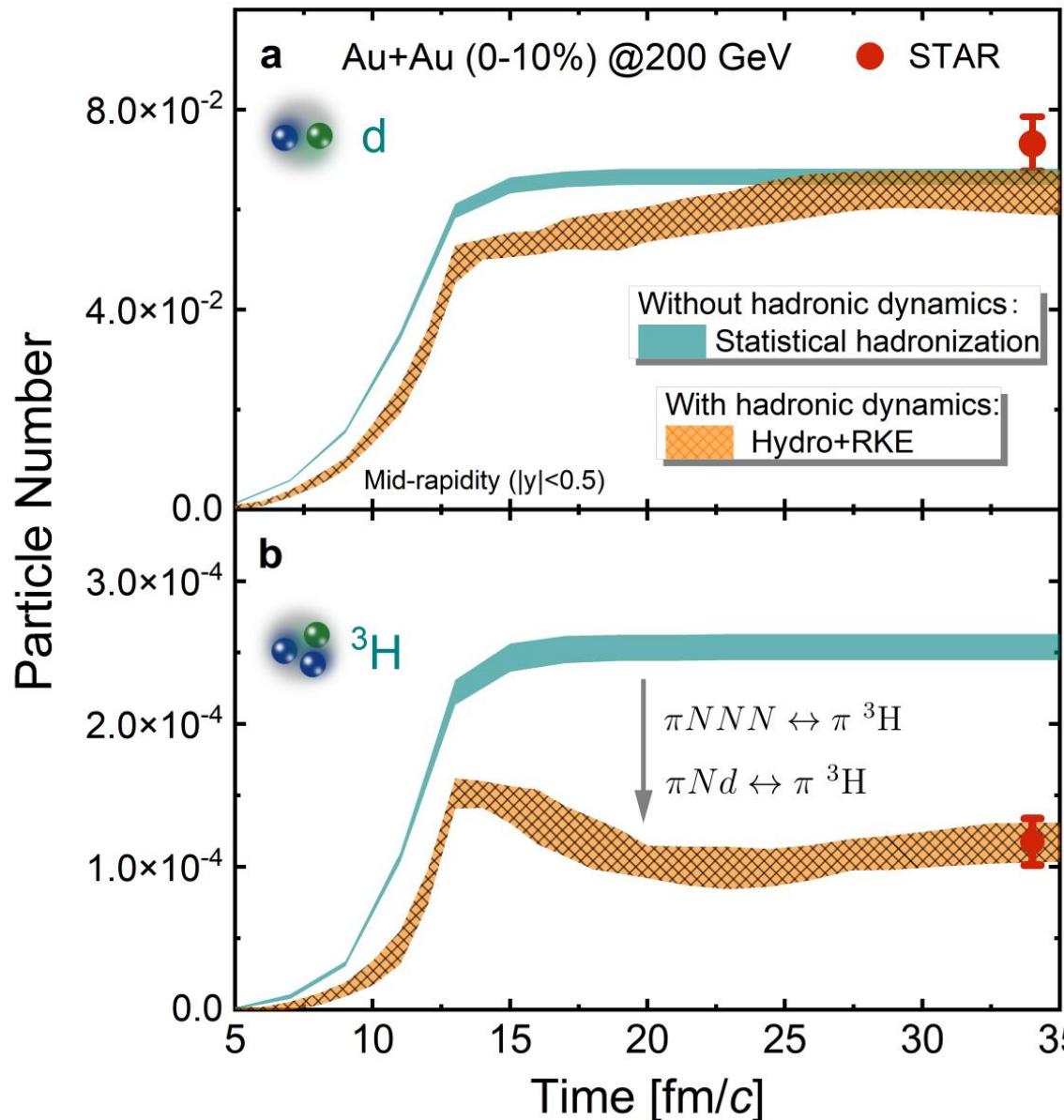


## 2. Hadronic Re-Scattering Effects at RHIC

(7)

K. J. Sun, R. Wang, C. M. Ko, Y. G. Ma, C. Shen, Nat. Commun. 15, 1074 (2024)

Data from STAR, PRL 130, 202301 (2023)



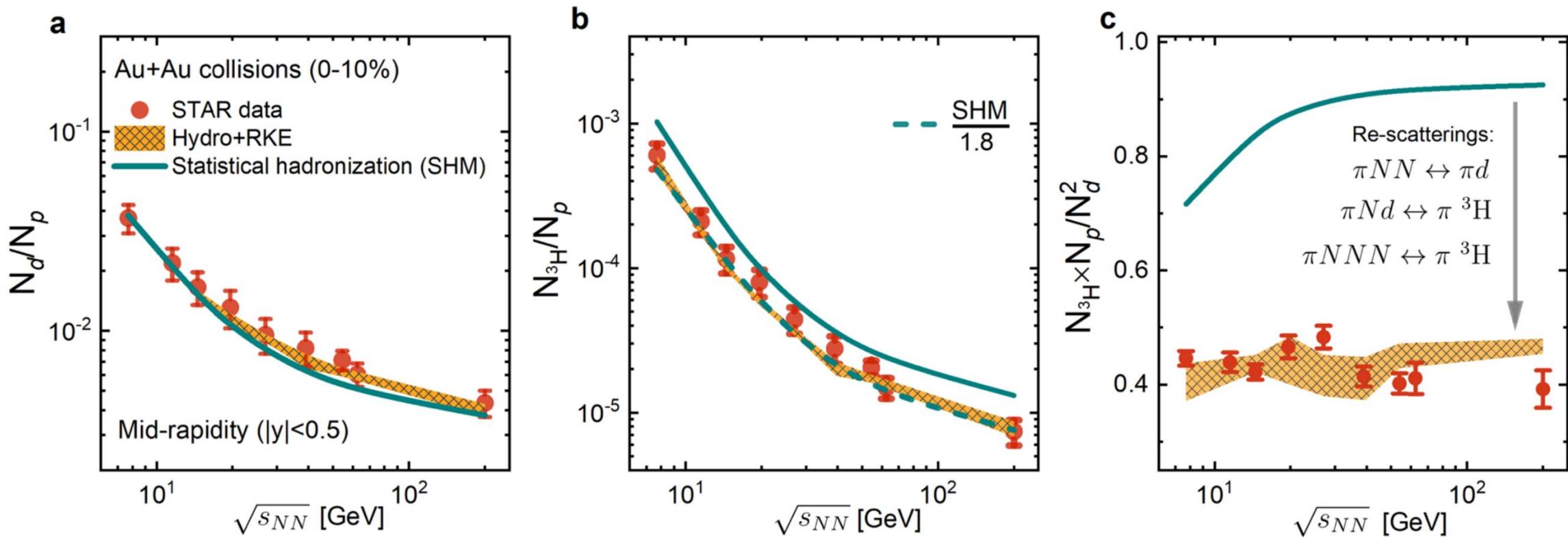
- $A = 2 \quad \pi NN \leftrightarrow \pi d, NNN \leftrightarrow Nd$
- $A = 3 \quad \pi NNN \leftrightarrow \pi t(h), \pi Nd \leftrightarrow \pi t(h), NNNN \leftrightarrow Nt(h), NNd \leftrightarrow Nt(h)$ , and etc.

## 2. Hadronic Re-Scattering Effects at RHIC

(8)

K. J. Sun, R. Wang, C. M. Ko, Y. G. Ma, C. Shen, *Nat. Commun.* 15, 1074 (2024)

Data from STAR, *PRL* 130, 202301 (2023)



Hadronic re-scatterings have small effects on the final deuteron yields (see also D. Oliinychenko et al. *PRC* 99, 044907 (2019)), but they reduce the triton yields by about a factor of 1.8.

## 2. Final-state coalescence

Density Matrix Formulation (**sudden** approximation)

$$N_A = \text{Tr}(\hat{\rho}_s \hat{\rho}_A) = g_c \int d\Gamma \rho_s(\{x_i, p_i\}) \times W_A(\{x_i, p_i\})$$

Two-body coalescence  $a + b \rightarrow c$ :

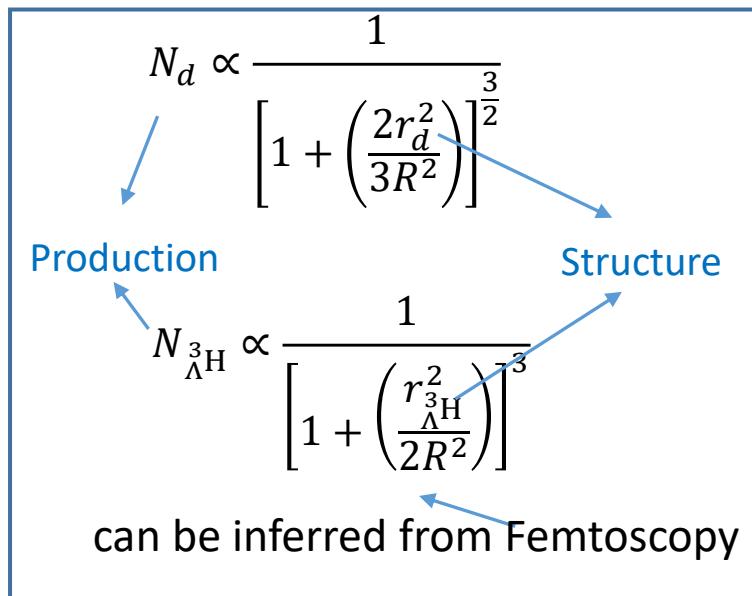
$$N_c = \frac{2J_c+1}{(2J_a+1)(2J_b+1)} \int \frac{dx_a dk_a}{(2\pi)^3} \frac{dx_b dk_b}{(2\pi)^3} f_a(x_a, k_a) f_b(x_b, k_b) W_c(x, k)$$

$$\approx \frac{2J_c+1}{(2J_a+1)(2J_b+1)} \frac{N_a N_b}{(\frac{m_a m_b T}{m_a+m_b} (R_a^2 + R_b^2))^{3/2}} \times \boxed{\frac{1}{(1 + \frac{\sigma^2}{R_a^2 + R_b^2})^{3/2}}}$$

$$f_a = \frac{N_a}{(m_a T R_a^2)^{3/2}} e^{-\frac{k_a^2}{2m_a T} - \frac{x_a^2}{2R_a^2}}$$

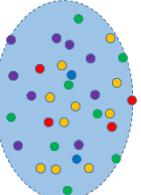
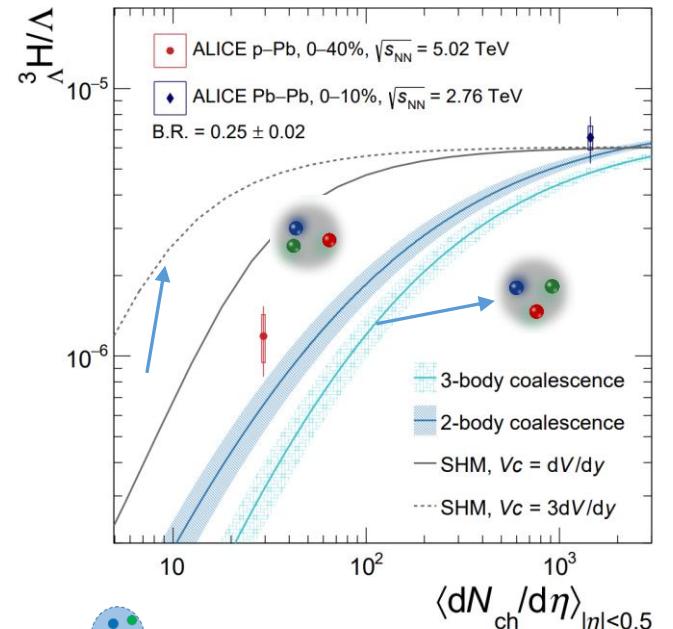
$$N_a = \int \frac{dx_a dk_a}{(2\pi)^3} f_a(x_a, k_a) \quad 1 = \int \frac{dx dk}{(2\pi)^3} W_c(x, k)$$

“Quantum mechanical correction”



ALICE Results (  $\Lambda^3H$  )

Phys. Rev. Lett. 128, 055203(2022)

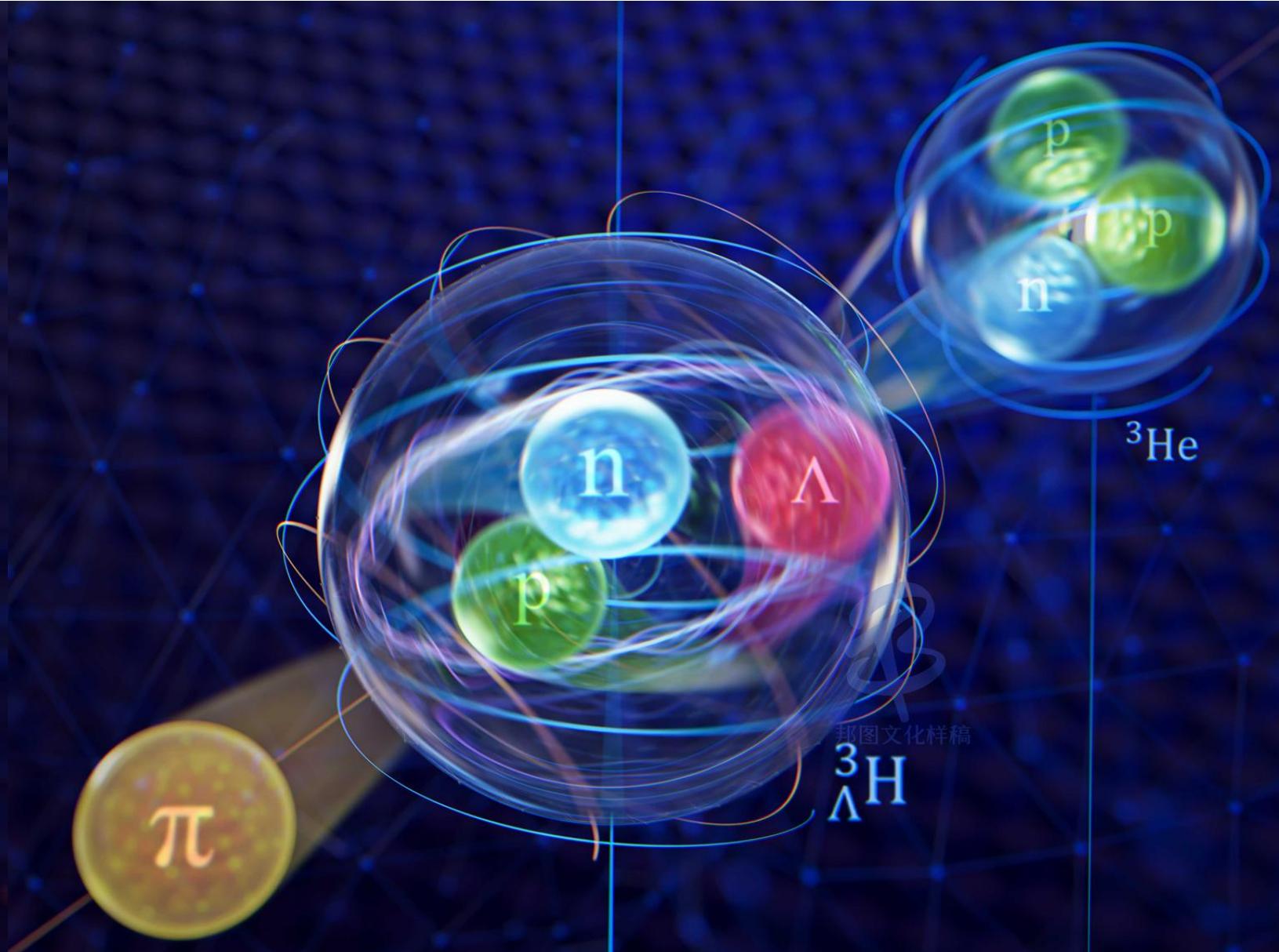


R. Scheibl and U. W. Heinz, PRC59. 1585(1999);

F. Bellini et al., PRC99,054905(2019);

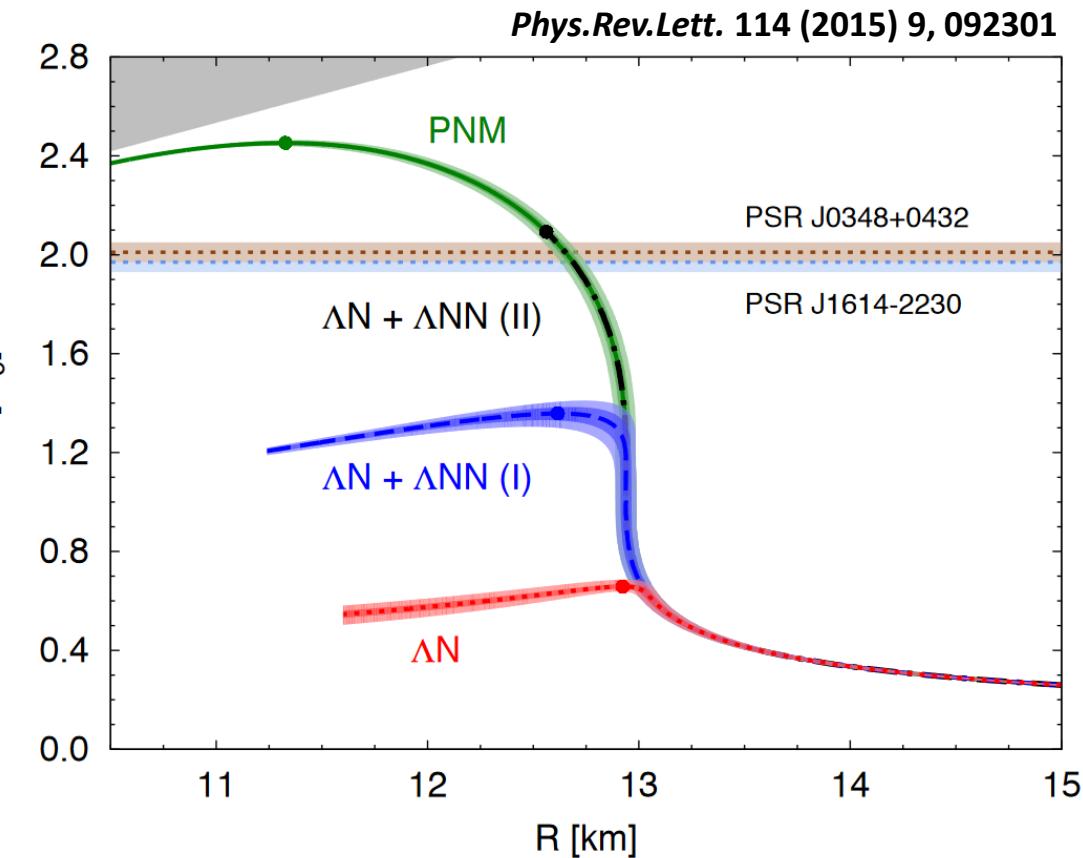
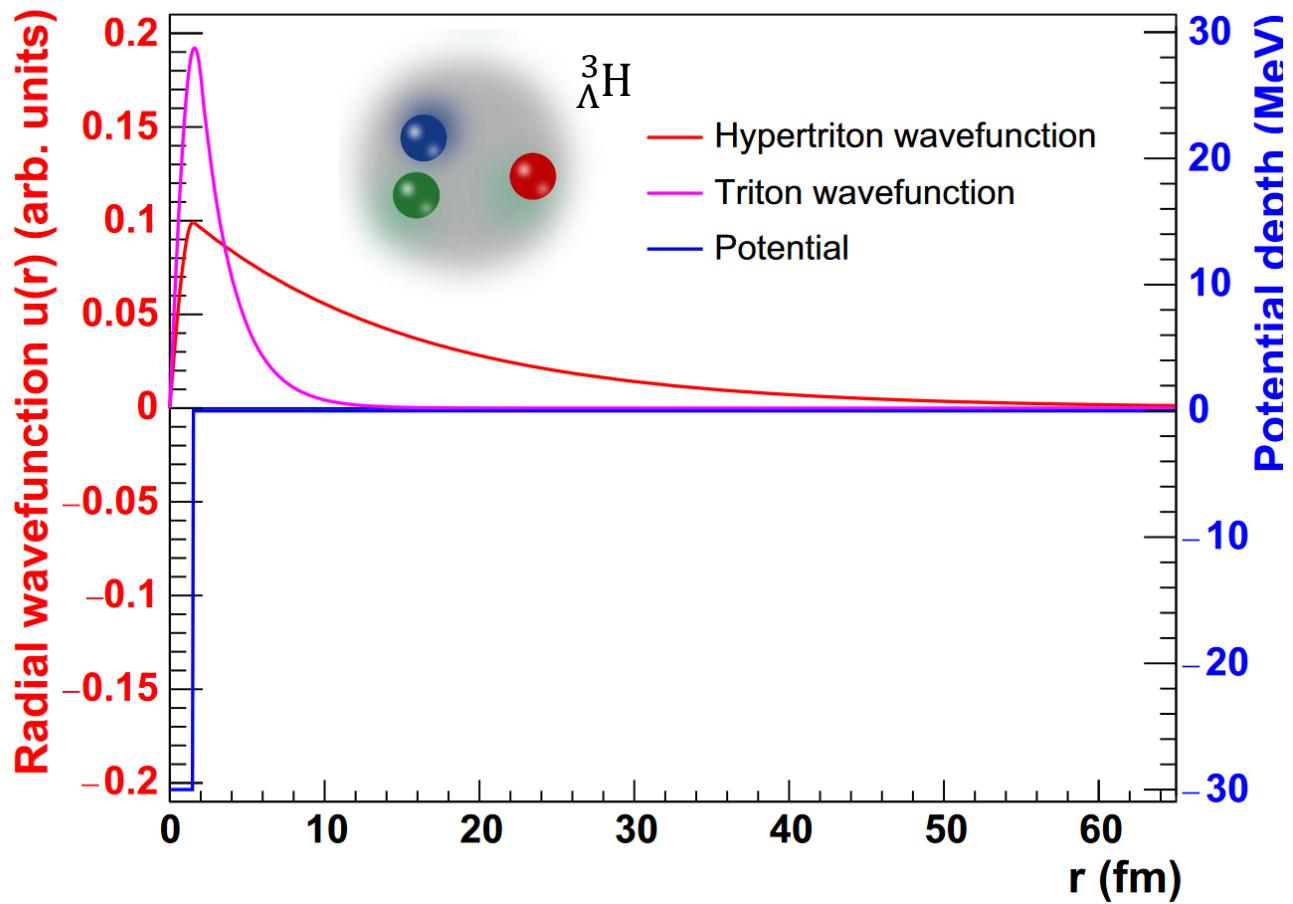
K. J. Sun, C. M. Ko and B. Dönigus, PLB 792, 132 (2019);

# (Anti-)Hypertriton Polarization



### 3. The halo-like nucleus: (anti-)hypertriton

(10)

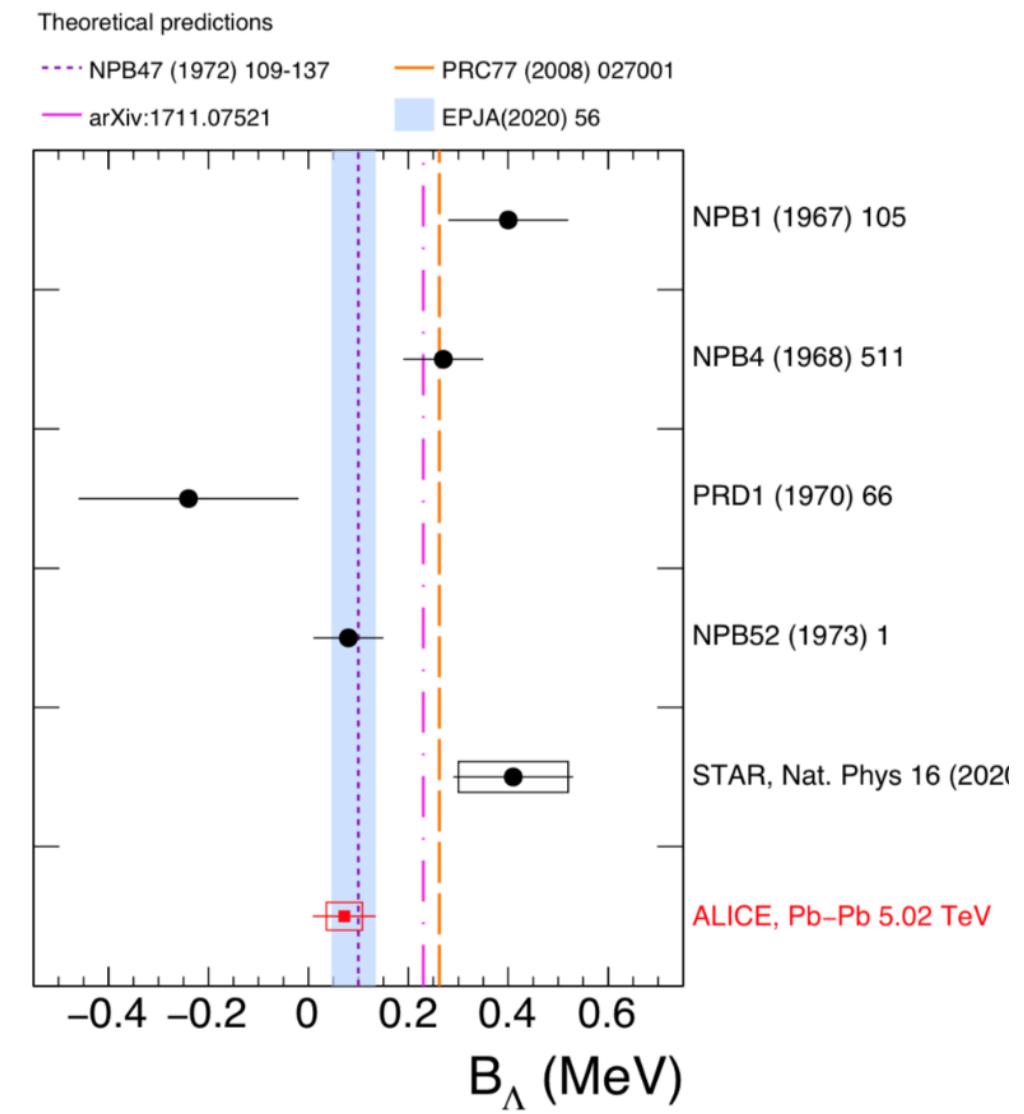
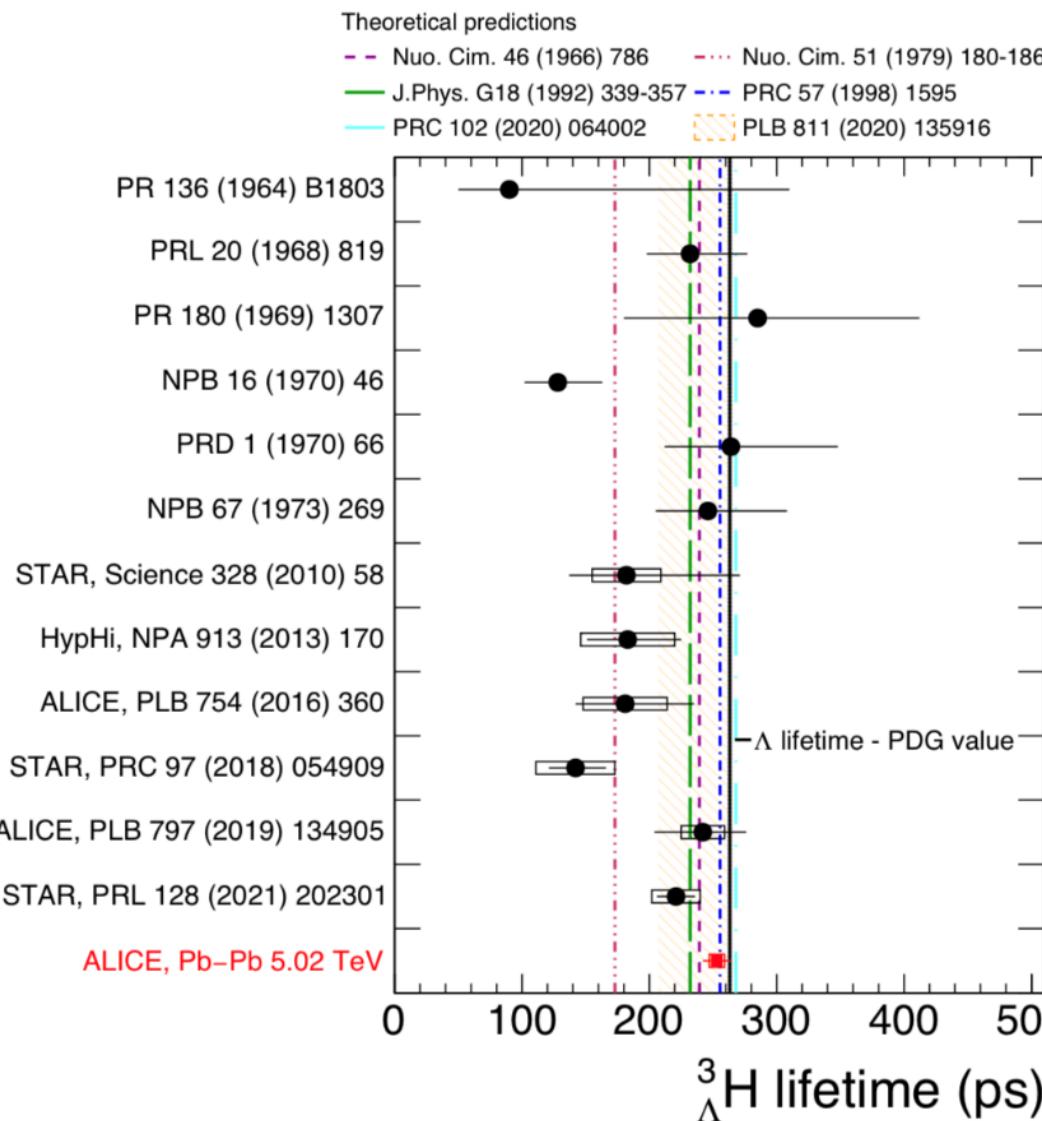


### 3. Binding energy and lifetime

(11)

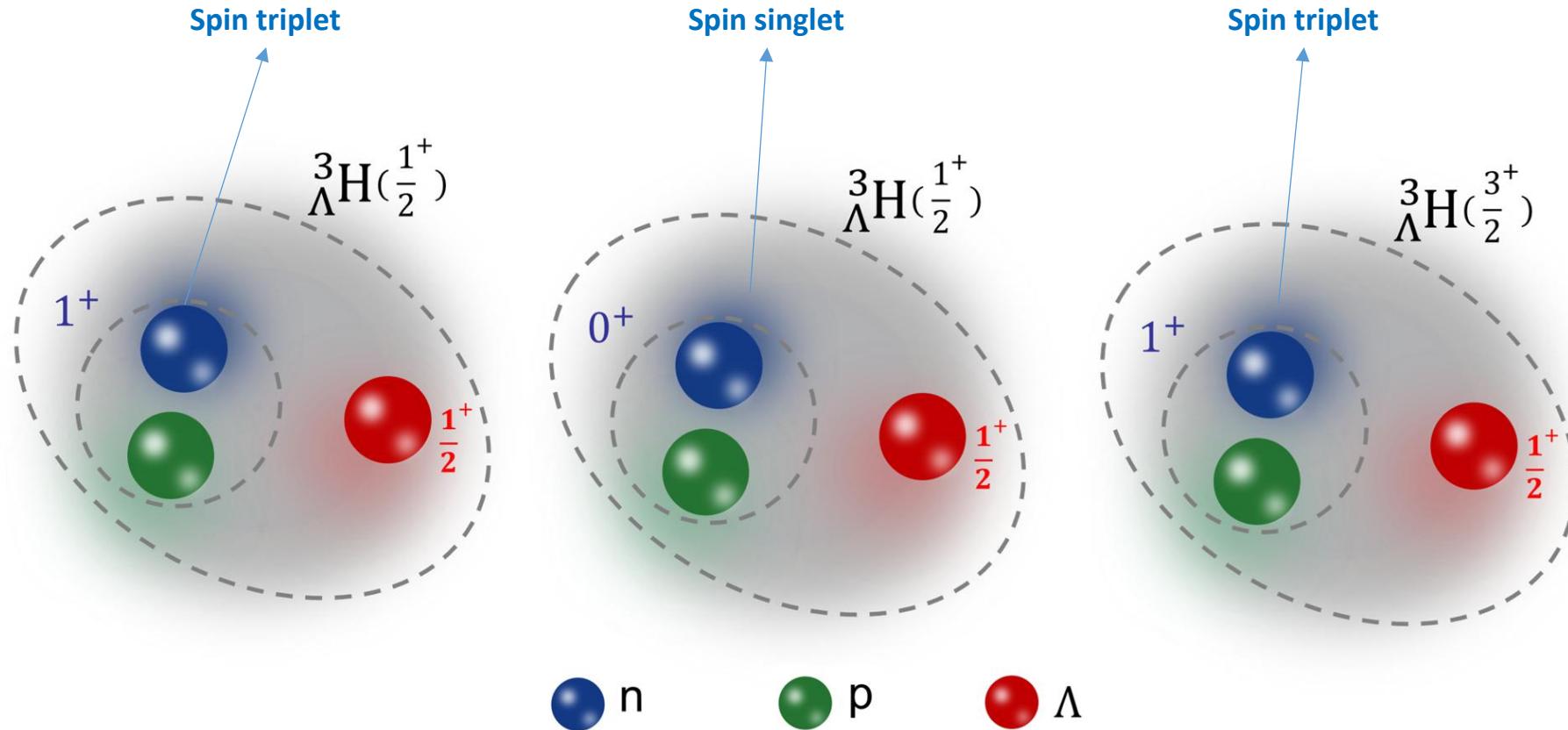
ALICE, PRL 131, 102302 (2023)

Y. G. Ma, Nucl. Sci. Tech. 3497 (2023)



### 3. Spin of (anti-)hypertriton ?

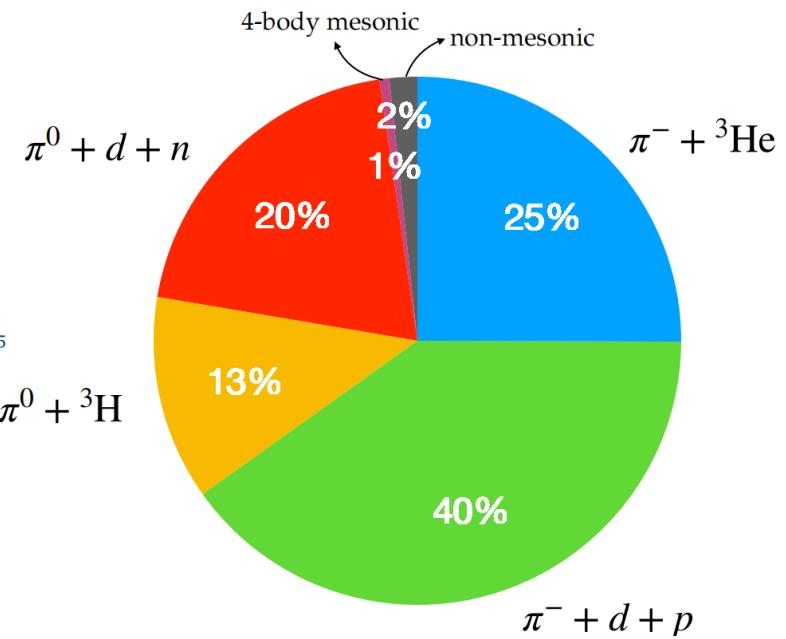
(12)



# 3. Spin of (anti-)hypertriton ?

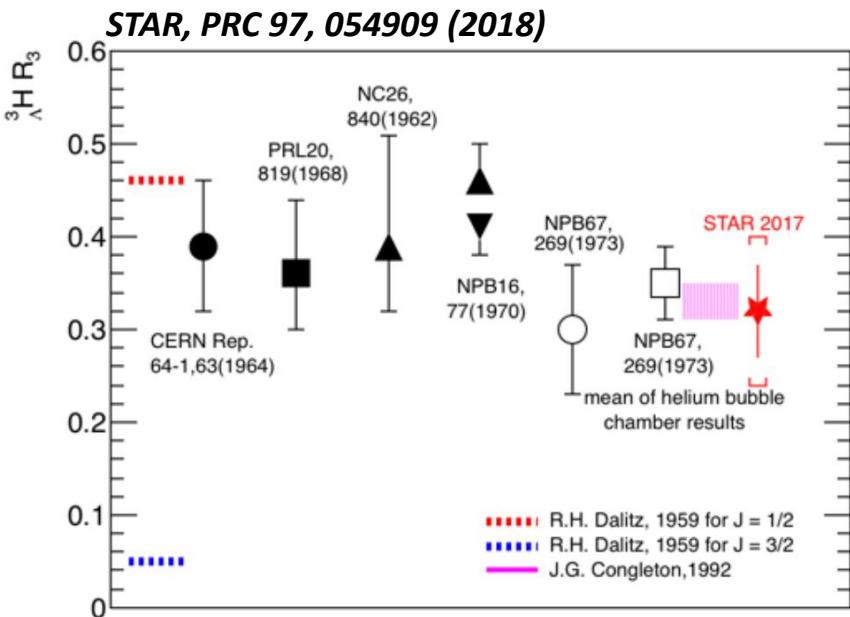
(13)

$$\begin{aligned} {}^3_{\Lambda}\text{H} &\rightarrow \pi^- + {}^3\text{He}, & {}^3_{\Lambda}\text{H} &\rightarrow \pi^0 + {}^3\text{H}, \\ {}^3_{\Lambda}\text{H} &\rightarrow \pi^- + d + p, & {}^3_{\Lambda}\text{H} &\rightarrow \pi^0 + d + n, \\ {}^3_{\Lambda}\text{H} &\rightarrow \pi^- + p + n + p, & {}^3_{\Lambda}\text{H} &\rightarrow \pi^0 + p + n + n. \end{aligned}$$



Relative branching ratio:

$$R_3 = \frac{\text{B.R.}({}^3_{\Lambda}\text{H} \rightarrow {}^3\text{He}\pi^-)}{\text{B.R.}({}^3_{\Lambda}\text{H} \rightarrow {}^3\text{He}\pi^-) + \text{B.R.}({}^3_{\Lambda}\text{H} \rightarrow d\pi^-)}$$



Favors spin 1/2

PHYSICAL REVIEW D 87, 034506 (2013)

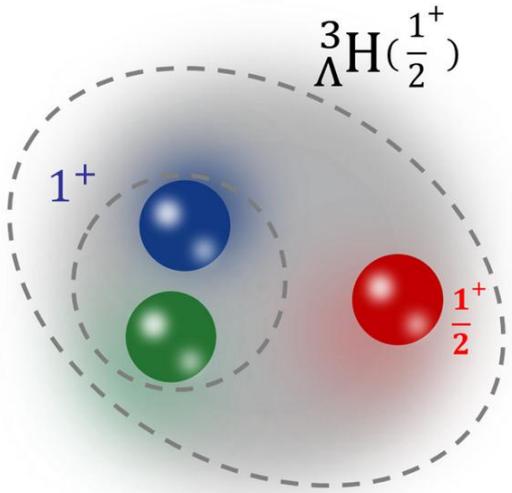
## Light nuclei and hypernuclei from quantum chromodynamics in the limit of SU(3) flavor symmetry

S. R. Beane,<sup>1</sup> E. Chang,<sup>2</sup> S. D. Cohen,<sup>3</sup> W. Detmold,<sup>4,5</sup> H. W. Lin,<sup>3</sup> T. C. Luu,<sup>6</sup> K. Orginos,<sup>4,5</sup> A. Parreño,<sup>2</sup> M. J. Savage,<sup>3</sup> and A. Walker-Loud<sup>7,8</sup>

| Label  | <i>A</i> | <i>s</i> | <i>I</i> | <i>J</i> <sup><i>π</i></sup> | Local<br>SU(3) irreps        | This<br>work         |
|--|----------|----------|----------|------------------------------|------------------------------|----------------------|
| <i>N</i>   | 1        | 0        | 1/2      | 1/2 <sup>+</sup>             | <b>8</b>                     | <b>8</b>             |
| <i>Λ</i>   | 1        | -1       | 0        | 1/2 <sup>+</sup>             | <b>8</b>                     | <b>8</b>             |
| <i>Σ</i>   | 1        | -1       | 1        | 1/2 <sup>+</sup>             | <b>8</b>                     | <b>8</b>             |
| <i>Ξ</i>   | 1        | -2       | 1/2      | 1/2 <sup>+</sup>             | <b>8</b>                     | <b>8</b>             |
| <i>d</i>   | 2        | 0        | 0        | 1 <sup>+</sup>               | <b>10</b>                    | <b>10</b>            |
| <i>nn</i>  | 2        | 0        | 1        | 0 <sup>+</sup>               | <b>27</b>                    | <b>27</b>            |
| <i>nΛ</i>  | 2        | -1       | 1/2      | 0 <sup>+</sup>               | <b>27</b>                    | <b>27</b>            |
| <i>nΛ</i>  | 2        | -1       | 1/2      | 1 <sup>+</sup>               | <b>8<sub>A</sub>, 10</b>     | -                    |
| <i>n\Sigma</i>   | 2        | -1       | 3/2      | 0 <sup>+</sup>               | <b>27</b>                    | <b>27</b>            |
| <i>n\Sigma</i>   | 2        | -1       | 3/2      | 1 <sup>+</sup>               | <b>10</b>                    | <b>10</b>            |
| <i>n\xi</i>  | 2        | -2       | 0        | 1 <sup>+</sup>               | <b>8<sub>A</sub></b>         | <b>8<sub>A</sub></b> |
| <i>n\xi</i>  | 2        | -2       | 1        | 1 <sup>+</sup>               | <b>8<sub>A</sub>, 10, 10</b> | -                    |
| <i>H</i>   | 2        | -2       | 0        | 0 <sup>+</sup>               | <b>1, 27</b>                 | <b>1, 27</b>         |
| ${}^3\text{H}, {}^3\text{H}$                                 | 3        | 0        | 1/2      | 1/2 <sup>+</sup>             | <b>35</b>                    | <b>35</b>            |
| ${}^3\text{H}(1/2^+)$  | 3        | -1       | 0        | 1/2 <sup>+</sup>             | <b>35</b>                    | -                    |
| ${}^3_{\Lambda}\text{H}(3/2^+)$                              | 3        | -1       | 0        | 3/2 <sup>+</sup>             | <b>10</b>                    | <b>10</b>            |
| ${}^3_{\Lambda}\text{He}, {}^3_{\Lambda}\text{H}, nn\Lambda$ | 3        | -1       | 1        | 1/2 <sup>+</sup>             | <b>27, 35</b>                | <b>27, 35</b>        |
| ${}^3_{\Lambda}\text{He}$                                    | 3        | -1       | 1        | 3/2 <sup>+</sup>             | <b>27</b>                    | <b>27</b>            |
| ${}^4\text{He}$  | 4        | 0        | 0        | 0 <sup>+</sup>               | <b>28</b>                    | <b>28</b>            |
| ${}^4_{\Lambda}\text{He}$                                    | 4        | -1       | 1/2      | 0 <sup>+</sup>               | <b>28</b>                    | -                    |
| ${}^4_{\Lambda\Lambda}\text{He}$                             | 4        | -2       | 1        | 0 <sup>+</sup>               | <b>27, 28</b>                | <b>27, 28</b>        |
| $\Lambda\Xi^0 pnn$   | 5        | -3       | 0        | 3/2 <sup>+</sup>             | <b>10 + ⋯</b>                | <b>10</b>            |

Favors spin 3/2

### 3. (Anti-)hypertriton polarization and its spin structure (14)



$$\begin{aligned}
 |\frac{1}{2}, \uparrow\rangle_{\Lambda^3} &= \frac{\sqrt{6}}{3} |\frac{1}{2}, \frac{1}{2}\rangle_n |\frac{1}{2}, \frac{1}{2}\rangle_p |\frac{1}{2}, -\frac{1}{2}\rangle_\Lambda \\
 &- \frac{\sqrt{6}}{6} (|\frac{1}{2}, \frac{1}{2}\rangle_n |\frac{1}{2}, -\frac{1}{2}\rangle_p |\frac{1}{2}, \frac{1}{2}\rangle_\Lambda \\
 &+ |\frac{1}{2}, -\frac{1}{2}\rangle_n |\frac{1}{2}, \frac{1}{2}\rangle_p |\frac{1}{2}, \frac{1}{2}\rangle_\Lambda), \\
 |\frac{1}{2}, \downarrow\rangle_{\Lambda^3} &= -\frac{\sqrt{6}}{3} |\frac{1}{2}, -\frac{1}{2}\rangle_n |\frac{1}{2}, -\frac{1}{2}\rangle_p |\frac{1}{2}, \frac{1}{2}\rangle_\Lambda \\
 &+ \frac{\sqrt{6}}{6} (|\frac{1}{2}, \frac{1}{2}\rangle_n |\frac{1}{2}, -\frac{1}{2}\rangle_p |\frac{1}{2}, -\frac{1}{2}\rangle_\Lambda \\
 &+ |\frac{1}{2}, -\frac{1}{2}\rangle_n |\frac{1}{2}, \frac{1}{2}\rangle_p |\frac{1}{2}, -\frac{1}{2}\rangle_\Lambda).
 \end{aligned}$$

**Coalescence model for hypertriton production (without baryon spin correlation)**

$$E_i \frac{d^3 N_{i,\pm\frac{1}{2}}}{d\mathbf{p}_i^3} = \int_{\Sigma^\mu} d^3 \sigma_\mu p_i^\mu w_{i,\pm\frac{1}{2}}(\mathbf{x}_i, \mathbf{p}_i) \bar{f}_i(\mathbf{x}_i, \mathbf{p}_i)$$

$$w_{i,\pm\frac{1}{2}} = \frac{1}{2} [1 \pm \mathcal{P}_i(\mathbf{x}_i, \mathbf{p}_i)]$$

$$\hat{\rho}_i = \text{diag} \left( \frac{1+\mathcal{P}_i}{2}, \frac{1-\mathcal{P}_i}{2} \right)$$

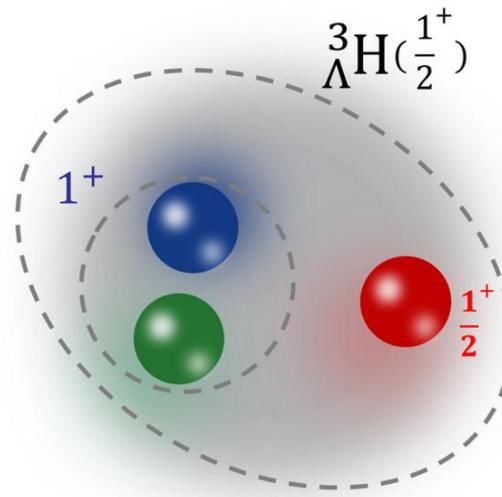
$$\bar{f}_i = \frac{g_i}{(2\pi)^3} \left[ \exp(p_i^\mu u_\mu/T)/\xi_i + 1 \right]^{-1}$$

$$\hat{\rho}_{np\Lambda} = \hat{\rho}_n \otimes \hat{\rho}_p \otimes \hat{\rho}_\Lambda$$

$$\begin{aligned}
 E \frac{d^3 N_{\Lambda^3, \pm\frac{1}{2}}}{d\mathbf{P}^3} &= E \int \prod_{i=n,p,\Lambda} p_i^\mu d^3 \sigma_\mu \frac{d^3 p_i}{E_i} \bar{f}_i(\mathbf{x}_i, \mathbf{p}_i) \\
 &\times \left( \frac{2}{3} w_{n,\pm\frac{1}{2}} w_{p,\pm\frac{1}{2}} w_{\Lambda,\mp\frac{1}{2}} + \frac{1}{6} w_{n,\pm\frac{1}{2}} w_{p,\mp\frac{1}{2}} w_{\Lambda,\pm\frac{1}{2}} \right. \\
 &\quad \left. + \frac{1}{6} w_{n,\mp\frac{1}{2}} w_{p,\pm\frac{1}{2}} w_{\Lambda,\pm\frac{1}{2}} \right) \\
 &\times W_{\Lambda^3 H}(\mathbf{x}_n, \mathbf{x}_p, \mathbf{x}_\Lambda; \mathbf{p}_n, \mathbf{p}_p, \mathbf{p}_\Lambda) \delta(\mathbf{P} - \sum_i \mathbf{p}_i)
 \end{aligned}$$

### 3. (Anti-)hypertriton polarization and its spin structure

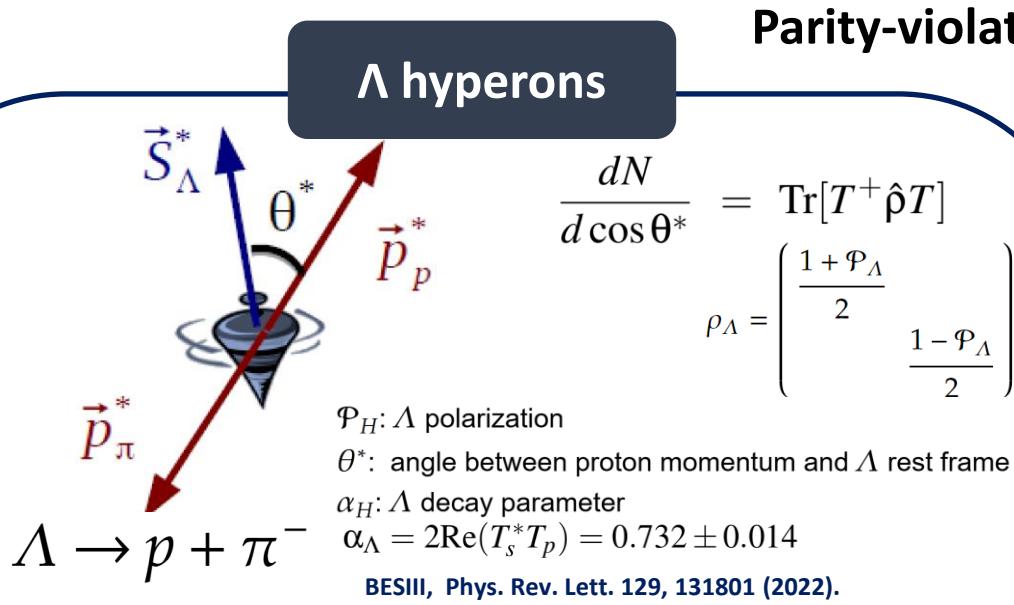
(15)



$$\begin{aligned}\mathcal{P}_{^3\Lambda H} &\approx \frac{\frac{2}{3}\mathcal{P}_n + \frac{2}{3}\mathcal{P}_p - \frac{1}{3}\mathcal{P}_\Lambda - \mathcal{P}_n\mathcal{P}_p\mathcal{P}_\Lambda}{1 - \frac{2}{3}(\mathcal{P}_n + \mathcal{P}_p)\mathcal{P}_\Lambda + \frac{1}{3}\mathcal{P}_n\mathcal{P}_p} \\ &\approx \frac{2}{3}\mathcal{P}_n + \frac{2}{3}\mathcal{P}_p - \frac{1}{3}\mathcal{P}_\Lambda \\ &\approx \mathcal{P}_\Lambda\end{aligned}$$

### 3. (Anti-)hypertriton polarization and its spin structure (16)

K. J. Sun et al., arXiv:2405.12015(2024)



#### The transition matrix

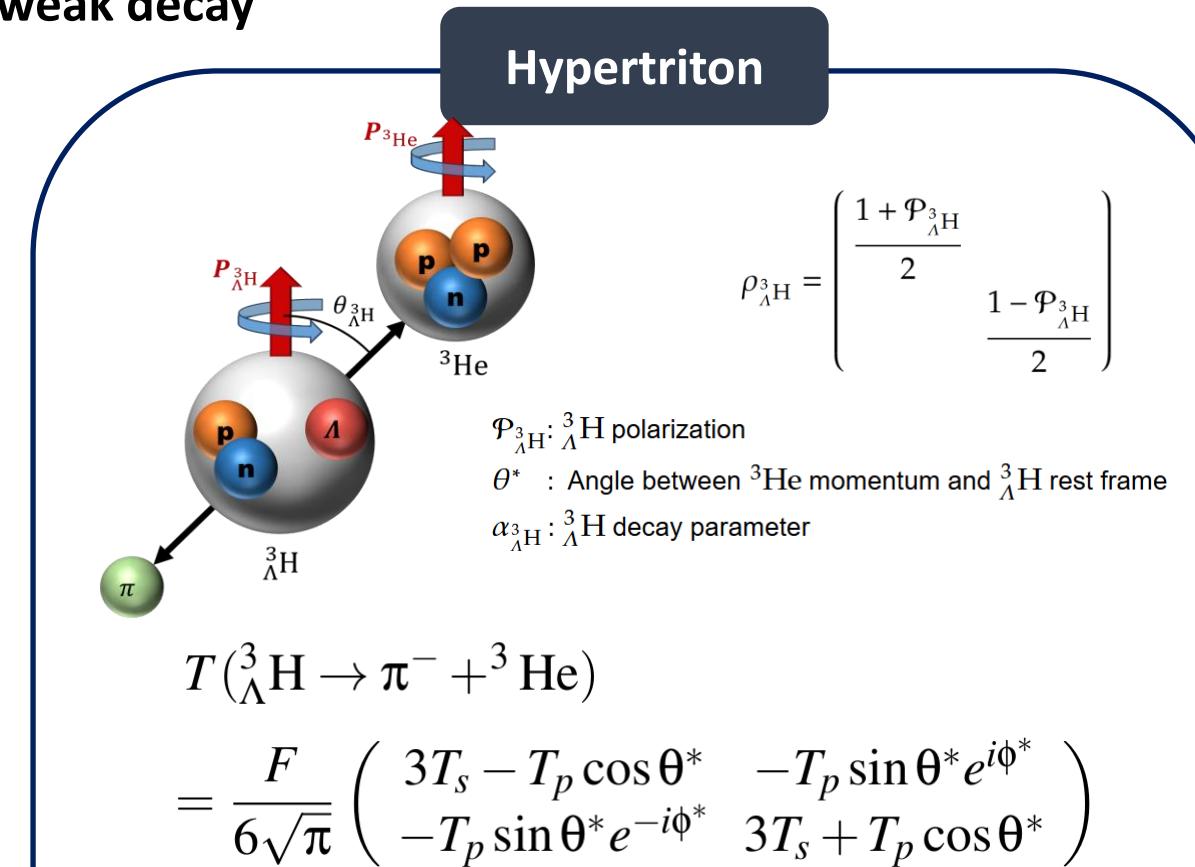
$$T(\Lambda \rightarrow \pi^- + p) = \frac{1}{\sqrt{4\pi}} \begin{pmatrix} T_s + T_p \cos \theta_p^* & T_p \sin \theta_p^* e^{i\phi_p^*} \\ T_p \sin \theta_p^* e^{-i\phi_p^*} & T_s - T_p \cos \theta_p^* \end{pmatrix}$$

#### The angular distribution

$$\frac{dN}{d \cos \theta^*} = \frac{1}{2} (1 + \alpha_H |\mathcal{P}_H| \cos \theta^*)$$

H denotes  $\Lambda$  and  $\bar{\Lambda}$

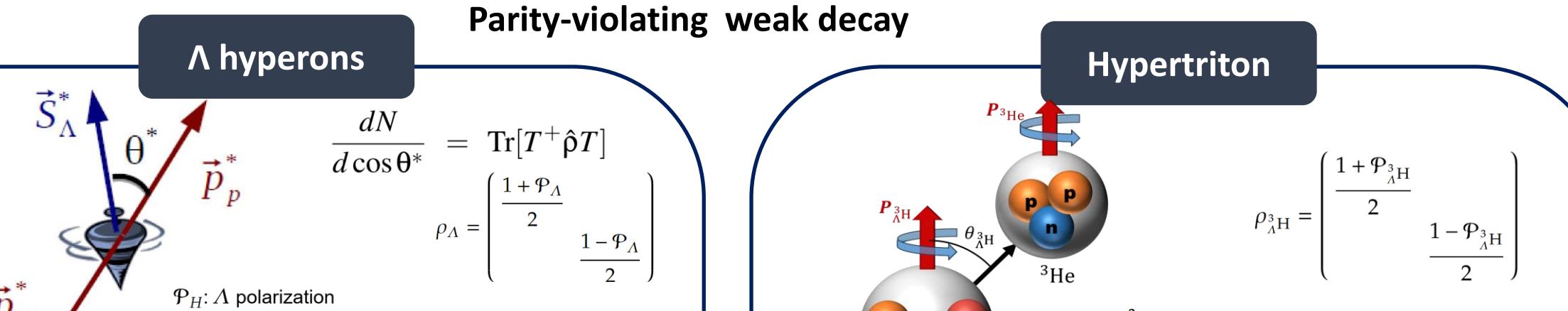
#### Parity-violating weak decay



$$\frac{dN}{d \cos \theta^*} = \frac{1}{2} (1 + \alpha_{^3\text{H}_\Lambda} \mathcal{P}_{^3\text{H}_\Lambda} \cos \theta^*)$$

### 3. (Anti-)hypertriton polarization and its spin structure (16)

K. J. Sun et al., arXiv:2405.12015(2024)



#### The transition matrix

$$T(\Lambda \rightarrow \pi^- + p) = \frac{1}{\sqrt{4\pi}} \begin{pmatrix} T_s + T_p \cos\theta_p^* & \\ T_p \sin\theta_p^* e^{-i\phi_p^*} & T_s - T_p \cos\theta_p^* \\ & -T_p \sin\theta_p^* e^{i\phi_p^*} \end{pmatrix}$$

#### The angular distribution

$$\frac{dN}{d\cos\theta^*} = \frac{1}{2} (1 + \alpha_H |\mathcal{P}_H| \cos\theta^*)$$

$H$  denotes  $\Lambda$  and  $\bar{\Lambda}$

**Sign flip!**

$$\alpha_{^3\text{H}} \approx -\frac{1}{3T_s^2 + \frac{1}{3}T_p^2} \alpha_\Lambda$$

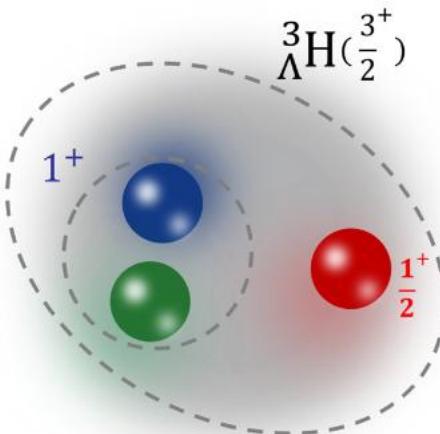
$$\approx -\frac{1}{2.58} \alpha_\Lambda$$

$$\pi^- + ^3\text{He})$$

$$3T_s - T_p \cos\theta^* \quad -T_p \sin\theta^* e^{i\phi^*} \\ -T_p \sin\theta^* e^{-i\phi^*} \quad 3T_s + T_p \cos\theta^*$$

$$\frac{dN}{d\cos\theta^*} = \frac{1}{2} (1 + \alpha_{^3\text{H}} \mathcal{P}_{^3\text{H}} \cos\theta^*)$$

### 3. (Anti-)hypertriton polarization and its spin structure (17)

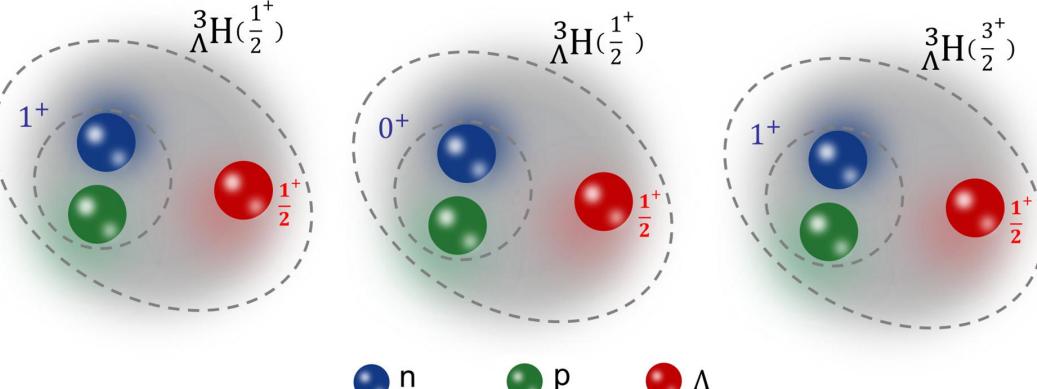


$$\hat{\rho}_{^3\Lambda H} \approx \text{diag} \left[ \frac{(1 + \mathcal{P}_\Lambda)^3}{4(1 + \mathcal{P}_\Lambda^2)}, \frac{(1 - \mathcal{P}_\Lambda)(1 + \mathcal{P}_\Lambda)^2}{4(1 + \mathcal{P}_\Lambda^2)}, \frac{(1 - \mathcal{P}_\Lambda)^2(1 + \mathcal{P}_\Lambda)}{4(1 + \mathcal{P}_\Lambda^2)}, \frac{(1 - \mathcal{P}_\Lambda)^3}{4(1 + \mathcal{P}_\Lambda^2)} \right]$$

$$T(^3_\Lambda H \rightarrow \pi^- + ^3 He) = \frac{FT_p}{\sqrt{6\pi}} \begin{pmatrix} e^{i\phi^*} \sin \theta^* & 0 \\ -\frac{2}{\sqrt{3}} \cos \theta^* & \frac{e^{i\phi^*} \sin \theta^*}{\sqrt{3}} \\ -\frac{e^{-i\phi^*} \sin \theta^*}{\sqrt{3}} & -\frac{2}{\sqrt{3}} \cos \theta^* \\ 0 & -e^{-i\phi^*} \sin \theta^* \end{pmatrix}$$

$$\frac{dN}{d\cos \theta^*} = \frac{1}{2} \left[ 1 + \left( \hat{\rho}_{\frac{1}{2}, \frac{1}{2}} + \hat{\rho}_{-\frac{1}{2}, -\frac{1}{2}} - \frac{1}{2} \right) (3 \cos^2 \theta^* - 1) \right]$$

$$\hat{\rho}_{\frac{1}{2}, \frac{1}{2}} + \hat{\rho}_{-\frac{1}{2}, -\frac{1}{2}} - \frac{1}{2} \approx -\frac{\mathcal{P}_\Lambda^2}{1 + \mathcal{P}_\Lambda^2} \approx -\mathcal{P}_\Lambda^2$$



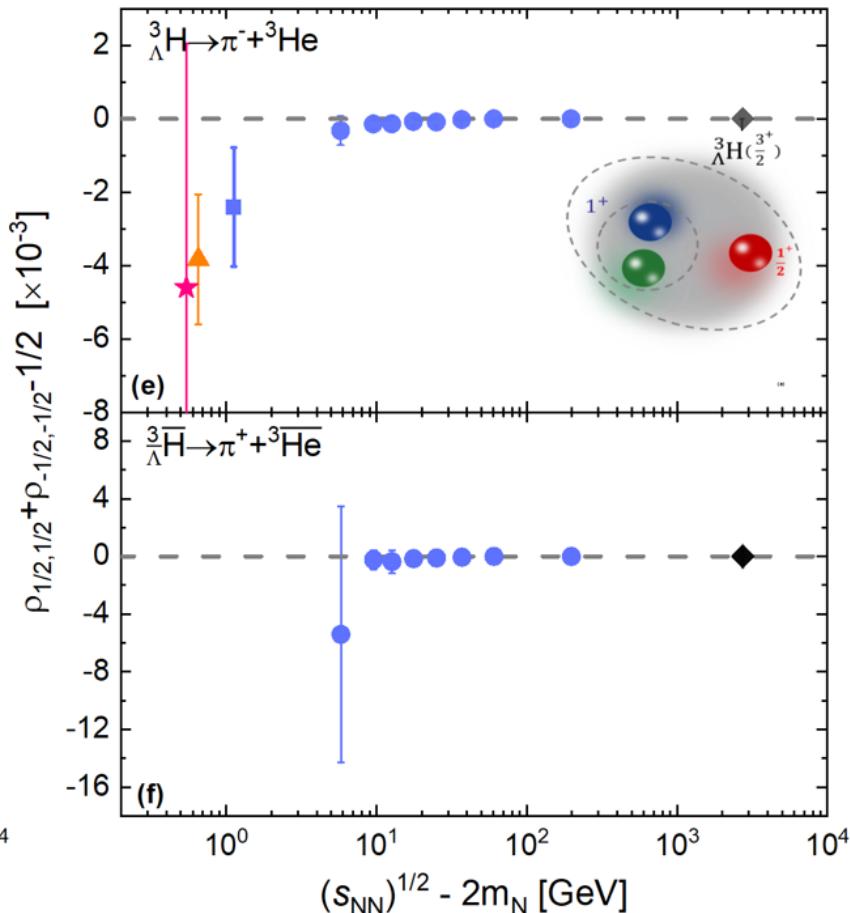
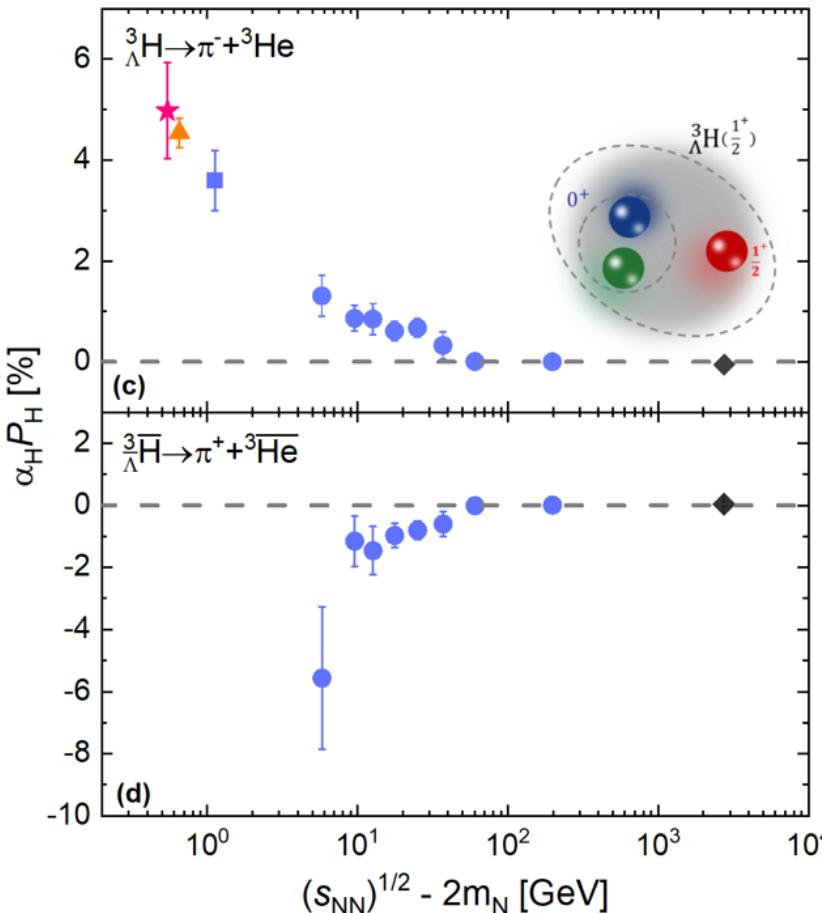
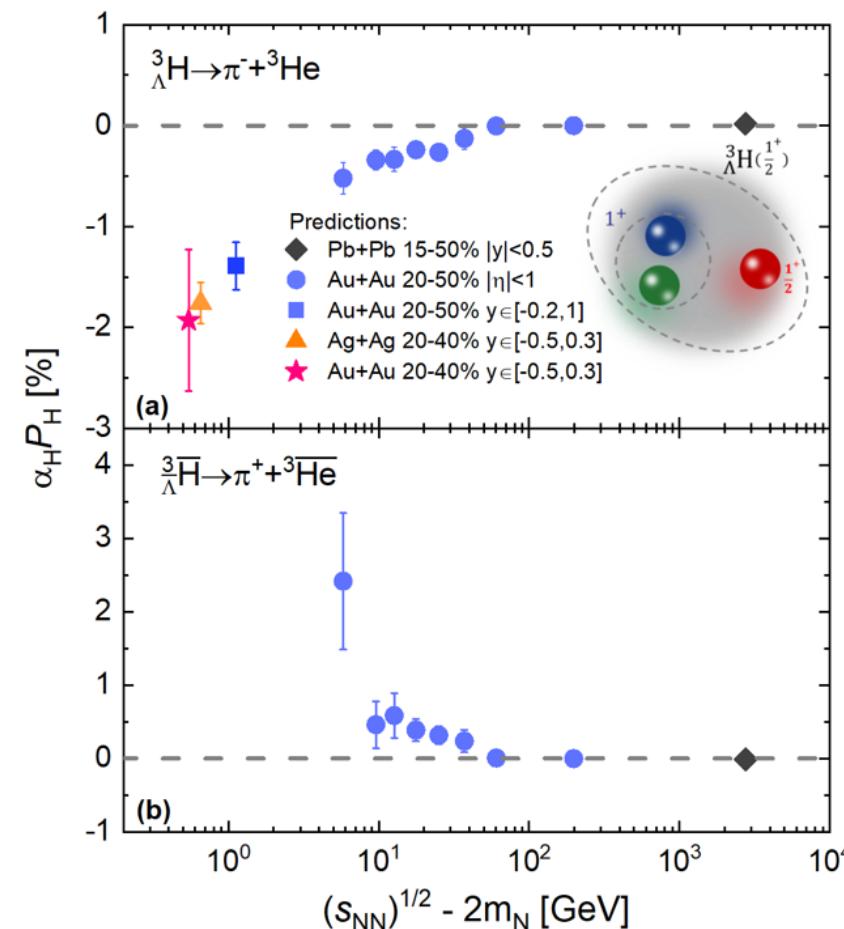
| $J^P$           | structure                                      | decay mode   | $\frac{dN}{d\cos \theta^*}$   |
|-----------------|--|--|---|
| $\frac{1}{2}^+$ | $\Lambda(\frac{1}{2}^+) - np(1^+)$             | ${}^3_\Lambda H \rightarrow \pi^- + {}^3 He$                     | $\frac{1}{2}(1 - \frac{1}{2.58}\alpha_\Lambda \mathcal{P}_\Lambda \cos \theta^*)$                 |
| $\frac{1}{2}^+$ | $\Lambda(\frac{1}{2}^+) - np(0^+)$             | ${}^3_\Lambda H \rightarrow \pi^- + {}^3 He$                     | $\frac{1}{2}(1 + \alpha_\Lambda \mathcal{P}_\Lambda \cos \theta^*)$                               |
| $\frac{3}{2}^+$ | $\Lambda(\frac{1}{2}^+) - np(1^+)$             | ${}^3_\Lambda H \rightarrow \pi^- + {}^3 He$                     | $\frac{1}{2}(1 - \mathcal{P}_\Lambda^2(3 \cos^2 \theta^* - 1))$                                   |
| $\frac{1}{2}^-$ | $\bar{\Lambda}(\frac{1}{2}^-) - \bar{n}p(1^-)$ | ${}^3_{\bar{\Lambda}} \bar{H} \rightarrow \pi^+ + {}^3 \bar{He}$ | $\frac{1}{2}(1 - \frac{1}{2.58}\alpha_{\bar{\Lambda}} \mathcal{P}_{\bar{\Lambda}} \cos \theta^*)$ |
| $\frac{1}{2}^-$ | $\bar{\Lambda}(\frac{1}{2}^-) - \bar{n}p(0^-)$ | ${}^3_{\bar{\Lambda}} \bar{H} \rightarrow \pi^+ + {}^3 \bar{He}$ | $\frac{1}{2}(1 + \alpha_{\bar{\Lambda}} \mathcal{P}_{\bar{\Lambda}} \cos \theta^*)$               |
| $\frac{3}{2}^-$ | $\bar{\Lambda}(\frac{1}{2}^-) - \bar{n}p(1^-)$ | ${}^3_{\bar{\Lambda}} \bar{H} \rightarrow \pi^+ + {}^3 \bar{He}$ | $\frac{1}{2}(1 - \mathcal{P}_{\bar{\Lambda}}^2(3 \cos^2 \theta^* - 1))$                           |

### 3. (Anti-)hypertriton polarization and its spin structure (18)

The measurement of hypertriton polarization provides a novel method to uniquely determine its internal spin structure

$$\alpha_{\Lambda^3 H} \approx -\frac{1}{2.58} \alpha_\Lambda$$

$$\alpha_{\Lambda^3 H} \approx \alpha_\Lambda$$



## 4. Effects of baryon spin correlation (19)

$$\begin{aligned}
\hat{\rho}_{np\Lambda} &= \hat{\rho}_n \otimes \hat{\rho}_p \otimes \hat{\rho}_\Lambda + \frac{1}{2^2} (c_{np}^{\alpha\beta} \hat{\sigma}_{n,\alpha} \otimes \hat{\sigma}_{p,\beta} \otimes \hat{\rho}_\Lambda \\
&\quad + c_{p\Lambda}^{\alpha\beta} \hat{\sigma}_{p,\alpha} \otimes \hat{\sigma}_{\Lambda,\beta} \otimes \hat{\rho}_n + c_{n\Lambda}^{\alpha\beta} \hat{\sigma}_{n,\alpha} \otimes \hat{\sigma}_{\Lambda,\beta} \otimes \hat{\rho}_p) \\
&\quad + \frac{1}{2^3} c_{np\Lambda}^{\alpha\beta\gamma} \hat{\sigma}_{n,\alpha} \otimes \hat{\sigma}_{p,\beta} \otimes \hat{\sigma}_{\Lambda,\gamma}, \\
P_{^3H} &\approx \frac{\frac{2}{3}\langle P_n \rangle + \frac{2}{3}\langle P_p \rangle - \frac{1}{3}\langle P_\Lambda \rangle - \langle P_n P_p P_\Lambda \rangle + C_-}{1 - \frac{2}{3}(\langle (P_n + P_p) P_\Lambda \rangle) + \frac{1}{3}\langle P_n P_p \rangle + C_+} \\
C_- &= -\frac{1}{4}(\langle c_{np}^{zz} P_\Lambda \rangle + \langle c_{p\Lambda}^{zz} P_n \rangle + \langle c_{n\Lambda}^{zz} P_p \rangle) - \frac{1}{4}\langle c_{np\Lambda}^{zzz} \rangle, \quad \text{'genuine' correlation terms} \\
C_+ &= \frac{1}{12}(\langle c_{np}^{zz} \rangle - 2\langle c_{p\Lambda}^{zz} \rangle - 2\langle c_{n\Lambda}^{zz} \rangle).
\end{aligned}$$

### Induced correlations

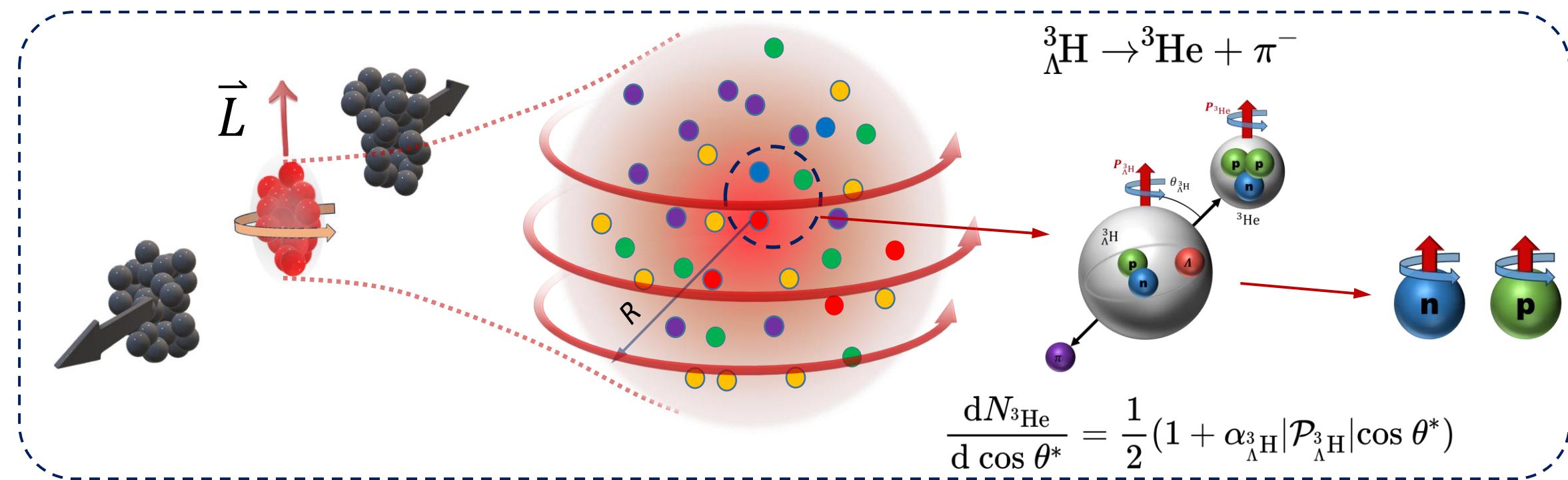
We can express the polarization of a particle as  $\mathcal{P} = \langle \mathcal{P} \rangle + \delta\mathcal{P}$  with  $\delta\mathcal{P}$  denoting its space and momentum dependent fluctuations, which leads to the relations  $\langle P_n P_p \rangle = \langle P_n \rangle \langle P_p \rangle + \langle \delta P_n \delta P_p \rangle$  and  $\langle P_n P_p P_\Lambda \rangle = \langle P_n \rangle \langle P_p \rangle \langle P_\Lambda \rangle + \langle \delta P_n \delta P_p \rangle \langle P_\Lambda \rangle + \langle \delta P_n \delta P_\Lambda \rangle \langle P_p \rangle + \langle \delta P_p \delta P_\Lambda \rangle \langle P_n \rangle + \langle \delta P_n \delta P_p \delta P_\Lambda \rangle$ . Assuming again  $\langle P_n \rangle \approx \langle P_p \rangle \approx \langle P_\Lambda \rangle$  and neglecting the three-body correlation, we then have

$$P_{^3H} \approx (1 - \langle \delta P_n \delta P_p \rangle - \langle \delta P_p \delta P_\Lambda \rangle - \langle \delta P_n \delta P_\Lambda \rangle) \langle P_\Lambda \rangle.$$

This result suggests that it is possible to extract the information on the spin-spin correlations among nucleons and  $\Lambda$  hyperons from the measurement of hypertriton polarization in heavy-ion collisions, although it is non-trivial in practice.

## 4. A possible way to probe polarization of nucleons

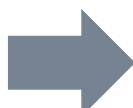
(20)



### Production of hypertriton

$$N_A = \int d\Gamma g_c p_s(x_i, p_i) \times W_c(x_i, p_i)$$

$$\mathcal{P}_{^3\text{H}} = \frac{N_\uparrow - N_\downarrow}{N_\uparrow + N_\downarrow}$$



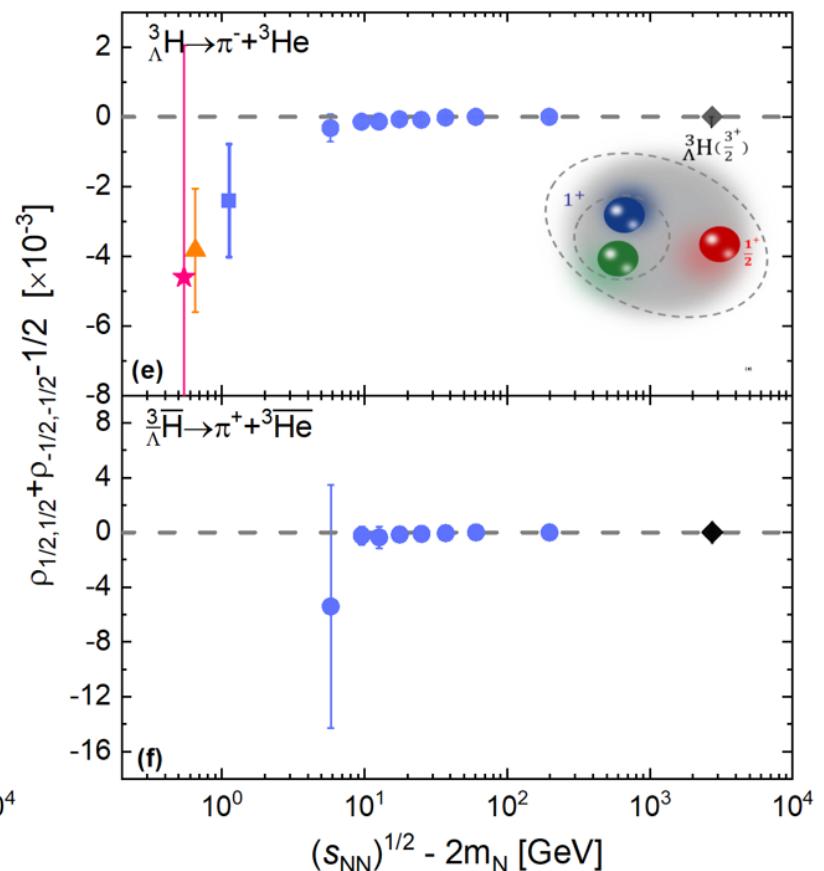
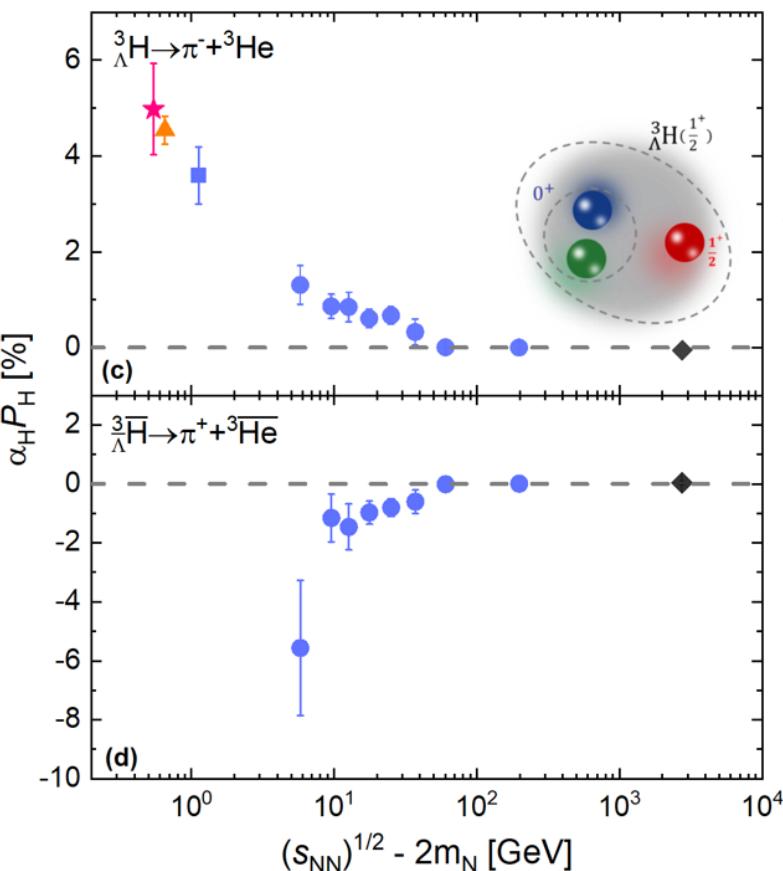
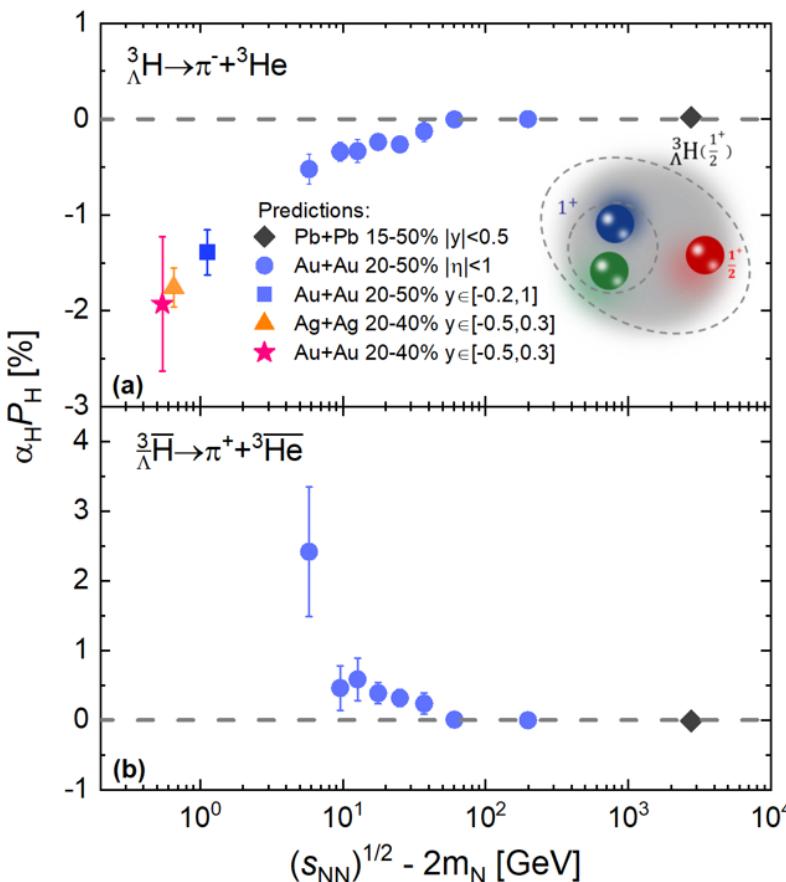
### Polarization of nucleons

$$\mathcal{P}_p + \mathcal{P}_n \approx \frac{1}{2} (3\mathcal{P}_{^3\text{H}} + \mathcal{P}_\Lambda)$$

# Summary and outlook

(21)

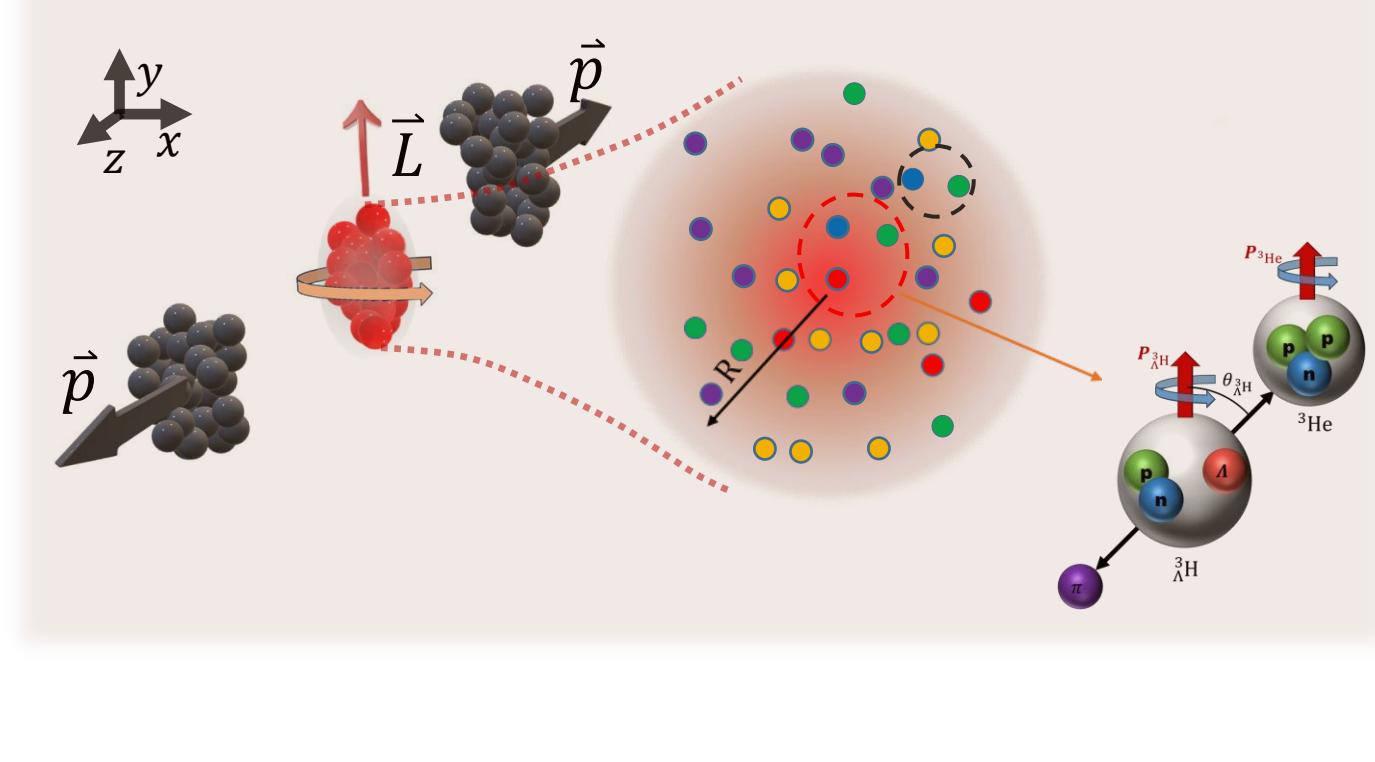
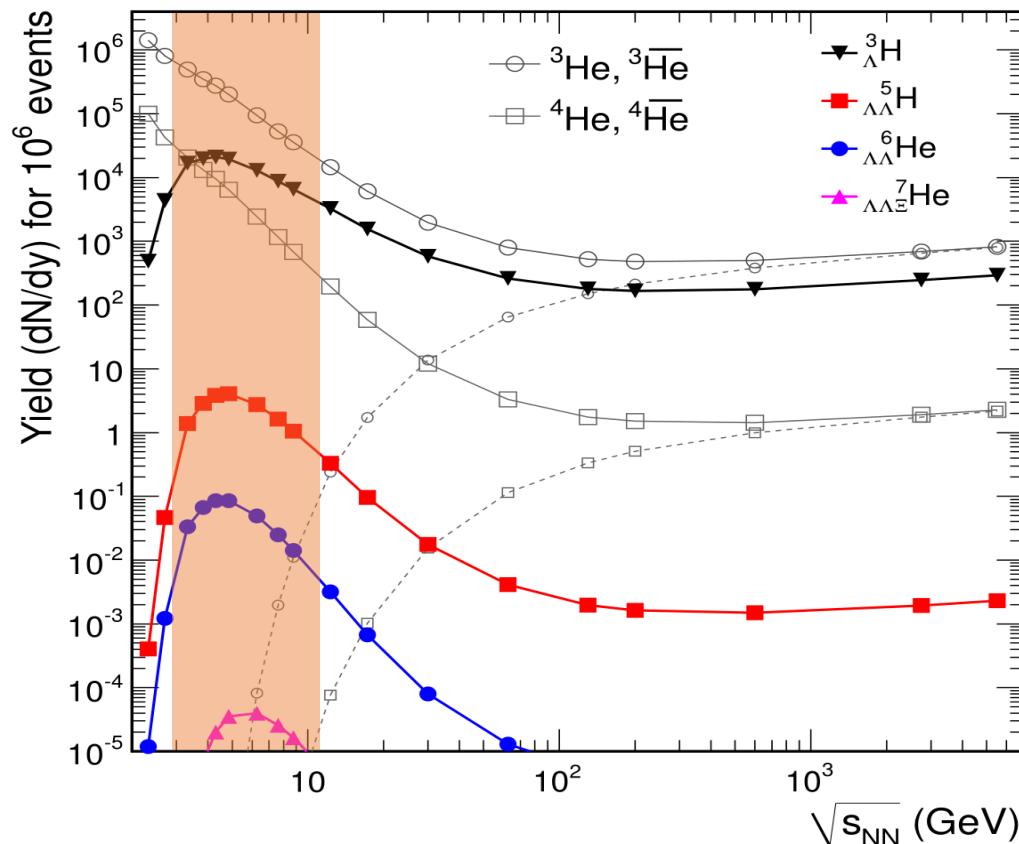
1. (Anti-)hypertriton is globally polarized in non-central heavy-ion collisions.
2. (Anti-)hypertriton polarization and its decay pattern provide a novel method to uniquely determine the spin structure of its wavefunction.



# Summary and outlook

(22)

A. Andronic et al., Phys. Lett. B 697, 203-207 (2011)



- FAIR/CBM (2.4-4.9 GeV)
- HIAF/CEE (2.1-4.5 GeV)
- NICA/MPD (4-11 GeV)

A novel tool to study the evolution of strongly-interacting matter at high-baryon density region

# Backup

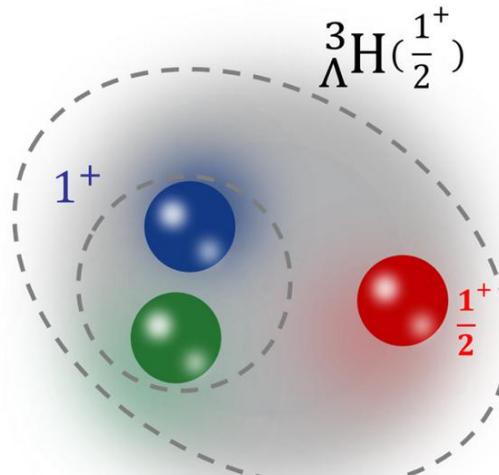
### 3. (Anti-)hypertriton polarization and its spin structure (8)

Parity-violating weak decay:

$$T(\Lambda \rightarrow \pi^- + p) = \frac{1}{\sqrt{4\pi}} \begin{pmatrix} T_s + T_p \cos \theta_p^* & T_p \sin \theta_p^* e^{i\phi_p^*} \\ T_p \sin \theta_p^* e^{-i\phi_p^*} & T_s - T_p \cos \theta_p^* \end{pmatrix}$$

$$T(\Lambda^3 \text{H} \rightarrow \pi^- + ^3 \text{He})$$

$$= \frac{F}{6\sqrt{\pi}} \begin{pmatrix} 3T_s - T_p \cos \theta^* & -T_p \sin \theta^* e^{i\phi^*} \\ -T_p \sin \theta^* e^{-i\phi^*} & 3T_s + T_p \cos \theta^* \end{pmatrix}$$



**Sign flip !**

The normalized angular distribution of the  ${}^3 \text{He}$  in the decay  $\Lambda^3 \text{H} \rightarrow \pi^- + {}^3 \text{He}$  is given by

$$\frac{dN}{d \cos \theta^*} = \text{Tr}[T^+ \hat{\rho} T] = \frac{1}{2}(1 + \alpha_{\Lambda^3 \text{H}} \mathcal{P}_{\Lambda^3 \text{H}} \cos \theta^*), \quad (7)$$

in terms of the hypertriton decay parameter  $\alpha_{\Lambda^3 \text{H}} \approx -\frac{1}{3T_s^2 + \frac{1}{3}T_p^2} \alpha_\Lambda \approx -\frac{1}{2.58} \alpha_\Lambda$ . The angular distribution of  ${}^3 \text{He}$  in the decay  $\Lambda^3 \text{H} \rightarrow \pi^- + {}^3 \text{He}$  can thus be further expressed as

$$\frac{dN}{d \cos \theta^*} \approx \frac{1}{2}(1 - \frac{1}{2.58} \alpha_\Lambda \mathcal{P}_\Lambda \cos \theta^*). \quad (8)$$

Compared to the angular distribution of the proton in the  $\Lambda$  decay, which has the form

$$\frac{dN}{d \cos \theta_p^*} = \frac{1}{2}(1 + \alpha_\Lambda \mathcal{P}_\Lambda \cos \theta_p^*), \quad (9)$$

the  ${}^3 \text{He}$  in  $\Lambda^3 \text{H}$  decay has an opposite sign in its angular dependence.

# 4. Effects of baryon spin correlation (17)

Z. T. Liang, Chirality 2023

$$\left| \rho_{00}^V - \frac{1}{3} \right| \gg P_\Lambda^2 \sim P_q^2$$

$$\rho_{00}^V - \frac{1}{3} \sim \langle P_q P_{\bar{q}} \rangle$$

The STAR data show that:  $\langle P_q P_{\bar{q}} \rangle \neq \langle P_q \rangle \langle P_{\bar{q}} \rangle$      $\langle P_q P_{\bar{q}} \rangle \gg \langle P_q \rangle \langle P_{\bar{q}} \rangle$

By studying  $P_H$ , we study the average of quark polarization  $P_q$ ;  
by studying  $\rho_{00}^V$ , we study the correlation between  $P_q$  and  $P_{\bar{q}}$ .

How to separate long range or local correlations

$$\rho_{10}^V = \frac{P_{qz}(1 + P_{\bar{q}y}) + (1 + P_{qy})P_{\bar{q}z} - iP_{qx}(1 + P_{\bar{q}y}) - i(1 + P_{qy})P_{\bar{q}x}}{\sqrt{2}(3 + \vec{P}_q \cdot \vec{P}_{\bar{q}})}$$

$$\rho_{0-1}^V = \frac{P_{qz}(1 - P_{\bar{q}y}) + (1 - P_{qy})P_{\bar{q}z} - iP_{qx}(1 - P_{\bar{q}y}) - i(1 - P_{qy})P_{\bar{q}x}}{\sqrt{2}(3 + \vec{P}_q \cdot \vec{P}_{\bar{q}})}$$

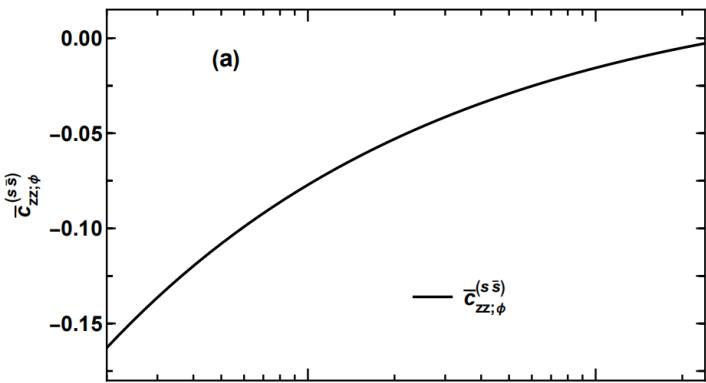
$$\rho_{1-1}^V = \frac{P_{qz}P_{\bar{q}z} - P_{qx}P_{\bar{q}x} + i(P_{qx}P_{\bar{q}y} + P_{qy}P_{\bar{q}x})}{3 + \vec{P}_q \cdot \vec{P}_{\bar{q}}}$$

sensitive to the long range correlation

They should be sensitive to the local correlations.

Global quark spin correlations in relativistic heavy ion collisions

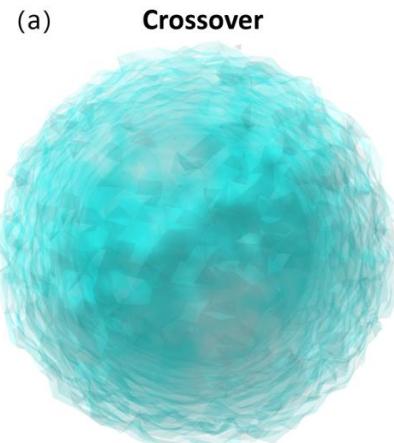
Ji-peng Lv,<sup>1,\*</sup> Zi-han Yu,<sup>1,†</sup> Zuo-tang Liang,<sup>1,‡</sup> Qun Wang,<sup>2,3,§</sup> and Xin-Nian Wang<sup>4,¶</sup>



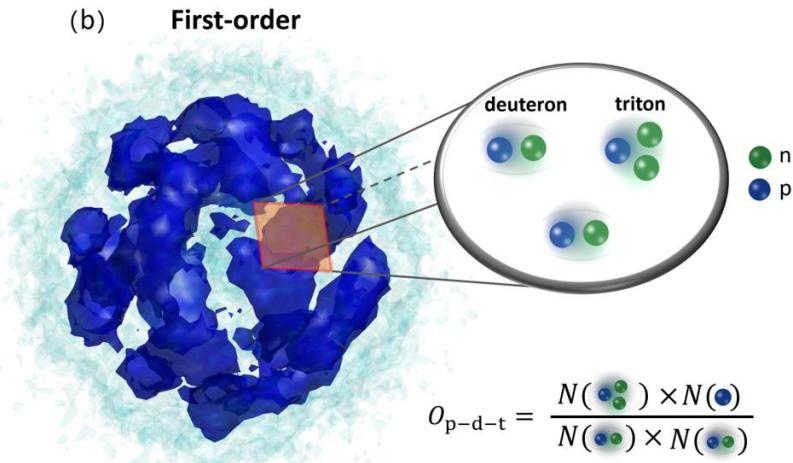
$\sqrt{s_{NN}}$  [GeV]

J. P. Lv et al., Phys. Rev. D 109, 114003 (2024)

C. M. Ko, NST 34, 80 (2023).



Crossover



First-order

$$O_{p-d-t} = \frac{N(\bullet\bullet) \times N(\bullet)}{N(\bullet\bullet) \times N(\bullet\bullet)}$$

$$N_d \approx N_d^{(0)}(1 + C_{np}) + \frac{3}{2^{1/2}} \left( \frac{2\pi}{mT} \right)^{3/2} \times \int d^3\mathbf{x}_1 d^3\mathbf{x}_2 C_2(\mathbf{x}_1, \mathbf{x}_2) \frac{e^{-\frac{(\mathbf{x}_1 - \mathbf{x}_2)^2}{2\sigma_d^2}}}{(2\pi\sigma_d^2)^{\frac{3}{2}}} \\ N_t \approx \frac{3^{3/2}}{4} \left( \frac{2\pi}{mT} \right)^3 \int d^3\mathbf{x}_1 d^3\mathbf{x}_2 d^3\mathbf{x}_3 \rho_{nnp}(\mathbf{x}_1, \mathbf{x}_2, \mathbf{x}_3) \times \frac{1}{3^{3/2}(\pi\sigma_t^2)^3} e^{-\frac{(\mathbf{x}_1 - \mathbf{x}_2)^2}{2\sigma_t^2} - \frac{(\mathbf{x}_1 + \mathbf{x}_2 - 2\mathbf{x}_3)^2}{6\sigma_t^2}},$$

$$\rho_{nnp}(\mathbf{x}_1, \mathbf{x}_2, \mathbf{x}_3) \approx \rho_n(\mathbf{x}_1)\rho_n(\mathbf{x}_2)\rho_p(\mathbf{x}_3) + C_2(\mathbf{x}_1, \mathbf{x}_2)\rho_p(\mathbf{x}_3) + C_2(\mathbf{x}_2, \mathbf{x}_3)\rho_n(\mathbf{x}_1) + C_2(\mathbf{x}_3, \mathbf{x}_1)\rho_n(\mathbf{x}_2) + C_3(\mathbf{x}_1, \mathbf{x}_2, \mathbf{x}_3)$$

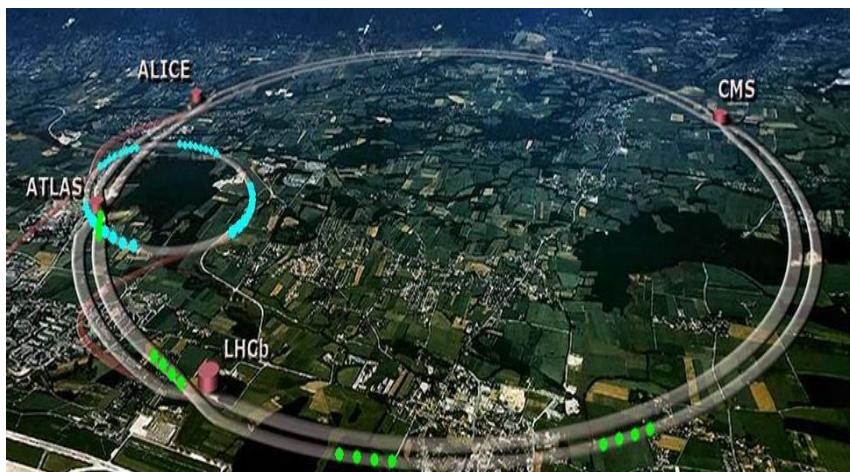
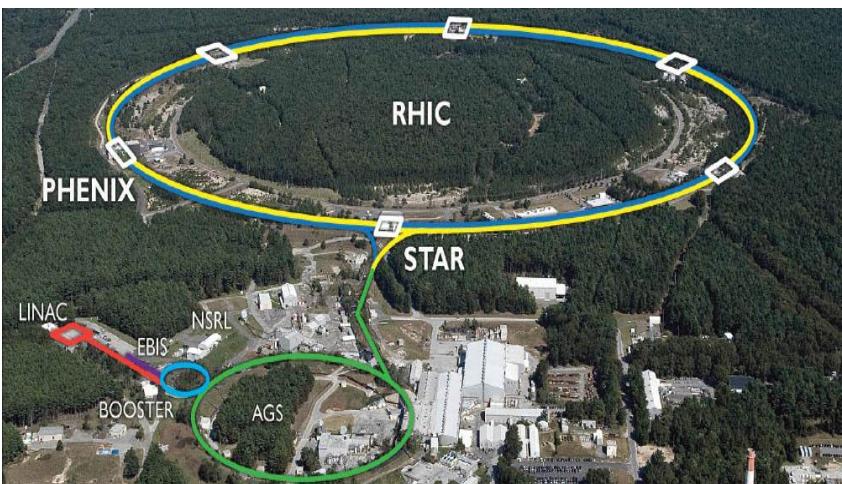
$$\frac{N_t N_p}{N_d^2} \approx \frac{1}{2\sqrt{3}} \left[ 1 + \Delta\rho_n + \frac{\lambda}{\sigma} G\left(\frac{\xi}{\sigma}\right) \right]$$

K. J. Sun, L. W. Chen, C. M. Ko, and Z. Xu, Phys. Lett. B 774, 103 (2017);  
K. J. Sun, C. M. Ko, and F. Li, PLB 816, 136258 (2021);

# Little Bang Nucleosynthesis

(2)

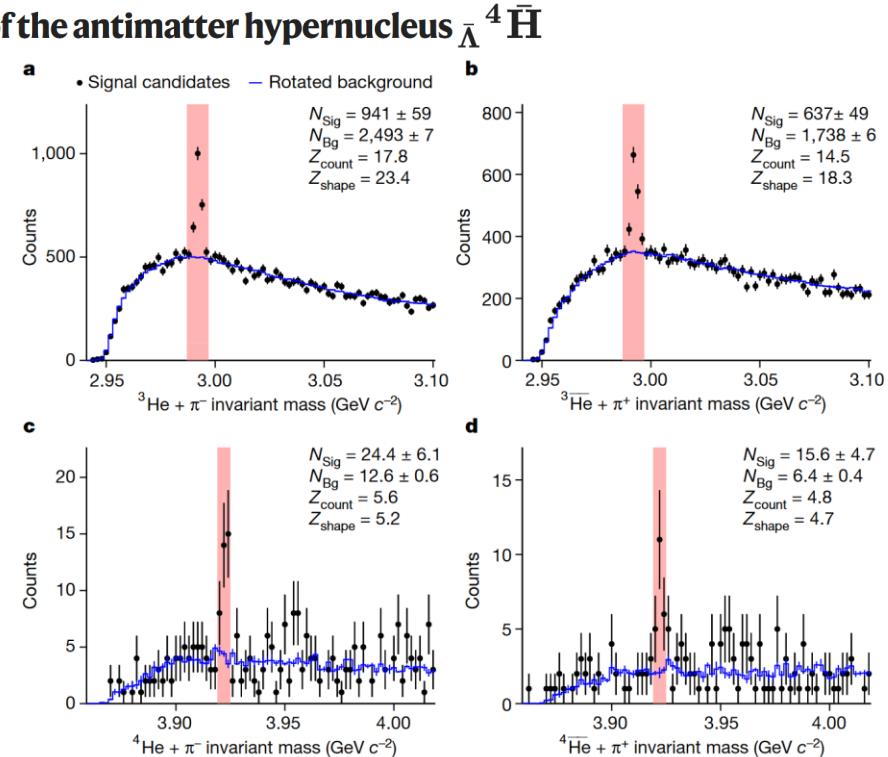
## Antimatter factory



## Observation of the antimatter hypernucleus $\bar{\Lambda}^4\bar{H}$

STAR Collaboration

[Nature \(2024\)](#) | [Cite this ar](#)



PHYSICAL REVIEW C 93, 064909 (2016)

## Antimatter $\bar{\Lambda}^4\bar{H}$ hypernucleus production and the ${}^3\text{H}/{}^3\text{He}$ puzzle in relativistic heavy-ion collisions

Kai-Jia Sun<sup>1</sup> and Lie-Wen Chen<sup>1,2,\*</sup>

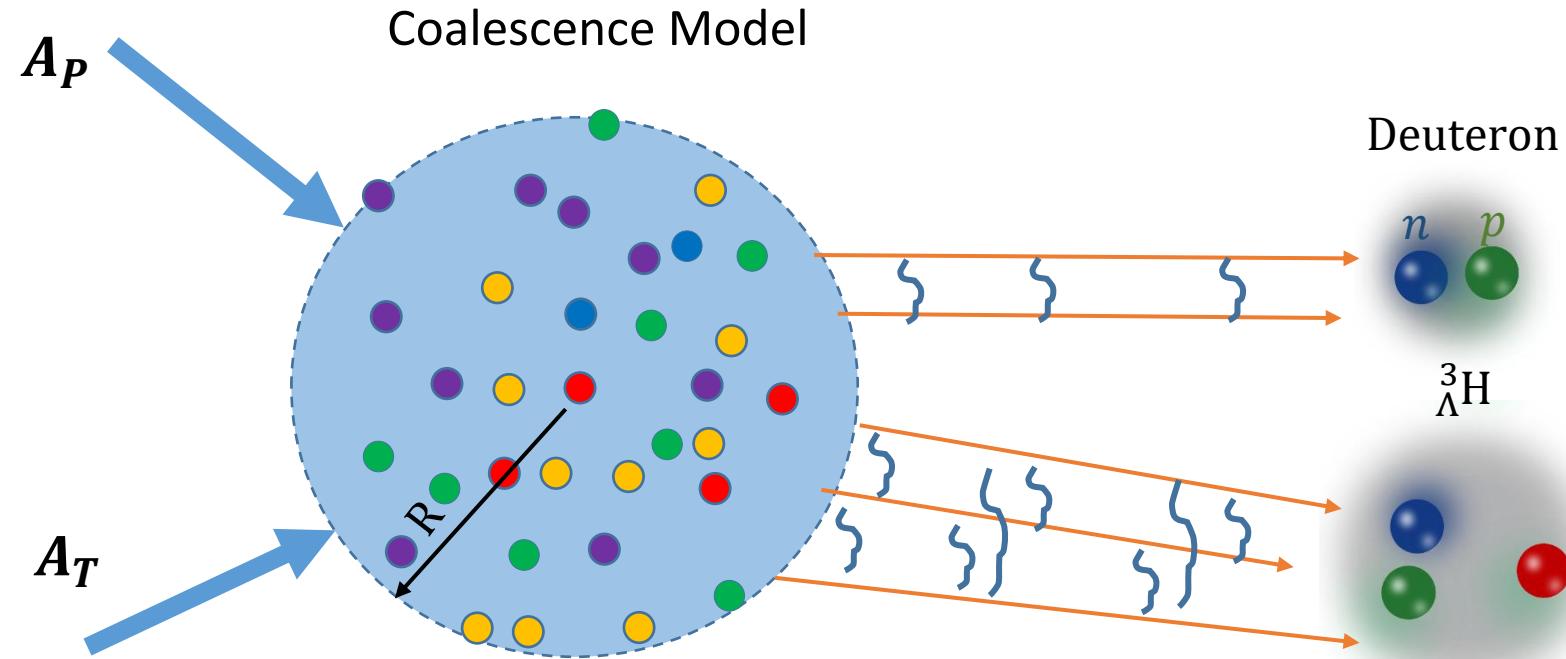
<sup>1</sup>Department of Physics and Astronomy and Shanghai Key Laboratory for Particle Physics and Cosmology, Shanghai Jiao Tong University, Shanghai 200240, China

<sup>2</sup>Center of Theoretical Nuclear Physics, National Laboratory of Heavy Ion Accelerator, Lanzhou 730000, China

(Received 26 April 2016; published 29 June 2016)

We show that the measured yield ratio  ${}^3\text{H}/{}^3\text{He}$  ( ${}^3\bar{\Lambda}\bar{H}/{}^3\text{He}$ ) in Au + Au collisions at  $\sqrt{s_{NN}} = 200$  GeV and in Pb + Pb collisions at  $\sqrt{s_{NN}} = 2.76$  TeV can be understood within a covariant coalescence model if (anti-) $\Lambda$  particles freeze out earlier than (anti-)nucleons but their relative freeze-out time is closer at  $\sqrt{s_{NN}} = 2.76$  TeV than at  $\sqrt{s_{NN}} = 200$  GeV. The earlier (anti-) $\Lambda$  freeze-out can significantly enhance the yield of (anti)hypernucleus  ${}^4\bar{\Lambda}\bar{H}$  ( ${}^4\bar{H}$ ), leading to that  ${}^4\bar{\Lambda}\bar{H}$  has a comparable abundance with  ${}^4\bar{\Lambda}\bar{H}$  and thus provides an easily measured antimatter candidate heavier than  ${}^4\text{He}$ . The future measurement on  ${}^4\text{H}/{}^4\bar{\Lambda}\bar{H}$  would be very useful to understand the (anti-) $\Lambda$  freeze-out dynamics and the production mechanism of (anti)hypernuclei in relativistic heavy-ion collisions.

# Final-state coalescence



Two-body coalescence  $a + b \rightarrow c$ :

$$N_c = \frac{2J_c + 1}{(2J_a + 1)(2J_b + 1)} \int \frac{dx_a dk_a}{(2\pi)^3} \frac{dx_b dk_b}{(2\pi)^3} f_a(x_a, k_a) f_b(x_b, k_b) W_c(x, k)$$

$$\approx \frac{2J_c + 1}{(2J_a + 1)(2J_b + 1)} \frac{N_a N_b}{(\frac{m_a m_b T}{m_a + m_b} (R_a^2 + R_b^2))^{3/2}} \times \frac{1}{(1 + \frac{\sigma^2}{R_a^2 + R_b^2})^{3/2}}$$

$$f_a = \frac{N_a}{(m_a T R_a^2)^{3/2}} e^{-\frac{k_a^2}{2m_a T} - \frac{x_a^2}{2R_a^2}}$$

$$N_a = \int \frac{dx_a dk_a}{(2\pi)^3} f_a(x_a, k_a)$$

$$W_c = 8e^{-x^2/\sigma^2 - \sigma^2 k^2}$$

$$1 = \int \frac{dx dk}{(2\pi)^3} W_c(x, k)$$

“Quantum mechanical correction”

Production

$$N_d \propto \frac{1}{\left[1 + \left(\frac{2r_d^2}{3R^2}\right)\right]^2}$$

Structure

$$N_{^3\text{H}} \propto \frac{1}{\left[1 + \left(\frac{r_{^3\text{H}}^2}{2R^2}\right)\right]^3}$$

can be inferred from Femtoscopy

Density Matrix Formulation  
(sudden approximation)

$$N_A = \text{Tr}(\hat{\rho}_s \hat{\rho}_A) \\ = g_c \int d\Gamma \rho_s(\{x_i, p_i\}) \times W_A(\{x_i, p_i\})$$

Wigner function of light cluster

Overlap between source distribution function and Wigner function of light nuclei

