



# 相对论重离子碰撞中矢量介子自旋 排列的最新实验测量进展

郝宝山 (Baoshan Xi)

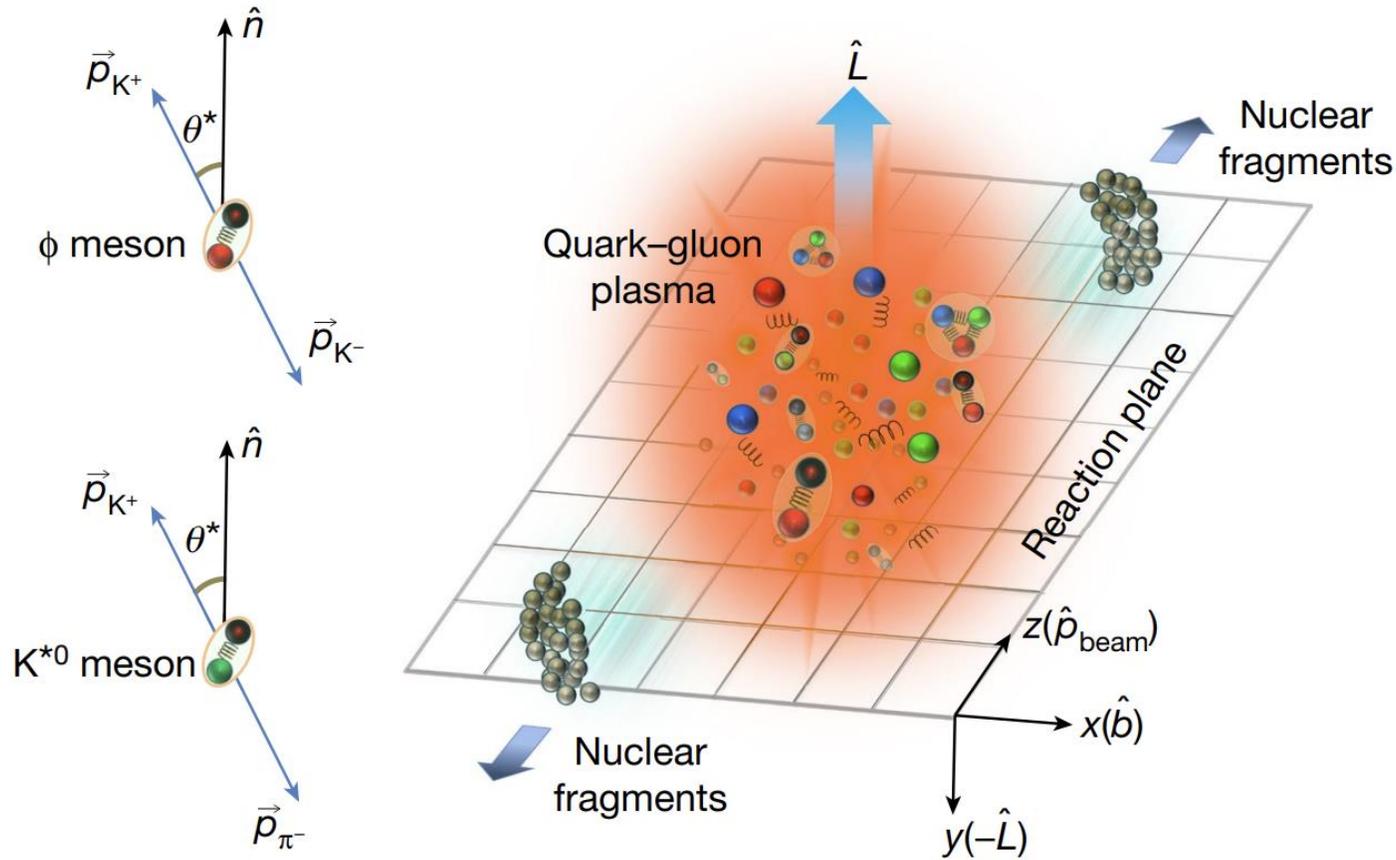
Fudan University

2024-10-13

# Outline

- Introduction & Development ( $\phi$ ,  $J/\Psi$  and so on)
- Experimental measurement and results
- Research motivation of global spin alignment of  $\rho^0$  meson
- Summary

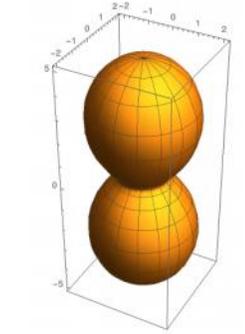
# QGP matter under rotation



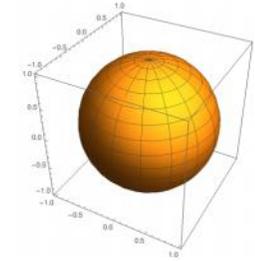
$$\frac{dN}{d\cos\theta^*} = N_0 \times \left[ (1 - \rho_{00}^{obs}) + (3\rho_{00}^{obs} - 1)\cos^2\theta^* \right]$$

**A deviation of  $\rho_{00}$  from 1/3 signals net spin alignment.**

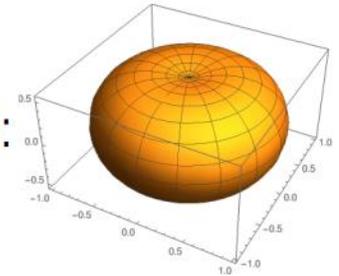
$\rho_{00} > 1/3$ :



$\rho_{00} = 1/3$ :



$\rho_{00} < 1/3$ :



STAR Collaboration, Nature 614 (2023) 7947.  
Z. T. Liang and X. N. Wang, Phys. Lett. B 629, (2005) 20.

# Progress in Theory and Experiment of Global Spin Alignment

- In theory, first predicted the possible existence of a global spin alignment (2005).
- STAR published experimental measurement results for  $\phi$  and  $K^{*0}$  in 2007 and 2023, respectively.
- Large global spin alignment for  $\phi$ -meson observed, so far can be accommodated by local fluctuation of strong force field. Various theory explanations are rapidly developing.
- ALICE measured  $J/\Psi$ ,  $D^{*+}$  and  $D^*$ .
- STAR is measuring  $\rho^0$ ,  $J/\Psi$ , and the rapidity ( $y$ ) dependence of  $\phi$ 's
- Theoretically studied the rapidity dependence of the global spin alignment of  $\phi$ .

Z. T. Liang and X. N. Wang, Phys. Rev. Lett. 96 (2006) 039901.

Z. T. Liang and X. N. Wang, Phys. Lett. B 629, (2005) 20.

X. L. Sheng, L. Oliva, Z. T. Liang, Q. Wang, X. N. Wang, Phys. Rev. Lett. 131 (2023) , 042304.

X. L. Sheng, L. Oliva, Q. Wang, Phys. Rev. D 101 (2020) 9.

X. L. Sheng, S. Pu, Q. Wang, Phys. Rev. C 108 (2023) 5, 054902.

STAR Collaboration, Phys. Rev. C 77 (2008) 061902.

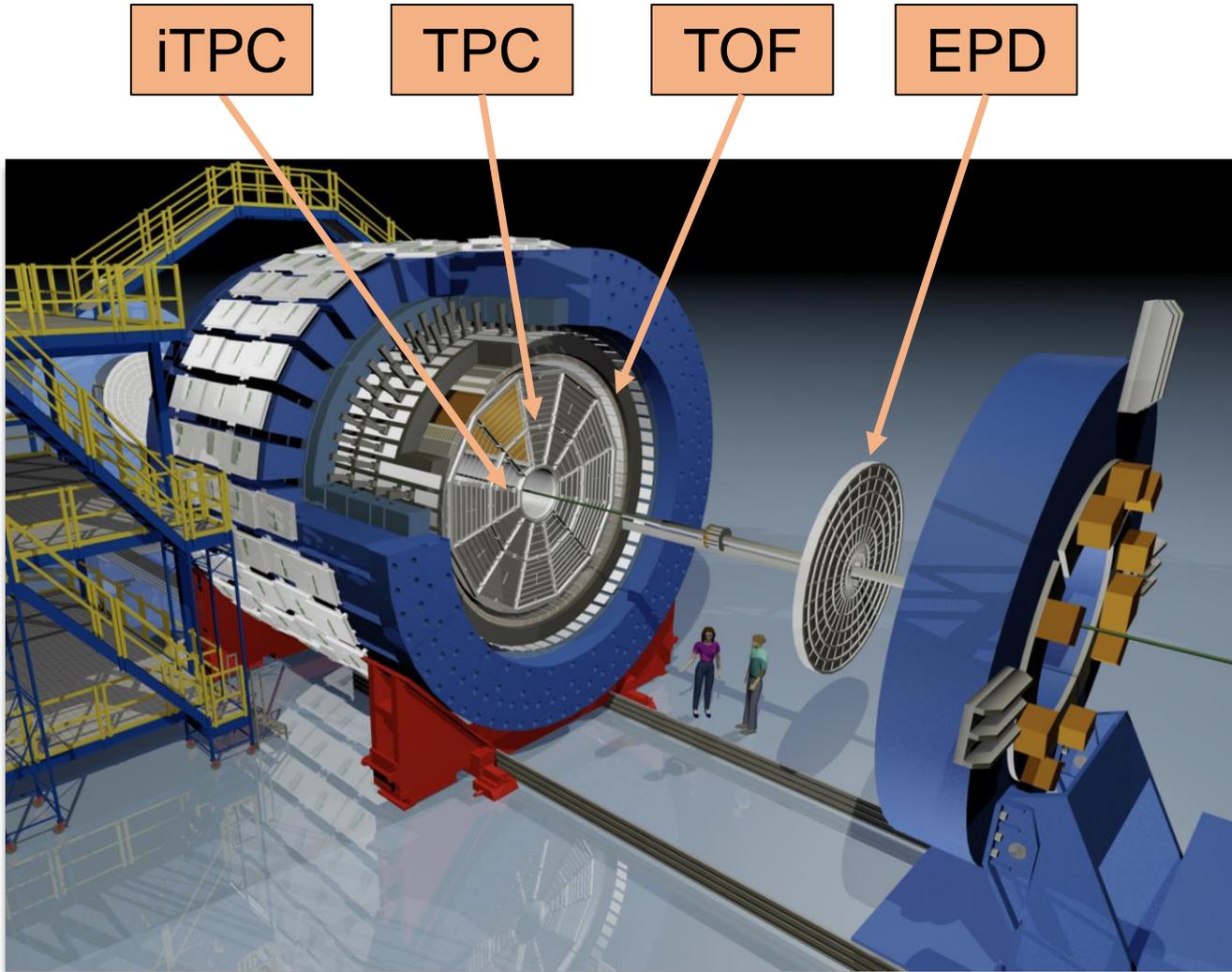
STAR Collaboration, Nature 614 (2023) 7947.

ALICE Collaboration, Phys. Rev. Lett. 131 (2023) 042303.

ALICE Collaboration, Phys. Lett. B 846 (2023) 137920.

<https://indico.cern.ch/event/1139644/contributions/5540118/>.

# The STAR Detector



Full azimuthal coverage

TPC\* :  $|\eta| < 1$

iTPC\* :  $|\eta| < 1.5$

tracking, centrality, particle identification, and event plane.

TOF\* :  $|\eta| < 0.9$

eTOF:  $-1.6 < \eta < -1.1$  (not shown)

particle identification

BBC :  $3.3 < |\eta| < 5$  (not shown)

EPD\* :  $2.1 < |\eta| < 5.1$

event plane reconstruction

**Greater EP resolution with EPD**

\*Used in  $\phi$  analysis

# Second order event plane

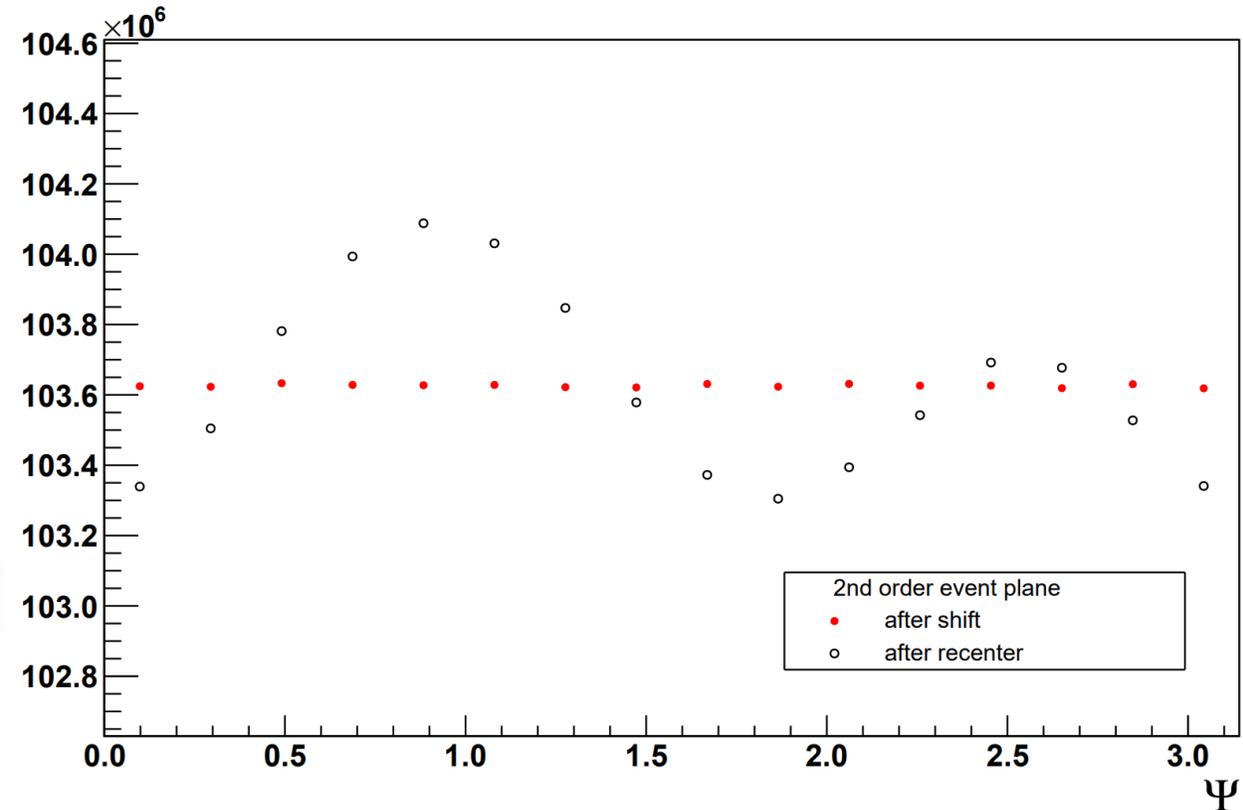
The second event plane is obtained from TPC (for 200GeV data) and flattened by recentering and shifting (performed every 10 runs).

After the recenter corrected, we get the recenter factors:

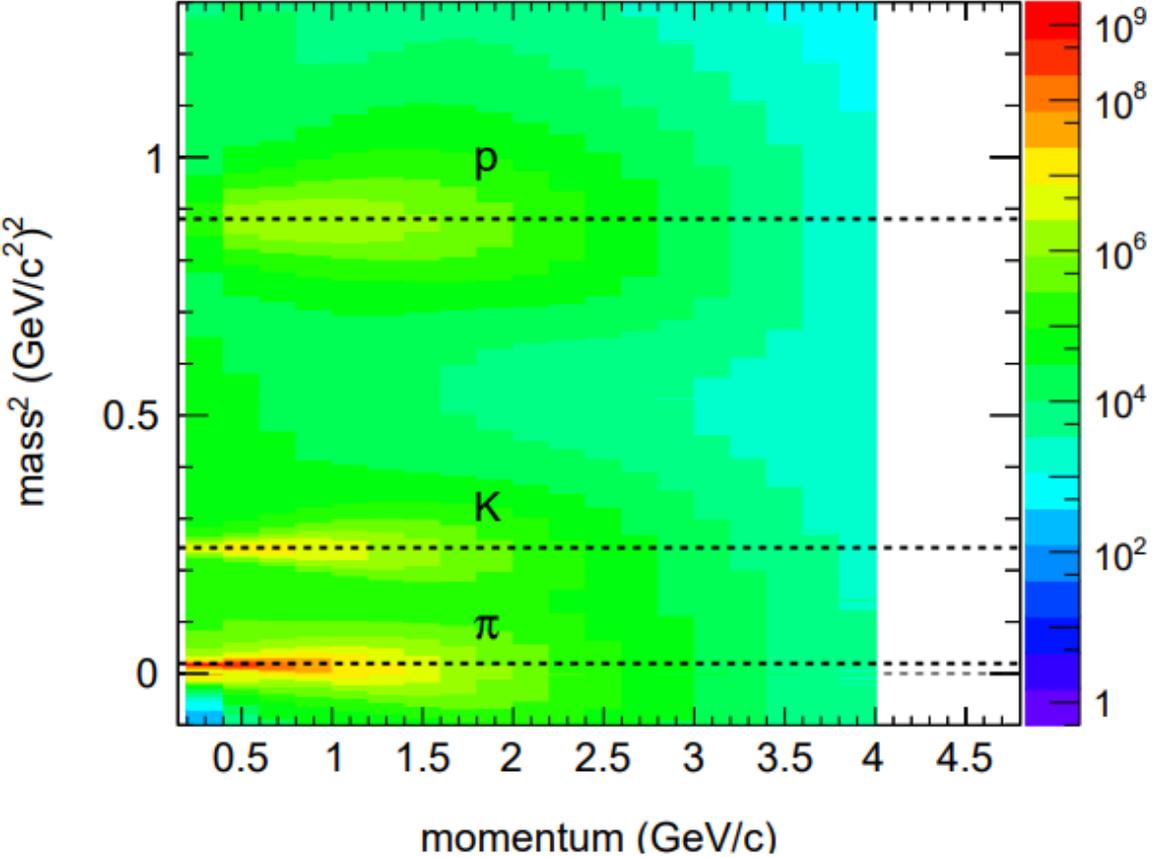
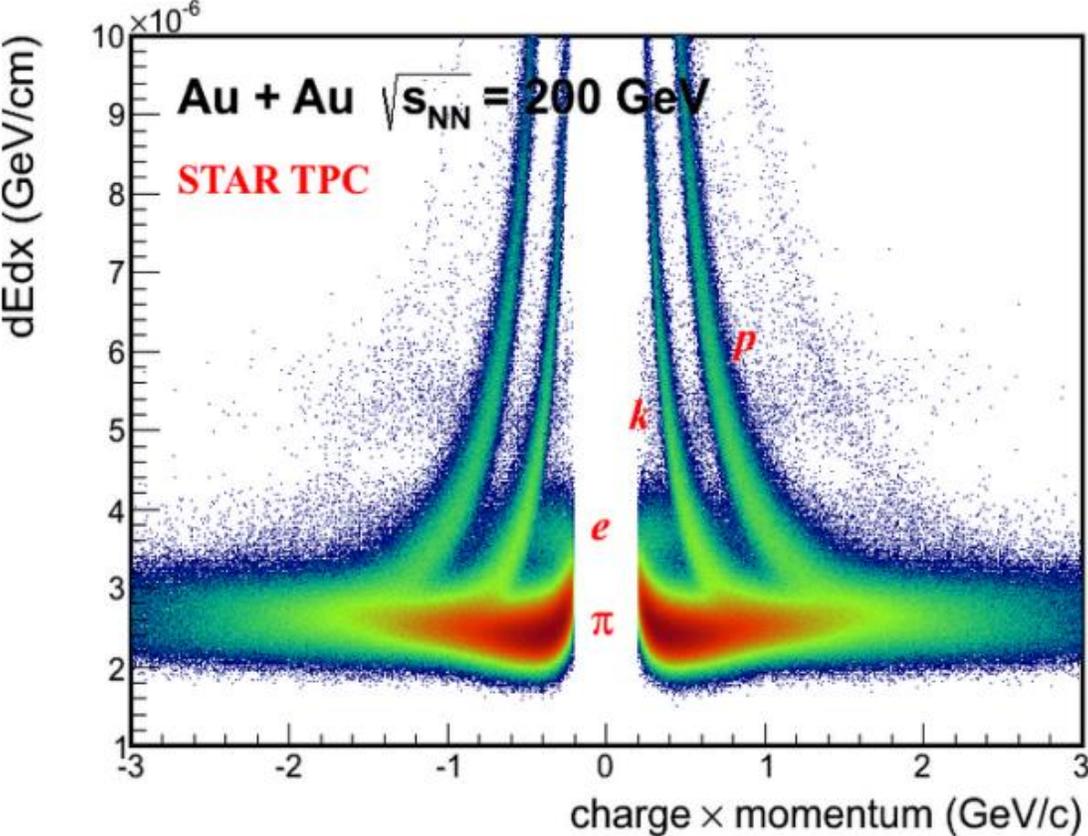
$$Q_x = \sum W \cdot p_T \cdot \cos 2\phi \quad Q_y = \sum W \cdot p_T \cdot \sin 2\phi$$

Then we do the shift

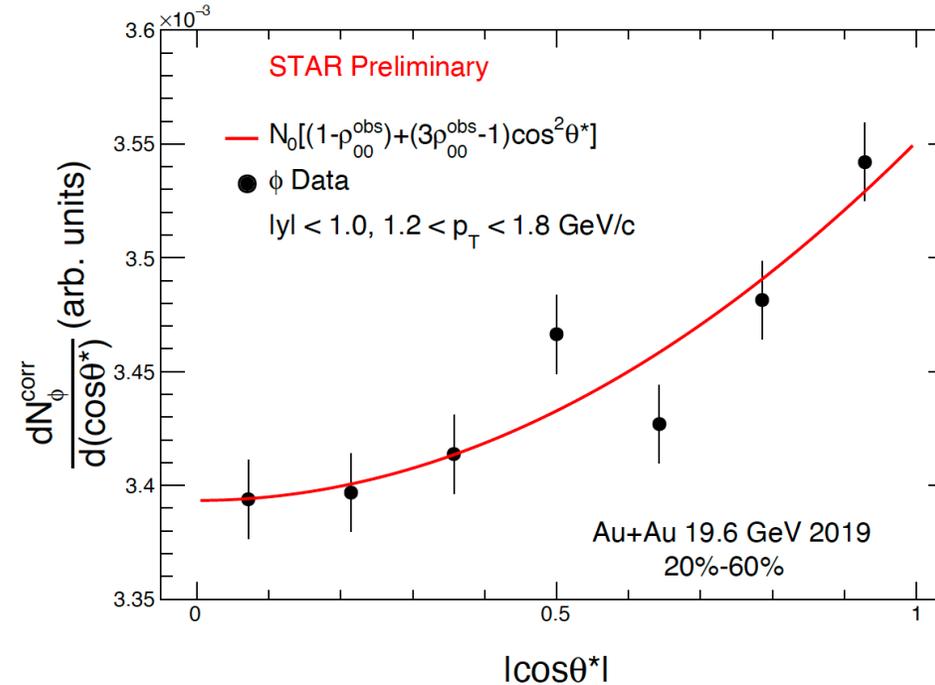
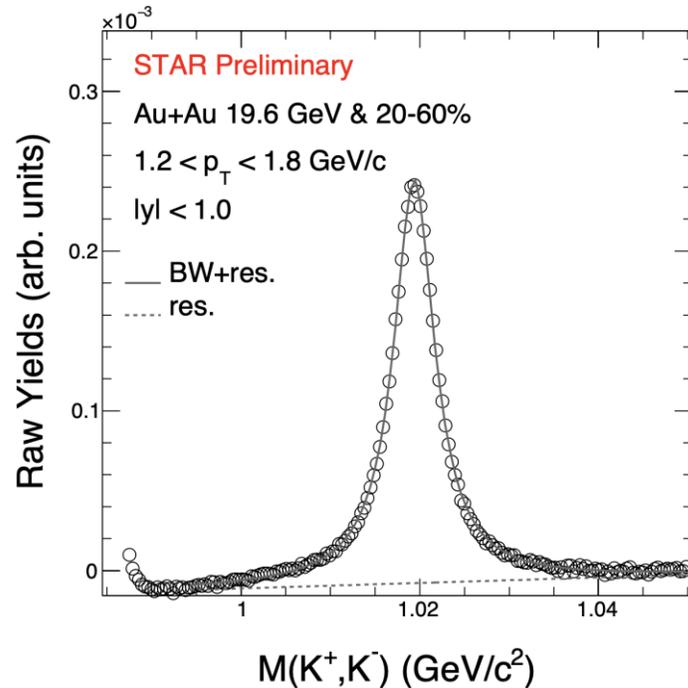
$$2\Psi_2 = 2\Psi'_2 + \sum_{i=1}^n \frac{2}{i} [-\langle \sin(i2\Psi'_2) \rangle \cos(i2\Psi'_2) + \langle \cos(i2\Psi'_2) \rangle \sin(i2\Psi'_2)]$$



# Particle identification



# How to measure( $\phi$ ): signal reconstruction



- Use event-mixing to subtract background and extract yields from histogram integration in seven  $\cos\theta^*$  bins.
- Yields vs.  $|\cos\theta^*|$  are corrected for the acceptance and efficiency.
- $\rho_{00}^{obs}$  is extracted from a fit to the corrected yields vs.  $|\cos\theta^*|$
- Then calculate  $\rho_{00}$  from  $\rho_{00}^{obs}$  accounting for EP resolution:  $\rho_{00} = \frac{1}{3} + \frac{4}{1+3R} \left( \rho_{00}^{obs} - \frac{1}{3} \right)$ .

K. Schilling, P. Seyboth and G. E. Wolf, Nucl. Phys. B 15 (1970) 397.  
 A. H. Tang , B. Tu and C. S. Zhou, Phys. Rev. C 98 (2018).

# How to measure: acceptance and EP resolution

## Corrections for finite EP resolution, efficiency, and acceptance

i)  $\phi$ -meson  $\rho_{00}$  analysis Detector efficiency within the acceptance is corrected using the STAR Monte Carlo embedding method<sup>8-10</sup>. To account for finite EP resolution and finite acceptance in pseudo-rapidity ( $\eta$ )<sup>11</sup>, the observed  $\cos \theta^*$  distribution is not fitted using Eq. 1 in the main text, but is instead described by the correction procedure derived in Ref. <sup>7</sup> wherein the data are fitted using

$$\left[ \frac{dN}{d \cos \theta^*} \right]_{|\eta|} \propto \left( 1 + \frac{B'F}{2} \right) + (A' + F) \cos^2 \theta^* + (A'F - \frac{B'F}{2}) \cos^4 \theta^*, \quad (1)$$

where

$$A' = \frac{A(1+3R)}{4+A(1-R)}, \quad B' = \frac{A(1-R)}{4+A(1-R)}, \quad (2)$$

and

$$A = \frac{3\rho_{00} - 1}{1 - \rho_{00}}, \quad (3)$$

$$\frac{dN}{d \cos \theta^* d\beta} \propto 1 + A \cos^2 \theta^* + B \sin^2 \theta^* \cos 2\beta + C \sin 2\theta^* \cos \beta.$$

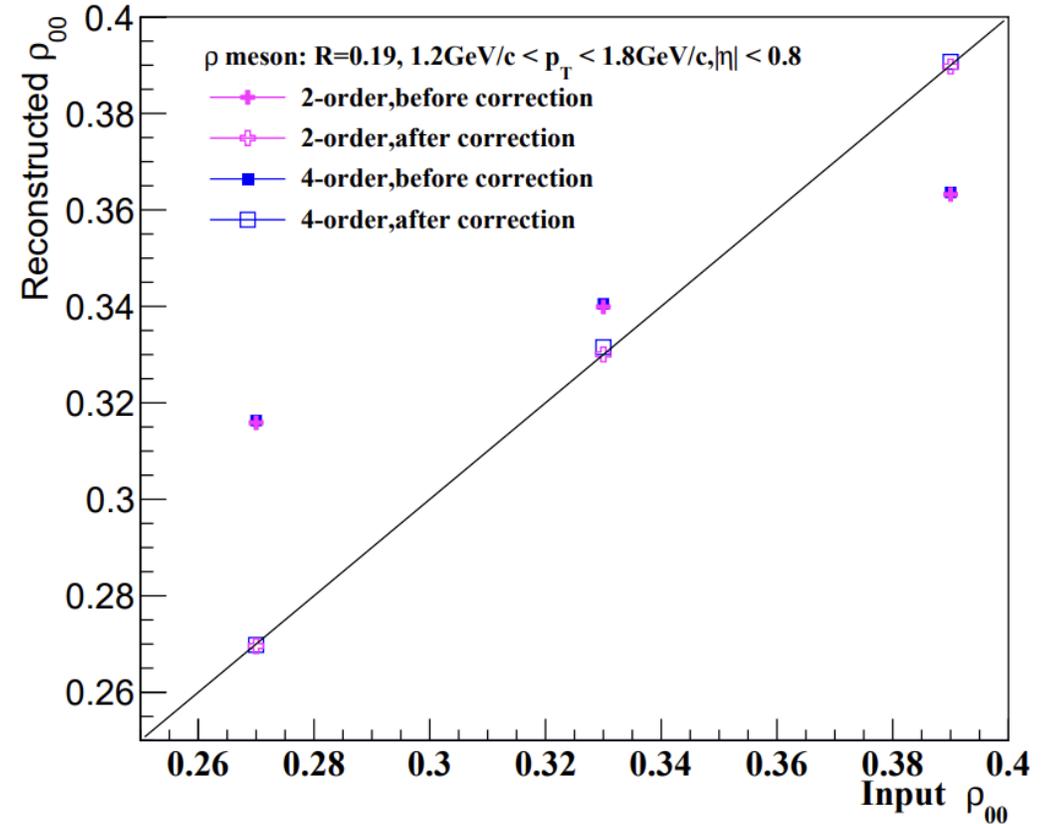
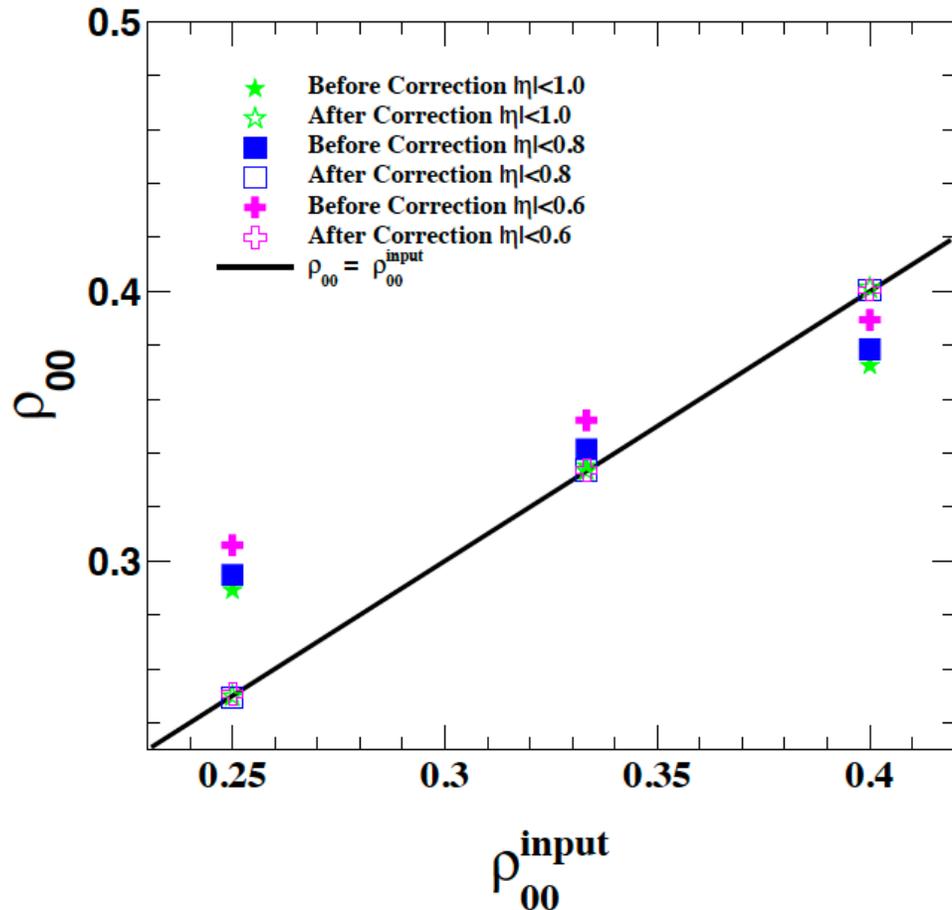
$$\left[ \frac{dN}{d \cos \theta^* d\beta} \right]_{|\eta|} \propto 2 + F - \frac{BF}{2} + \frac{3G}{4} - \frac{BG}{2} + \left[ 2A - F(1 - A - B) - G \left( \frac{3}{2} - \frac{3A}{4} - \frac{3B}{2} \right) \right] \cos^2 \theta^* + \left[ -F \left( A + \frac{B}{2} \right) + G \left( \frac{3}{4} - \frac{3A}{2} - \frac{3B}{2} \right) \right] \cos^4 \theta^* + \left[ G \left( \frac{3A}{4} + \frac{B}{2} \right) \right] \cos^6 \theta^*.$$

$$A = \frac{A'(1+3R)}{4+A'(1-R)}, \quad B = \frac{A'(1-R)}{4+A'(1-R)}, \quad A' = \frac{3\rho_{00} - 1}{1 - \rho_{00}}$$

We followed the same correction procedure as used in published paper. Instead of expanding the  $\cos \theta$  distribution to the 2nd order, we expanded it to the fourth order. In practice this makes the fitting better but has little effect on the extracted  $\rho_{00}$ .

STAR Collaboration, Nature 614 (2023) 7947.

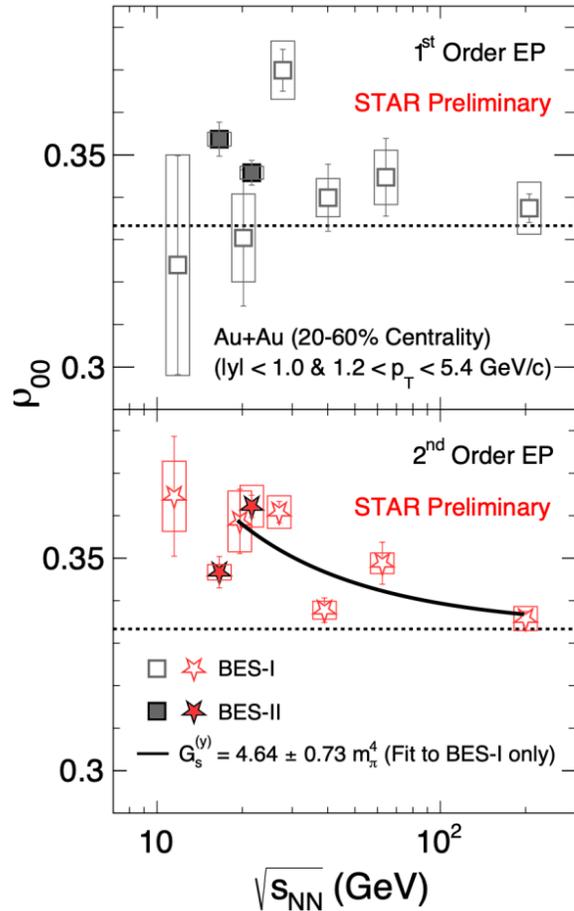
# Simulation: acceptance and EP resolution



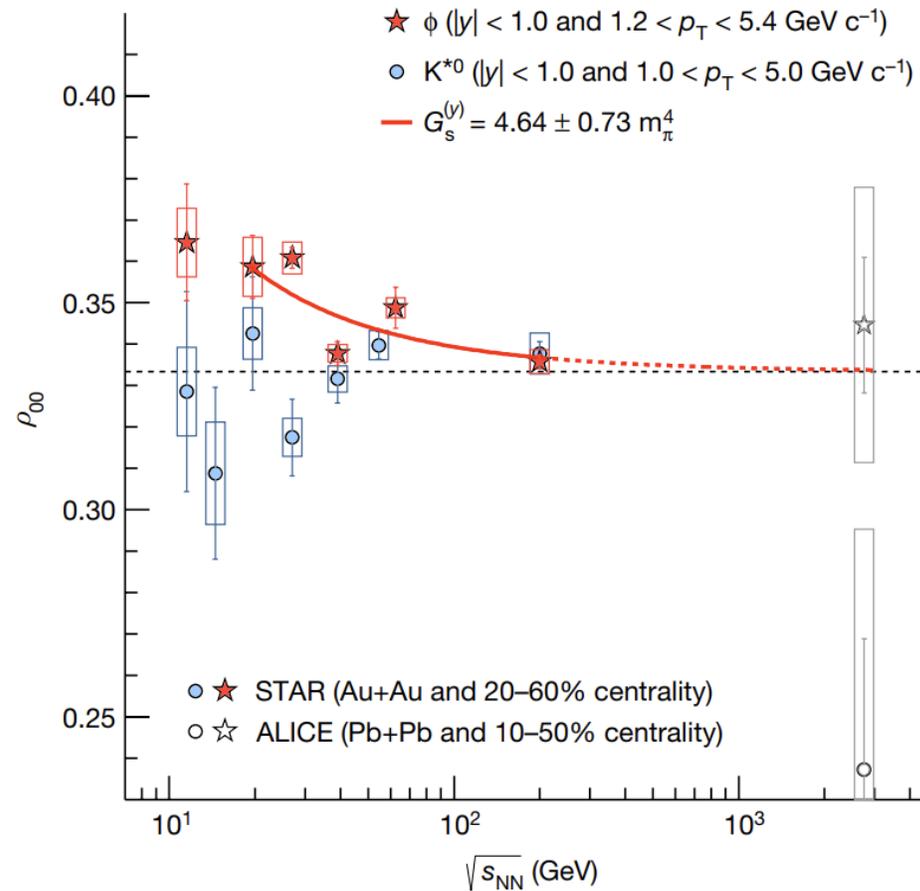
A. H. Tang , B. Tu and C. S. Zhou, Phys. Rev. C 98 (2018), 044907.

The simulation verified the effectiveness of the correction method considering event plane resolution and acceptance.

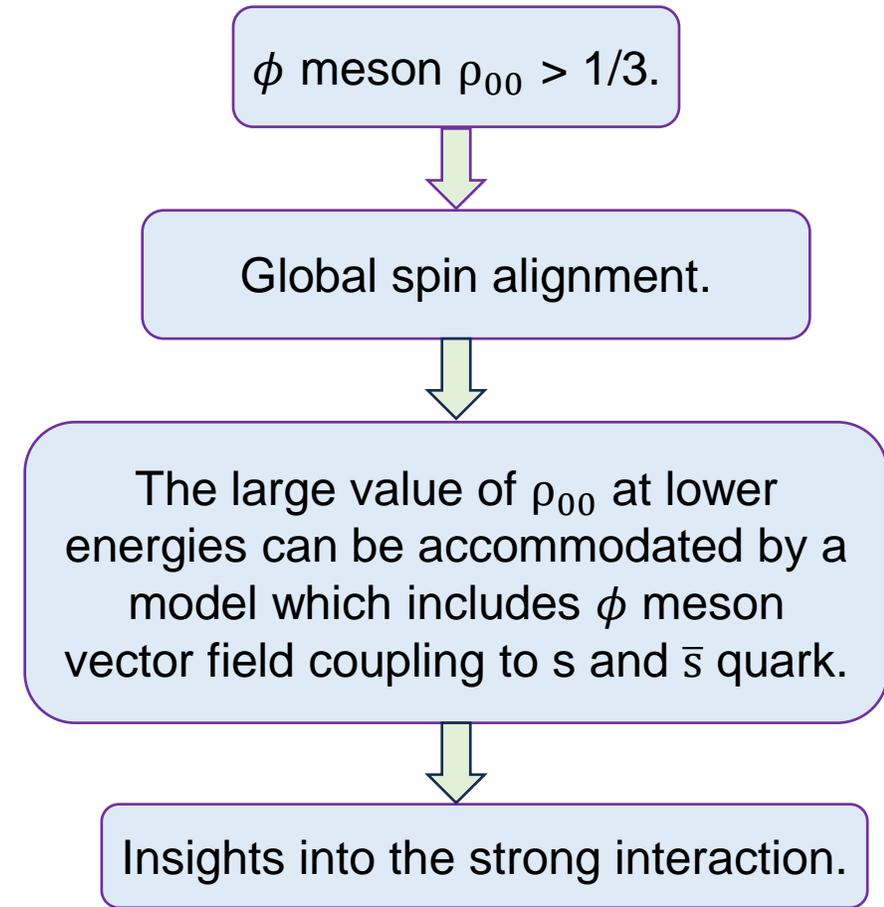
# STAR measurement of global spin alignment



Gavin for STAR, sQM 2024.



STAR Collaboration, Nature 614 (2023) 7947.



# Motivation for studying $\phi$ spin alignment in BES-II

- Large global spin alignment for  $\phi$ -meson observed, so far can be accommodated by local fluctuation of strong force field. Various theory explanations are rapidly developing.

STAR Collaboration, Nature 614 (2023) 7947.

X. L. Sheng, L. Oliva, Q. Wang, Phys. Rev. D 101 (2020) 9.

- Increasing with decreasing energy, will trend remains at lower energies?

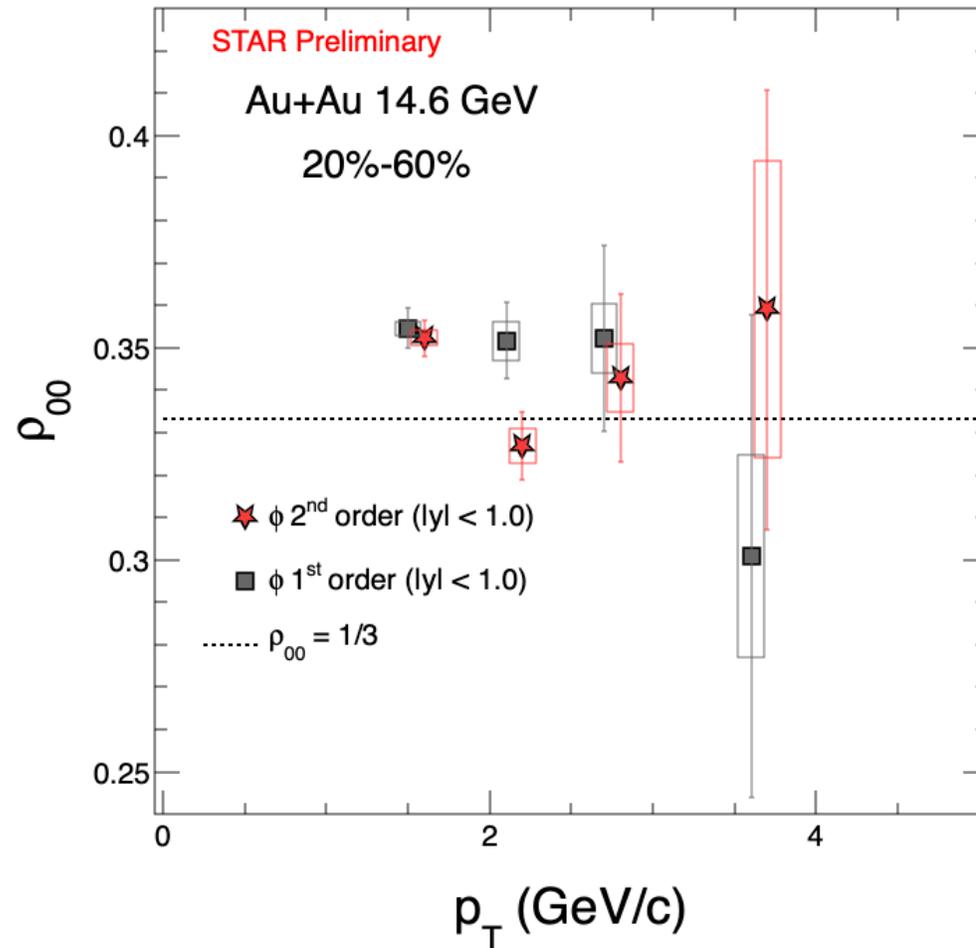
Clarify  $\rho_{00}$  behavior in lower energy regime.

- BES-II provides significantly more statistics for lower collision energies.

BES-I 19.6 GeV:  $\sim 1.9 \times 10^7$  events after cuts.

BES-II 19.6 GeV:  $\sim 4.6 \times 10^8$  events after cuts.

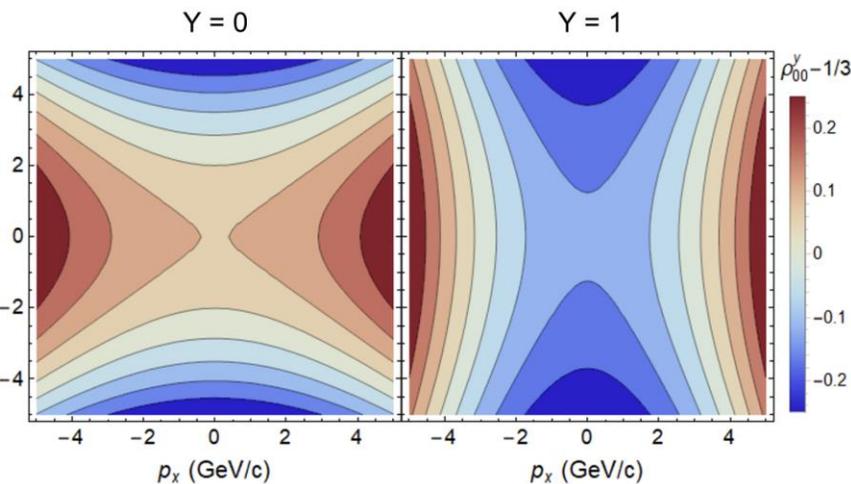
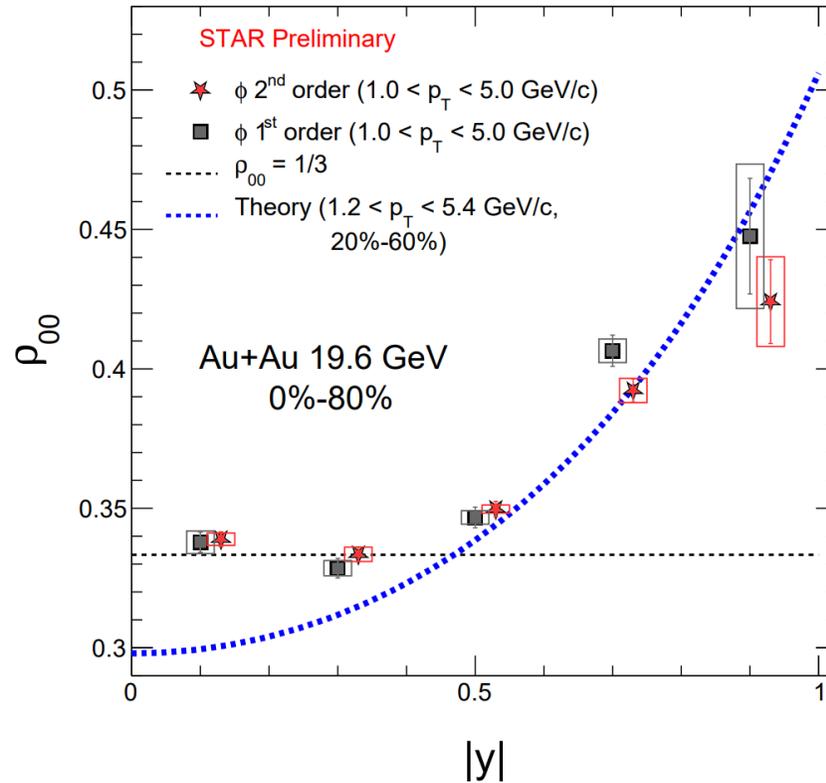
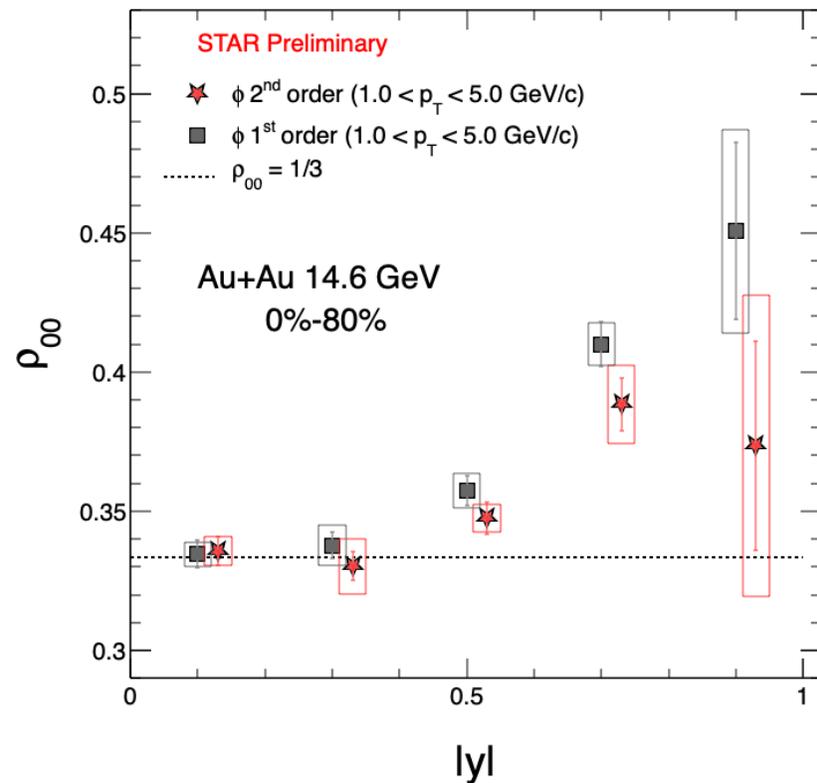
# $\rho_{00}$ : 1<sup>st</sup> order EP vs. 2<sup>nd</sup> order EP in BES-II



B. Xi for STAR, Quark Matter 2023.

Results obtained with 1st and 2nd order EP are consistent with each other. Both of them show significant signals.

# $\rho_{00}$ of $\phi$ as a function of rapidity ( $y$ )

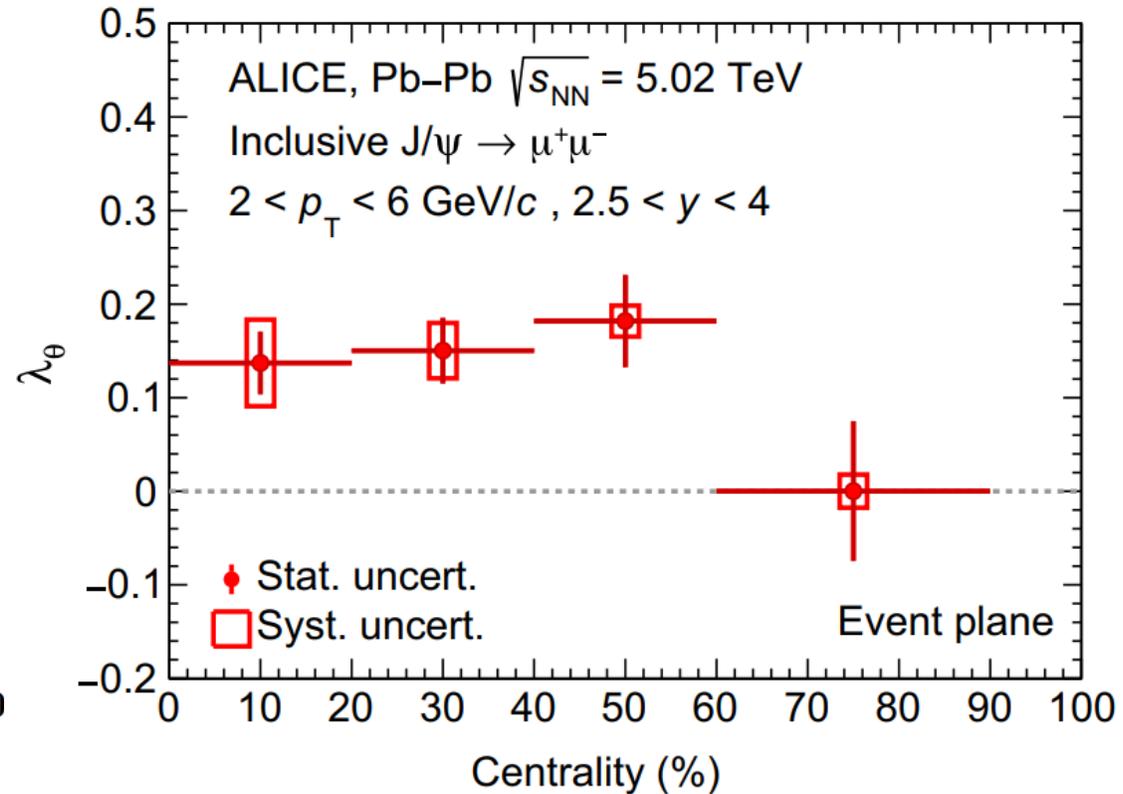
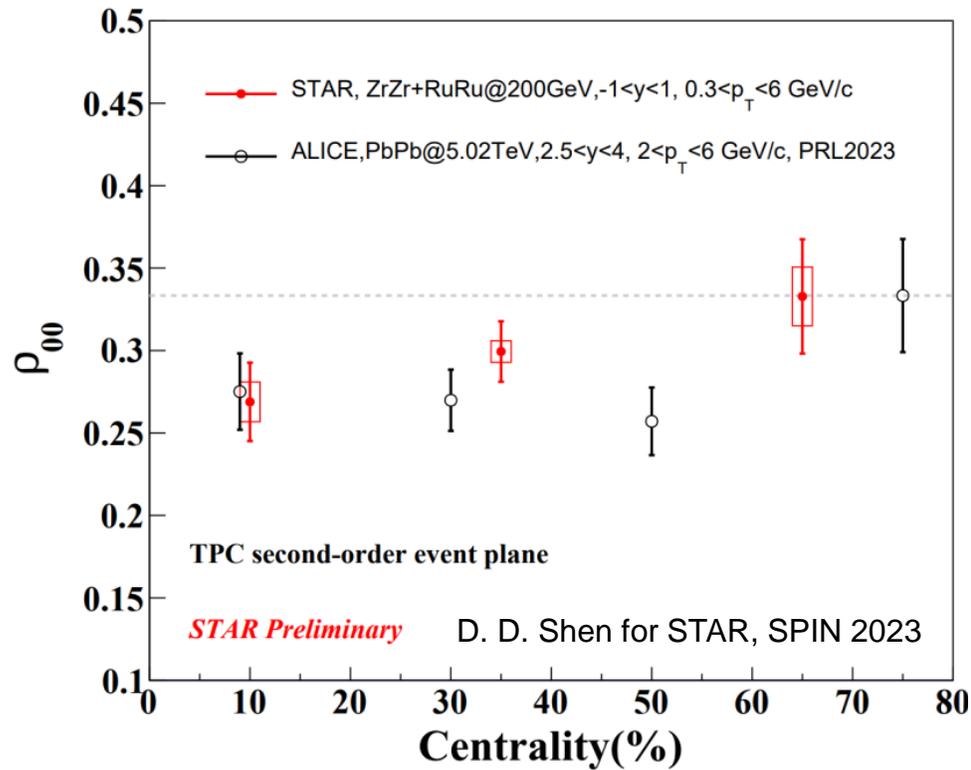


B. Xi for STAR, Quark Matter 2023.

X. L. Sheng, S. Pu, Q. Wang, Phys. Rev. C 108 (2023) 5, 054902.

**The trend is consistent with theoretical calculation.**  
**The behavior is due to larger fluctuation in the direction perpendicular to the motion direction.**

# $\rho_{00}$ of $J/\Psi$ in Isobar (RuRu and ZrZr) and PbPb



$$\lambda_\theta = \frac{1 - 3\rho_{00}}{1 + \rho_{00}}$$

$J/\Psi$ , like  $\phi$ , has also two quarks from same flavor family, making its  $\rho_{00}$  an interesting measurement.

$$\frac{dN}{d \cos \theta^*} \propto (1 + \rho_{00}) + (1 - 3\rho_{00}) \cos^2 \theta^*$$

ALICE Collaboration, Phys. Rev. Lett. 131 (2023) 042303.

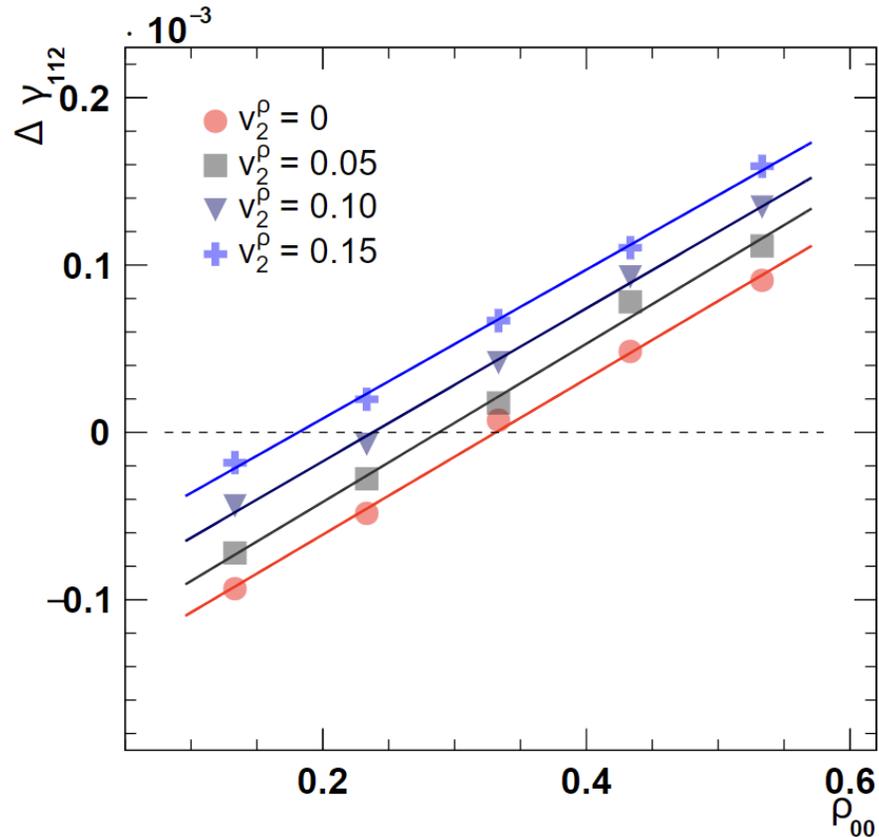
SPIN 2023, <https://indico.jlab.org/event/663/contributions/13003/>.

P. Faccioli, C. Lourenco, J. Seixas and H. K. Wohri, Eur. Phys. J. C 69, 657 (2010).

# Motivation for studying $\rho^0$ spin alignment

- $\rho^0$ -meson consists of a mixture of quark flavor  $(u\bar{u} - d\bar{d})/\sqrt{2}$ .
- $\rho^0$ -mesons are mostly regenerated, not directly from quark coalescence.
- The two reasons above make it not a good probe to local fluctuation of strong force field like  $\phi$  ( $s\bar{s}$ ) meson.
- However, the global spin alignment of  $\rho^0$ -meson has notable implication on CME analyses.

# Implications of $\rho^0$ spin alignment on CME search



Global spin alignment of  $\rho^0$  meson can contribute to background in CME observables, similar to resonance  $v_2$  effect.

( $\rho_{00} > 1/3$ ) will enhance apparent values of CME observables.  
( $\rho_{00} < 1/3$ ) will decrease apparent values of CME observables.

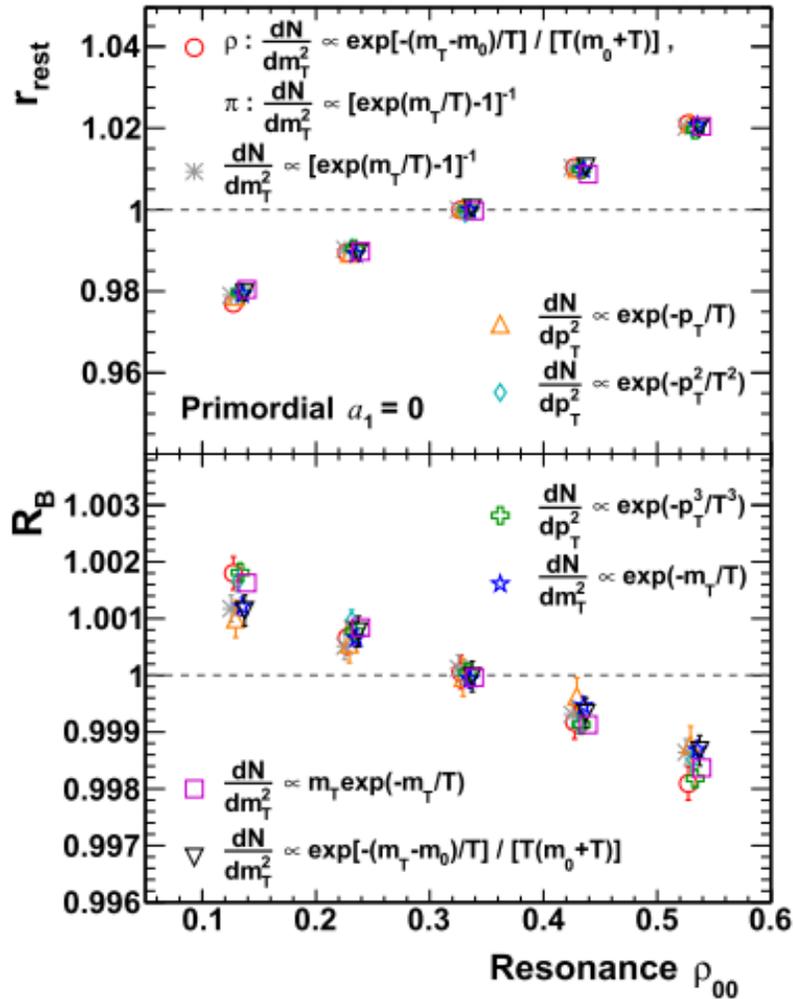
A. H. Tang, Chin. Phys. C 44 (2020) 054101.

D. Shen, J. Chen, A. Tang and G. Wang, Phys. Lett. B 839 (2023) 137777.

Toy model simulations of the  $\pi$ - $\pi$   $\Delta\gamma_{112}$  correlation as a function of  $\rho^0$  meson  $\rho_{00}$  with various inputs of  $v_2^\rho$ .

**To assess its effect in CME observables, it would be desirable to study  $\rho^0$  meson  $\rho_{00}$ .**

# Implication of $\rho^0$ spin alignment on CME search



The influence of  $\rho_{00}$  of  $\rho^0$ -meson on CME observations (signed balance function)

To assess its effect in CME observables, it would be desirable to study  $\rho^0$  meson  $\rho_{00}$ .

Fig. 9. (color online)  $r_{rest}$  and  $R_B$  as a function of resonance  $\rho_{00}$  for various transverse spectra. Choices of spectra are the same as in Fig. 7.

A. H. Tang, Chinese Phys. C 44 (2020) 5, 054101

# Measurement: yield extraction for $\rho^0$ meson

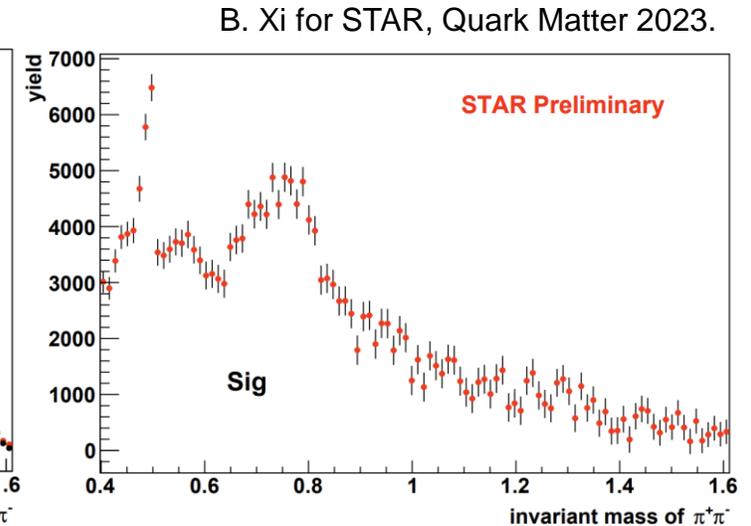
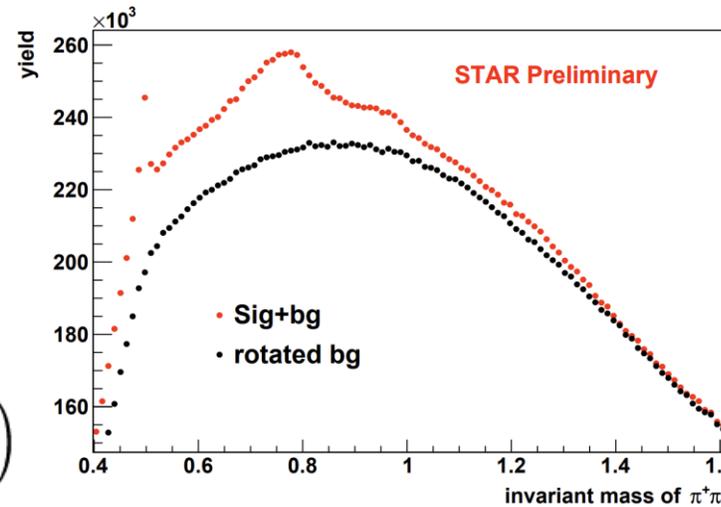
- We first subtract the rotated background. Then do normalization and use a second order polynomial to take care of residual background.
- We fit with contributions from 7 particles (hadronic cocktail fitting) :  $\omega, \rho^0, f_0, f_2, \sigma^0, k_S^0, \eta$ .
- At each  $p_T$  interval, we fit the mass and width of the particles.

$$F(\rho^0, \text{ or } f_0, f_2, \sigma^0) = PS(M_{\pi\pi}) \times BW(M_{\pi\pi})$$

$$BW(M_{\pi\pi}) = \frac{AM_{\pi\pi}M_0\Gamma(M_{\pi\pi})}{\left[ (M_0^2 - M_{\pi\pi}^2)^2 + M_0^2\Gamma^2(M_{\pi\pi}) \right]}$$

$$PS(M_{\pi\pi}) = \frac{M_{\pi\pi}}{\sqrt{M_{\pi\pi} + p_T^2}} \times \exp\left(-\sqrt{M_{\pi\pi} + p_T^2}/T\right)$$

$$\Gamma(M_{\pi\pi}) = \left[ \frac{(M_{\pi\pi}^2 - 4m_\pi^2)}{(M_0^2 - 4m_\pi^2)} \right]^{(2J+1)/2} \times \Gamma_0 \times (M_0/M_{\pi\pi})$$



J. Adams, et al. (STAR Collaboration), Phys. Rev. Lett. 92, (2004) 092301.  
Prabhat R. Pujahari (for the STAR collaboration), Nucl. Phys. A 862, (2011) 297.

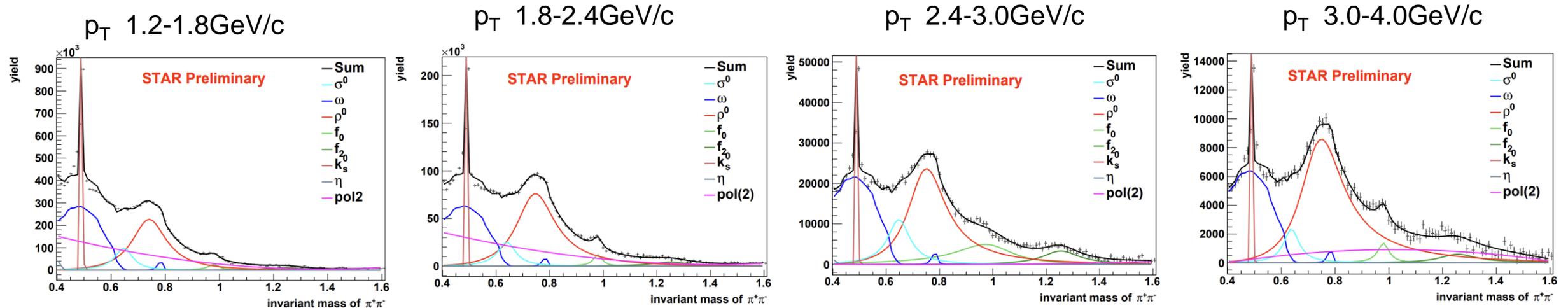
# Measurement: overall fitting (Integrating 7 $\cos\theta^*$ bins together)

We do an overall fitting for each  $p_T$  bin at 60-80% with the following constraints of parameters:

- $\rho^0$  : mass: 0.7-0.8                      width: free parameter
- $f_0$  : mass: 0.98 (fixed)                      width: free parameter
- $\sigma^0$  : mass: 0.4 - 0.8                      width: 0.1 - 0.8(PDG)
- $k_S^0$  : Gaussian function with mass and width as free parameters.
- $\omega, \eta$ : with function shapes obtained from hijing simulation.
- $f_2$  : mass 1.275    width 0.185

Through cocktail fitting, we get the mass and width of particles.

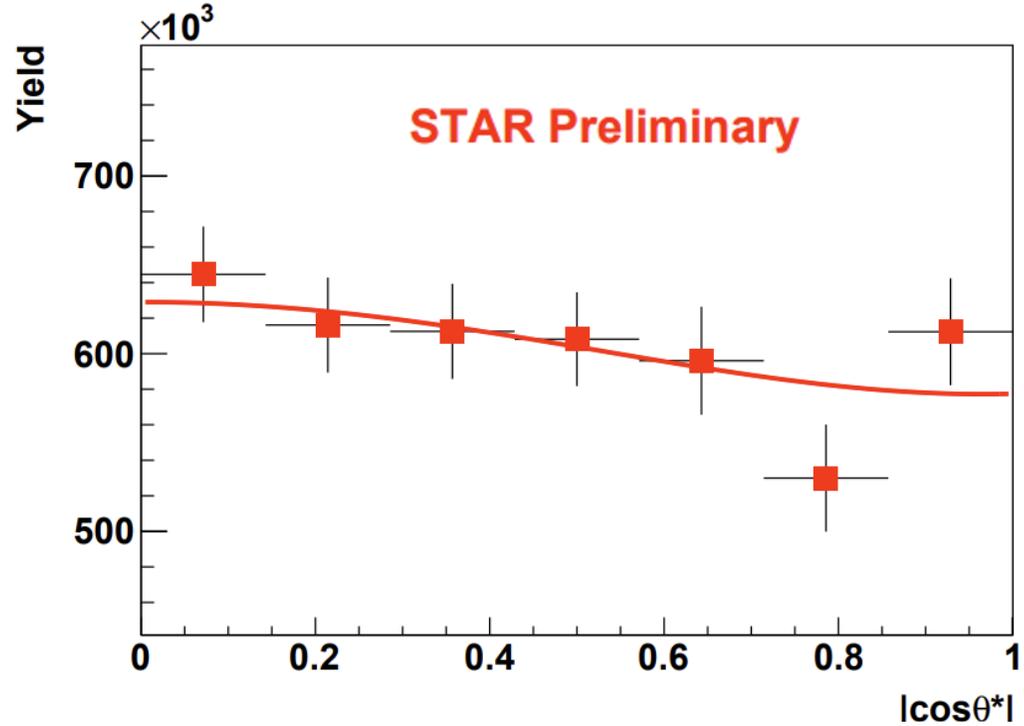
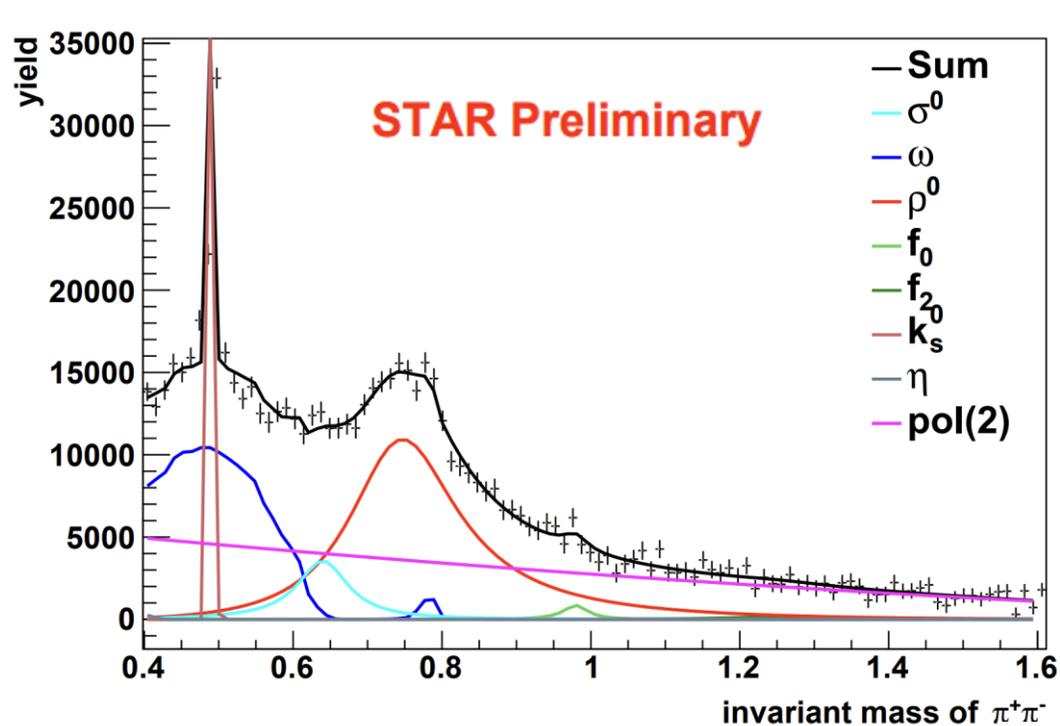
As an example, we show the overall fitting results of AuAu collision in different  $p_T$  at 60-80% centrality.



# Measurement: for each $\cos\theta^*$ bin

Based on the information obtained from the overall fitting, we fix the mass and width of  $\sigma^0$  and  $k_s^0$ , and fix the width of  $\rho^0$  and  $f_0$ .

Through cocktail fitting, the yields of  $\rho^0$  can be obtained. Then corrected for efficiency.



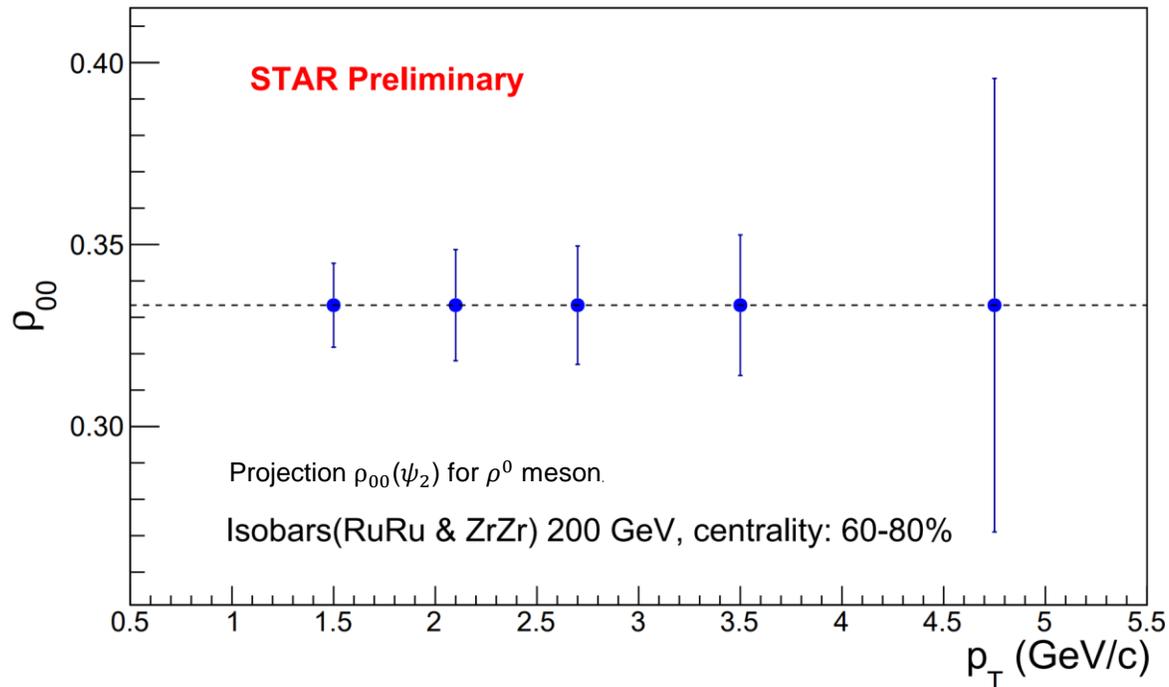
$$\left[ \frac{dN}{d\cos\theta^*} \right]_{|\eta|} \propto \left( 1 + \frac{B'F}{2} \right) + (A' + F - B'F) \cos^2\theta^* + \left( A'F + \frac{B'F}{2} \right) \cos^4\theta^*$$

AuAu for run 2011 at 200 GeV, Centrality: 60-80%,  $p_T$ : 1.8-2.4 GeV/c.

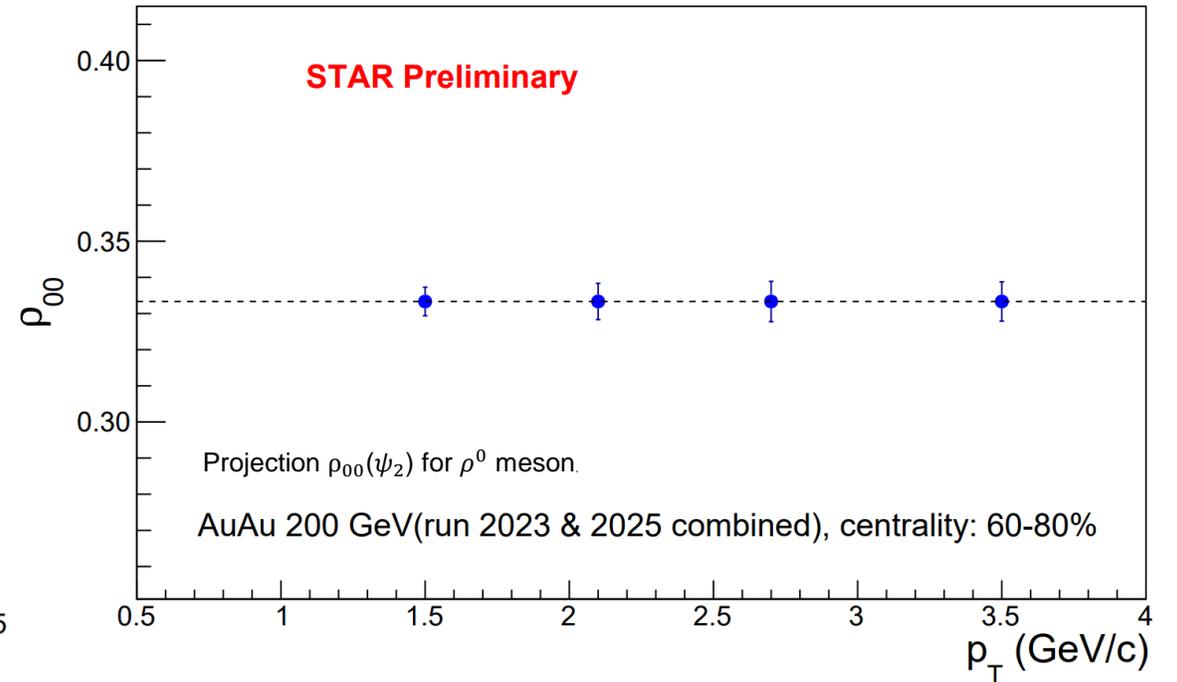
B. Xi, EPJ Web of Conferences 296, 04004 (2024).

# Error projection for $\rho^0$ meson $\rho_{00}$ for isobar and AuAu 200 GeV

Statistical error projection  
for isobar at 200 GeV.



Statistical error projection for AuAu 200 GeV,  
run 2023 & 2025 combined.



B. Xi for STAR, Quark Matter 2023.

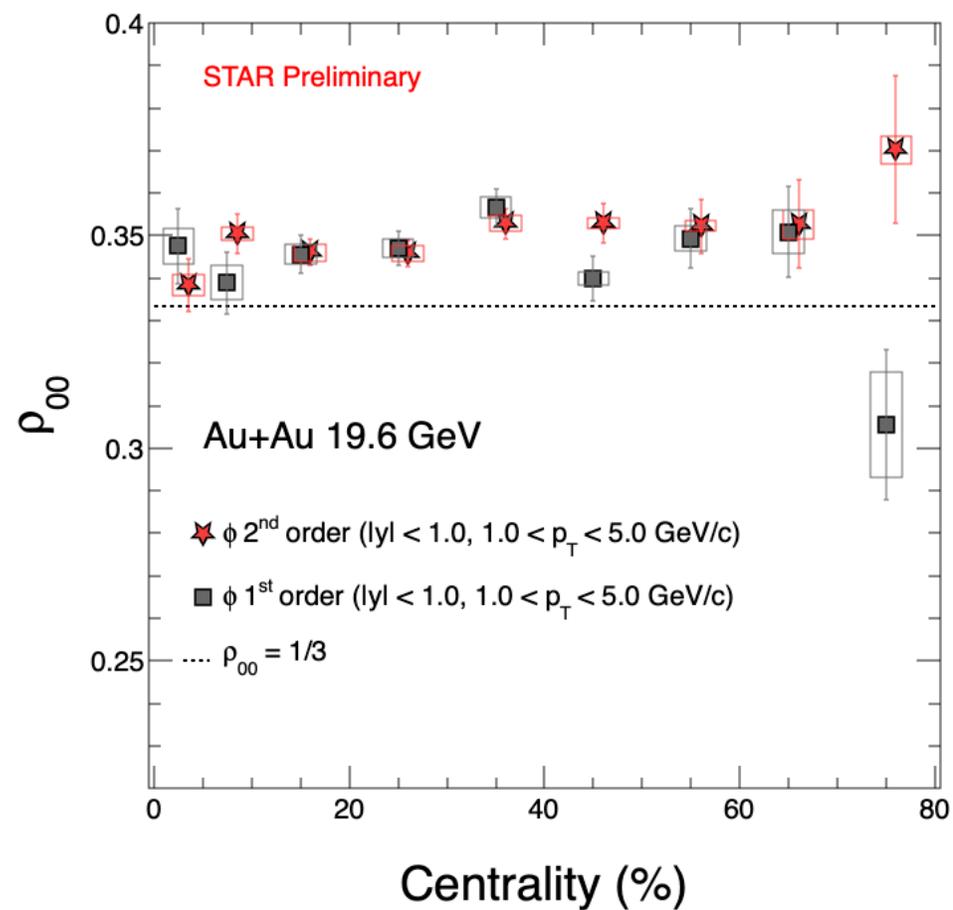
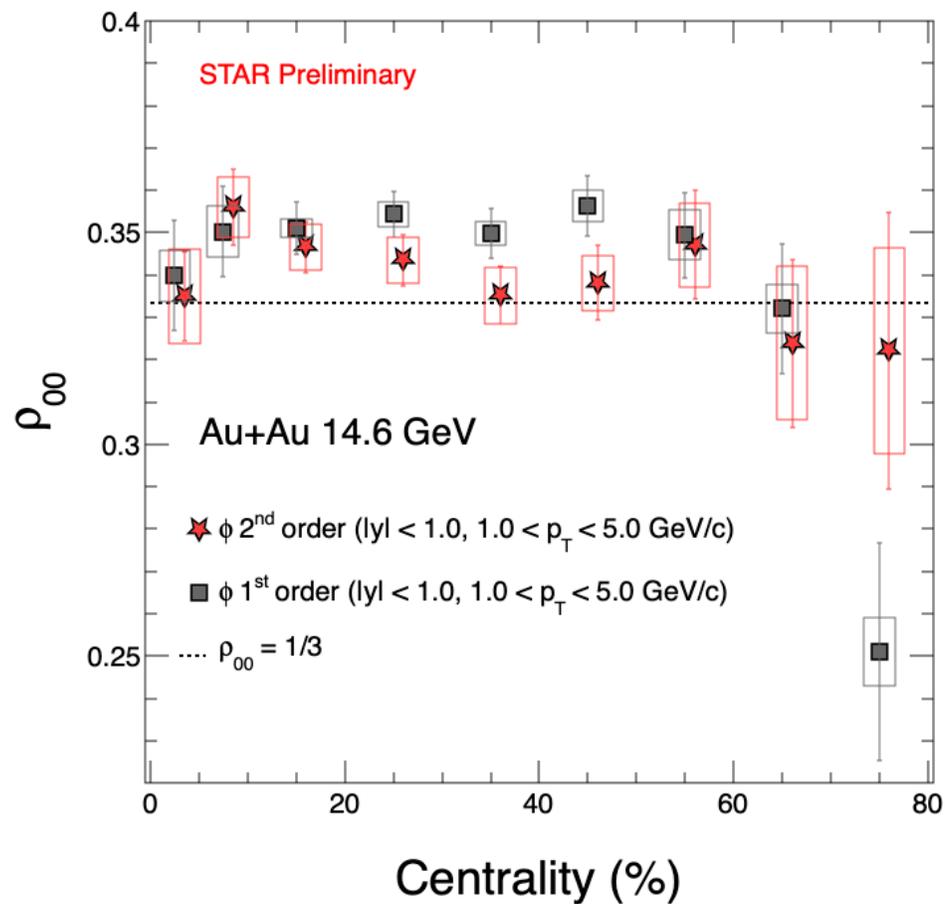
# Summary

- Summarized the experimental and theoretical progress of global spin alignment. Discussed latest experimental progress of  $J/\Psi$  and  $\phi$ .
- STAR measured global spin alignment of  $\phi$  and  $K^{*0}$ . Currently, the measurements for  $J/\Psi$ ,  $\rho^0$  and rapidity dependent of  $\phi$  are in progress.
- Discussed the measurement method and correction of global spin alignment.
- Physical significance of the global spin alignment of  $\rho^0$  mesons.

Thank You

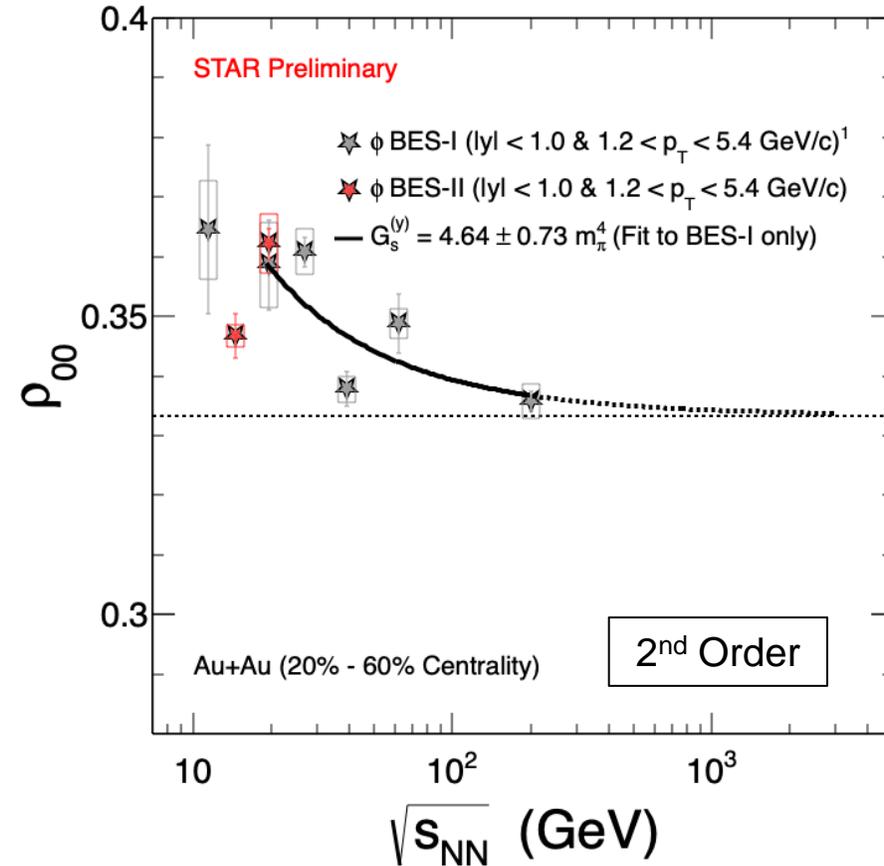
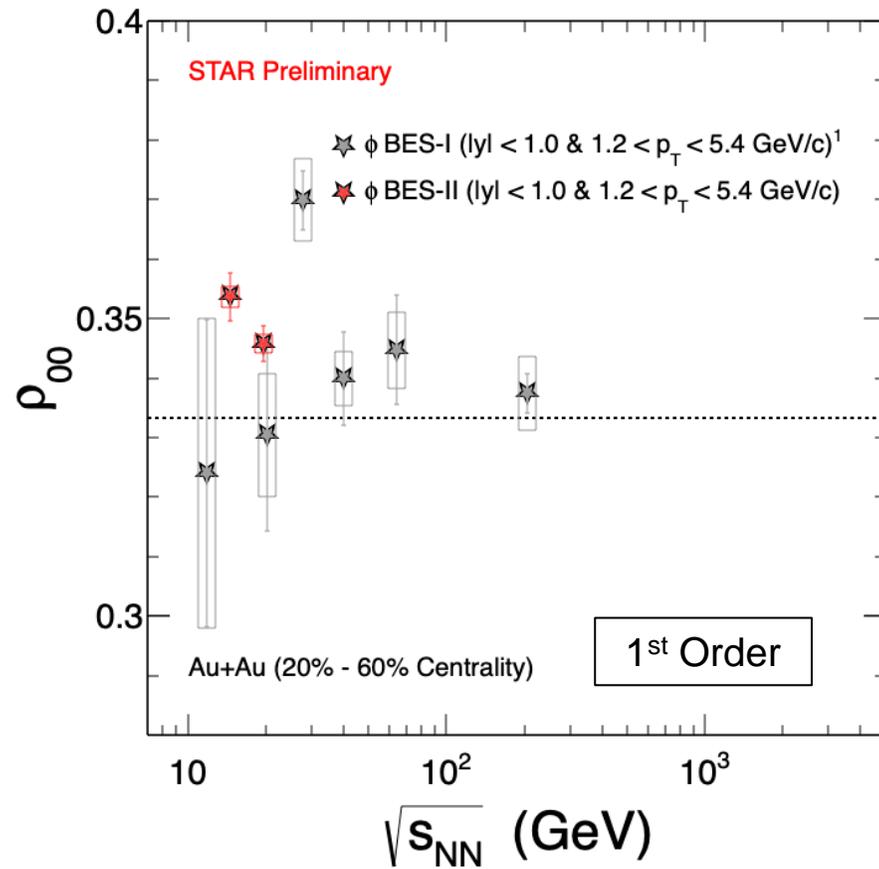
# Backup Slides

# $\rho_{00}$ of $\phi$ as a function of centrality



1st and 2nd order EP results have similar centrality dependence.

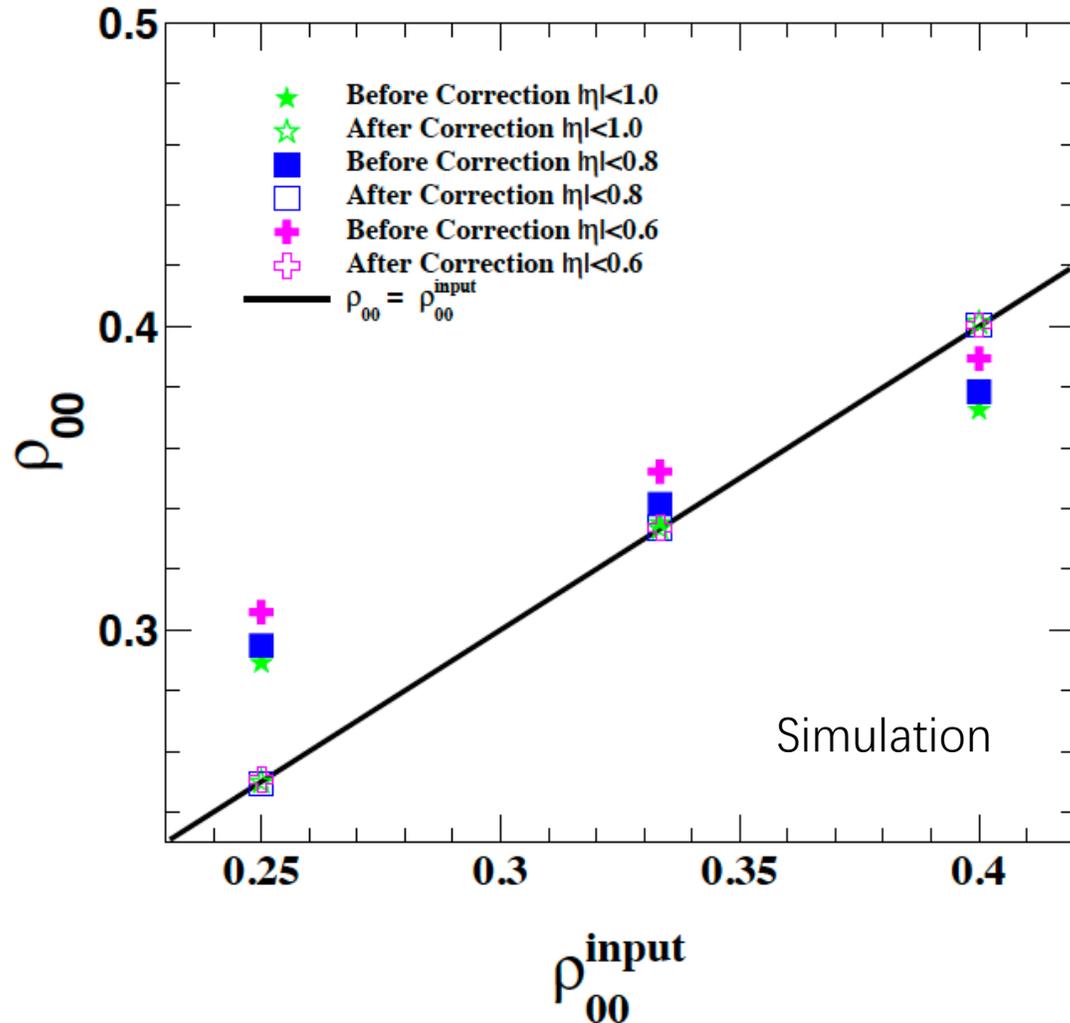
# $\rho_{00}$ of $\phi$ as a function of collision energy ( $\sqrt{s_{NN}}$ )



- [1] STAR Collaboration, Nature 614 (2023) 7947.  
 [2] Sheng et al., Phys. Rev. D 101 (2020) 9, 096005.  
 [3] Sheng et al., Phys. Rev. D 102 (2020) 5, 056013.

**Significant global spin alignment confirmed at 14.6 GeV and 19.6 GeV.**

# Acceptance and Resolution Correction QA



- We simulated  $\phi$ -mesons with the following inputs:
  - 3 input  $\rho_{00}$  values:  $\{0.25, 0.3333, 0.4\}$
  - $p_T = 1.5 \text{ GeV}/c$
  - $v_2 = 0.075$
  - $y = [-1.0, 1.0]$  flat
- We cut on the daughters with three different  $|\eta|$  cuts:  $|\eta| < \{0.6, 0.8, 1.0\}$ .
- What we show is that before corrections the  $\rho_{00}$  we recover is not the same as the input. After we apply the acceptance and resolution corrections, the reconstructed values are equivalent to the input.

# Datasets and cuts of $\phi$ spin alignment in BES-II

Au+Au 14.6 GeV BES-II (2019) (minbias)

Au+Au 19.6 GeV BES-II (2019) (minbias)

## Event cuts:

$|V_z| < 70 \text{ cm}$  ,  $V_r < 2.0 \text{ cm}$  ,  $n\text{BToFMatch} > 2$

Pile-up rejection cuts

## Track cuts:

$n\text{HitsFit} > 15$  ,  $n\text{HitsFit}/n\text{HitsMax} > 0.52$  ,

$|\eta| < 1$  ,  $dca < 2.0 \text{ cm}$  ,

$p_T > 0.1 \text{ GeV}/c$  &&  $p < 10 \text{ GeV}/c$

$0.16 < \text{mass2} < 0.36$  ,  $|\text{nSigmaKion}| < 2.5$

# Systematic of $\phi$ -analysis in BES-II

- $n\sigma_\pi$ : 2.0, 2.5, 3.0
- dca : 2.0, 2.5, 3.0
- Background normalization range: [1.04, 1.05] , [0.99, 1.0] , average of both
- Yield extraction method: bin counting, integration
- Yield extraction range:  $2.0\sigma$ ,  $2.5\sigma$ ,  $3.0\sigma$

For a given source of systematic uncertainties, we obtained  $\rho_{00}$  with the cut for this sources changed, and other cuts are at the central value. Assuming uniform probability distributions between the maximum and minimum values, the value of the systematic uncertainty for a source is:

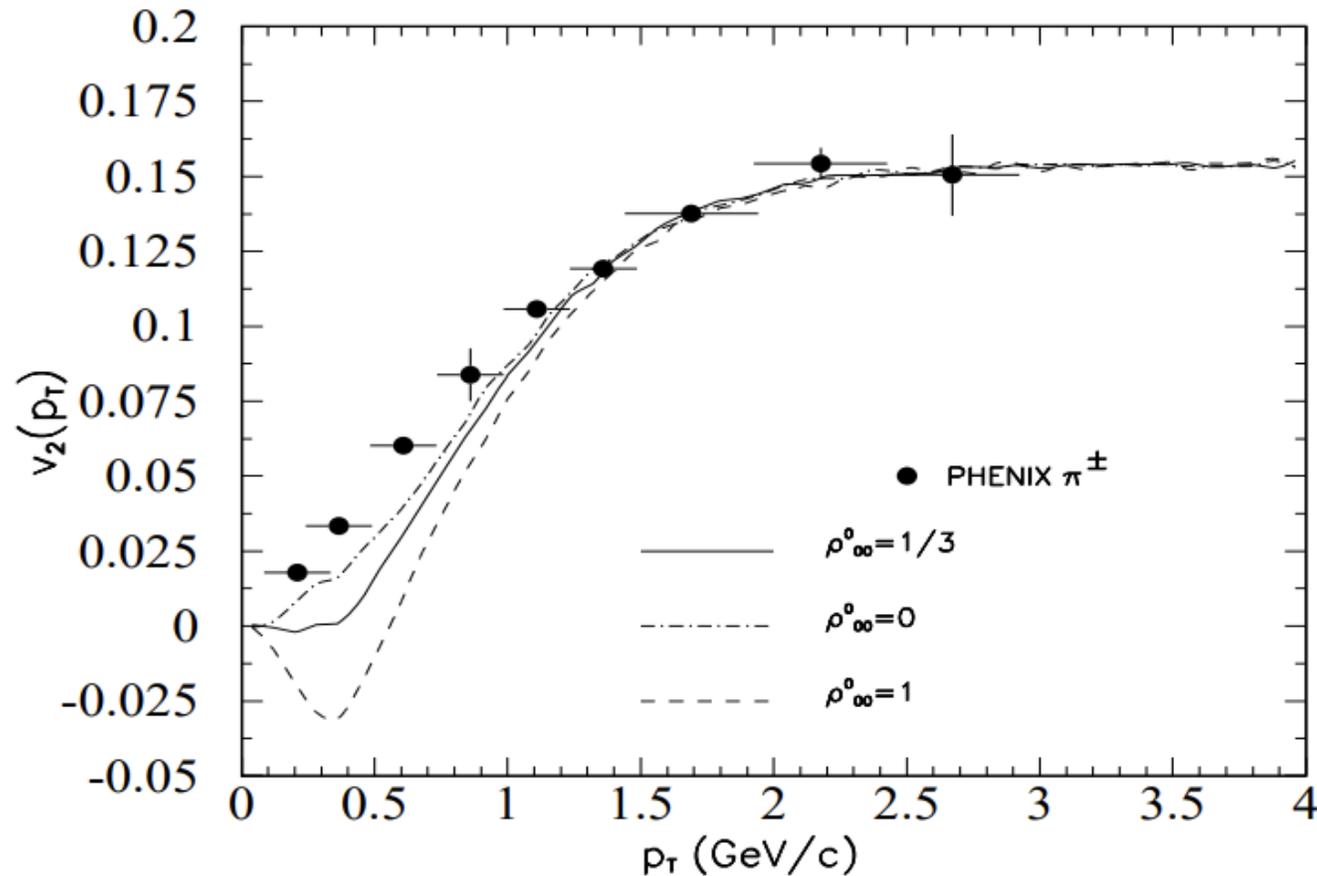
$$\Delta\rho_{00,sys}^i = \frac{\rho_{00,max}^i - \rho_{00,min}^i}{\sqrt{12}} \quad (39)$$

and then combine different sources of uncertainties:

$$\Delta\rho_{00,sys} = \sqrt{\sum_i (\Delta\rho_{00,sys}^i)^2} \quad (40)$$

\* For rapidity dependence, we took the statistical weighted average of the symmetric negative and positive bins as the central value. The difference between points was added as a source of systematic error.

# The $v_2$ of $\pi$ considering the decay of $\rho^0$ mesons with spin alignment effect



Z. T. Liang and X. N. Wang, Spin alignment of vector mesons in non-central A+A collisions, Phys. Lett. B 629, (2005) 20.

# For $\rho^0$ : datasets and cuts of Run 2011 and isobar

## Minimum Bias Event of AuAu 200GeV from 2011

(~500 M before event cuts. ~300M after event cuts)

### Event cuts:

$|V_z| < 30.0$  cm,  $V_r < 2.0$  cm,  $|V_z - V_zVPD| < 3.0$  cm

Number ToF matched point  $> 3$

### Track cuts: (Same for different datasets)

$n_{\text{HitsFit}} > 15$ ,  $n_{\text{HitsFit}}/n_{\text{HitsMax}} > 0.55$

$|\eta| < 0.8$ ,  $dca < 2.0$  cm,  $p_T > 0.2$  GeV/c &&  $p < 10$  GeV/c

$|n_{\text{SigmaPion}}| < 1.5$ ,  $-0.1 < \text{mass2} < 0.2$

## Minimum Bias Event of isobar

( For RuRu: ~2000 M before event cuts. ~1600M after event cuts.

For ZrZr: ~2200 M before event cuts. ~1800M after event cuts.)

### Event cuts:

$-35$  cm  $< V_z < 25$  cm,  $V_r < 2.0$  cm,  $|V_z - V_zVPD| < 3.0$  cm

Number ToF matched point  $> 3$

# Systematics of $\rho^0$ -analysis

- $n\sigma_\pi$ : 1.0, 1.5, 2.0
- dca : 1.5, 2.0, 2.5
- Normalization factor (small medium large)
- Fitting procedure
- Count and integration range:  $0.5*\Gamma$ ,  $1.0*\Gamma$ ,  $1.5*\Gamma$
- Yield extraction method: bin counting, integration

Bin counting are used in all parts(as in the analysis of  $\phi$  Nature).